Beyond Standard Model from Flavor Physics.

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OUTLINE

- Basic idea of talk: what is the SM, why BSM, how to look for BSM physics in experiments like Belle 2.
- Start with : What is a fundamental particle?
- How do fundamental particles interact- Standard Model.
- Why Beyond Standard Model (BSM) is needed.
- Some BSM models and their signatures.

Particle Types

- Particles come in two types: Bosons and Fermions
- A collection of bosons and fermions will fill up a given set of energy levels differently
- Bosons satisfy Bose-Einstein Statistics. Fermions satisfy the Fermi-Dirac Statistics.
- Bosons have integral spin (measured in terms of \hbar) and are responsible for forces between particles
- Bosons: Spin 0 (Higgs), Spin 1 (Photon), Spin 2 (Graviton)
- \circ Fermions have half integral spin and are the matter particles: E.g.: Spin 1/2 (electron), Spin 1/2 (neutrino- $\nu)$

Fundamental matter particle

- Generally, a fundamental particle is characterized by its spin and charge.
- Spin: Intrinsic angular momentum of a particle- intrinsic property just like mass and charge.



A purely left-handed(LH) or a right- handed (RH) particle is massless.

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STANDARD MODEL OF ELEMENTARY PARTICLES



Forces

Standard Model

• SM has a gauge sector: Coupling of the matter field with the gauge bosons: 3 parameters, g_W , g_S and g_{EM} .



EM force

Language of PP- Feynman Diagrams

• Particles and their interactions are described by Feynman Diagrams

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Strong Force

QCD

QCD binds quarks in mesons, baryons, tetraquarks, pentaquarks



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Weak Force - Charged



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Weak force - Neutral

Neutral Current Interactions

Neutral Current Interaction are mediated through the exchange of Z^0 . e - v and p - v scattering е ν Ζ е IJ d р р Ζ ν ν (3) э. Alakabha Datta (UMiss) Flavor Sector of the SM and New Physics Nov 3, 2016

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Weak Interactions are Different

• Any quark (u and d) and lepton (e and ν_e) can split up into a left handed (LH) and a right handed (RH) piece.

 $u \equiv u_L + u_R$



• W^{\pm} couple only to LH particles and RH antiparticles.

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Basic Forces and the SM

- Electromagnetic force couple equally to the LH and RH particles if they are electrically charged.
- Strong Force couple equally to the LH and RH particles if they are color charged.
- Weak Interactions only couple to LH particles.
- Neutrino only has weak interactions. In the SM there is only a LH neutrino, ν_L , and no RH neutrino, ν_R , and neutrinos are massless.

Higgs Force- mass generation

- A purely left-handed(LH) or a right- handed (RH) particle is massless.
- However certain particles can change from LH to RH via interacting with a Higgs boson. In that case the particle is sum of the LH and RH particles: $e = e_L + e_R$.
- Such particles can become massive.



The Higgs Mechanism- Flavor sector

We are aware of phase change- for example steam to water to ice.

When the universe was very hot at a temperature equivalent to 100 GeV (300 K $\equiv 10^{-2}$ eV) there was a phase change when the Higgs, *W*, *Z*, the quarks and leptons went from being massless to having masses.



Evidence of BSM physics

Clearly Established

- Neutrino Masses.
- Dark matter.

Several Hints

- Anomalies in semileptonic *B* decays.
- Neutrino anomalies.
- Hints of new scalar bosons.

What Is a Neutrino?

A neutrino is a subatomic particle and also a fundamental particle.

- Symbol is lowercase Greek letter nu: $\boldsymbol{\nu}$
- Electrically neutral
- Nearly massless
- Travels close to the speed of light
- Reacts to gravity and the weak nuclear force
- Mostly passes through matter
- Oscillates between flavors (electron, muon, tau)
- Has antimatter equivalents
- Nuclear processes produce neutrinos
- Billions of neutrinos pass through your body every second
- Neutrinos account for around 2-3% of the Sun's energy



sciencenotes.org



The image has been obtained with a 503 days exposure, by registering neutrinos emitted from the solar core and detected in a 50 000-ton water pool located 1 km underground. At night, neutrinos were transparently traversing the whole earth before being registered in this image

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Neutrinos Oscillate



A particle has wave like behavior with $E = \sqrt{p^2 c^2 + m^2 c^4} = \hbar \omega$ and $p = \frac{h}{\lambda}$.

$$\begin{array}{ll} |\nu_e\rangle & = & \cos\theta \, |\nu_1\rangle + \sin\theta \, |\nu_2\rangle \\ |\nu_\mu\rangle & = & -\sin\theta \, |\nu_1\rangle + \cos\theta \, |\nu_2\rangle \end{array}$$

 $|
u_1
angle$ and $|
u_2
angle$ have different masses (or frequencies)

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Neutrino Oscillation Model





$$\begin{aligned} |\nu_e\rangle &= \cos\theta \,|\nu_1\rangle \, e^{-i\omega_1 t} + \sin\theta \,|\nu_2\rangle \, e^{-i\omega_2 t} \\ |\nu_\mu\rangle &= -\sin\theta \,|\nu_1\rangle \, e^{-i\omega_1 t} + \cos\theta \,|\nu_2\rangle \, e^{-i\omega_2 t} \end{aligned}$$



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Neutrino Oscillations





$$\begin{aligned} |\nu_e\rangle &= \cos\theta |\nu_1\rangle \, e^{-i\omega_1 t} + \sin\theta |\nu_2\rangle \, e^{-i\omega_2 t} \\ |\nu_\mu\rangle &= -\sin\theta |\nu_1\rangle \, e^{-i\omega_1 t} + \cos\theta |\nu_2\rangle \, e^{-i\omega_2 t} \end{aligned}$$

Sterile Neutrino

- Neutrino Oscillation indicate that neutrinos have mass.
- $\circ\,$ In general this means that there must be something more than just a LH, ν_L , which is massless.

A possibility is there are new RH ν_R states. This RH neutrino, ν_R , is called the sterile neutrino.

- Weak Interactions only couple to LH neutrinos. Hence the sterile neutrino has no known interactions except for gravitational interaction.
- How can we detect the sterile neutrino. Can it be dark matter?

Energy Distribution of the Universe

ENERGY DISTRIBUTION OF THE UNIVERSE



- Normal Matter: Radiation + Matter.
- Dark Energy is the vacuum energy- Cosmological constant.

• DM is like very weakly interacting non- relativistic (moving slowly) Alakabha Datta (UMiss) Beyond Standard Model from Flavor J June 18, 2024 22/33

Evidence for Dark Matter



$$v^2 = \frac{GM}{r}$$

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Evidence for Dark Matter





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Evidence for Dark Matter- Bullet Cluster



DM is very weakly interaction and can go through each other. https://www.youtube.com/watch?v=eC5Lwjsgl4l

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SM problems - Many

- To many particles: For first generation: $(u_L, d_L) \times 3 + (u_R, d_R) \times 3 + (\nu_L, e_L) + e_R = 15$ states. For 3 generations 45 particles. Cannot be a fundamental theory!
- There is no theory to explain the masses of the fermions and the mixing.
- There is no dark matter candidate.
- Mechanism for neutrino mass.
- Does not include gravity.

Possible BSM: Grand Unification



If forces are unified then new particles should exist such as Leptoquarks and diquarks

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LQ signatures

LQ's can change the rates and distributions of many processes. eg: ${\sf B}$ anomalies

In charged current processes such as $B o D^* au
u_ au$



In neutral current processes such as $B \to K \ell^+ \ell^-$



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Sterile Neutrino Detection

- In general sterile neutrino is a neutrino that does not have any interactions of the the standard model.
- The sterile neutrino can mix with the ordinary neutrino and so $\nu\leftrightarrow$ $\nu \varsigma.$
- Can be detected in beta decay type reactions: $n \rightarrow p + e^- + \bar{\nu}_s$ where ν_s has a mass.
- Sterile neutrino can have interact through new forces with the usual SM particles.

Sterile Neutrino Signatures

For semileptonic $\bar{B} \to D^{*+} \ell^- \bar{\nu}_\ell$, N can be produced.

$$B \to D^{(*)} \ell \nu_{\ell} \to D^{(*)} \ell (\nu_1 \cos \theta + N \sin \theta)$$

Missing mass : $p_M^2 = (p_B - p_{D^{(*)}} - p_{\ell})^2$



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Dark Matter Signature







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Dark Matter Signature

Decays to invisible states. Penguin Flavor Changing Neutral Current (FCNC)



Look for rare decays $B \to K + \text{inv}$, $D \to \pi + \text{inv}$, $K \to \pi + \text{inv}$

Eg: $b \rightarrow sZ_D \rightarrow s + \chi \bar{\chi}$

Evidence for enhancement in $B \rightarrow K + inv$ by Belle 2.

Summary

- SM is a very successful theory.
- Clear evidence and various hints of BSM physics.
- Many ways to probe various BSM models in Belle 2 and Flavor experiments.
- Exciting results may be revealed in the near future.