

Theory introduction to the dark sector

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KEK

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Overview

Chapter 1: Introduction
Why dark sectors. Theory motivations.
Experimental targets

Chapter 2: Minimal dark sector models
dark photon,

dark scalar, sterile neutrinos

Chapter 3: Axions and axion-like particles
The effective field theory

Chapter 4: Non-minimal theories for dark sectors
Inelastic Dark matter, $L_\mu - L_\tau$

YESTERDAY

TODAY

The Higgs portal

SM + singlet scalar:

$$\mathcal{L} \supset -\underbrace{\frac{\xi}{2}|H|^2 s^2}_{\text{Higgs portal operator}} + \frac{\mu_s^2}{2}s^2 - \frac{\lambda_s}{4!}s^4 + \mu^2|H|^2 - \lambda|H|^4 + \text{couplings within the dark sector}$$

In addition to dark matter models, new (light) scalars appear in e.g., theories that address the hierarchy problem (relaxion), or that explain the baryon-antibaryon asymmetry

The Higgs portal

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Electroweak symmetry breaking:

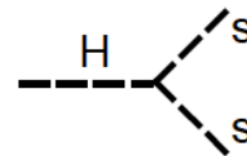
If the scalar, s , gets a VEV, then it will mix with the SM Higgs:

$$\tan \theta_s \sim \frac{\xi v_h v_s}{m_h^2 - m_s^2}$$

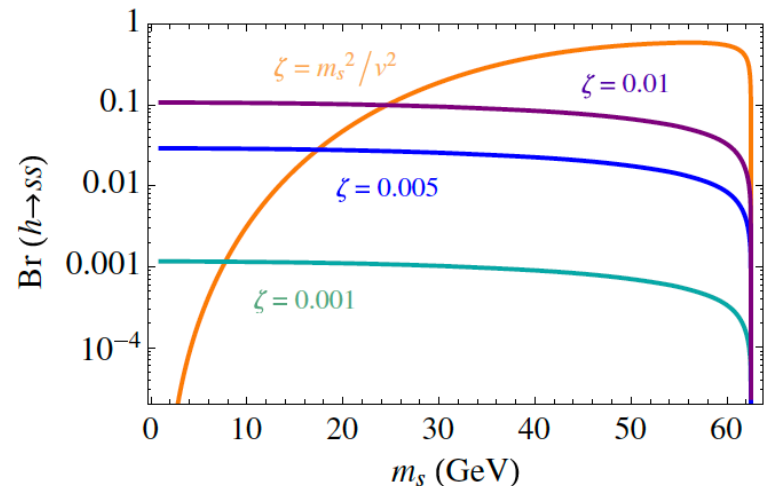


Constraints from Higgs coupling measurements since $\kappa_i \sim \cos \theta_s$

➔ The mixing cannot be too large (model independent bound)



Higgs exotic decays



Curtin et al., 1312.4992

Visible decays of the dark scalar

If the scalar is heavier than ~10 GeV:

- If $\theta_s = 0$, s is stable
 → Scalar invisible decay
- If $\theta_s \neq 0$, s will decay to SM particles

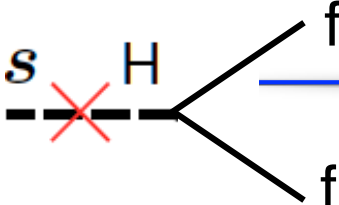
$$\Gamma(s \rightarrow f\bar{f}) = \sin^2\theta_s \frac{N_c m_s m_f^2}{8\pi v^2} \beta_f^3$$

Main BRs: $bb, \pi\pi, cc, \dots$



$\sin(\theta_s) \geq O(10^{-5})$
S decays promptly

$\sin(\theta_s) \leq O(10^{-5})$
S has displaced decays



As we have learnt for the dark photon, also the dark scalar will mainly decay to DM if $2m_{DM} < m_s$

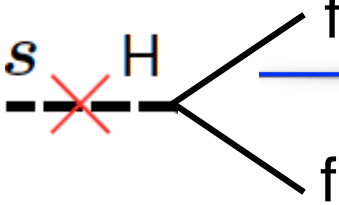
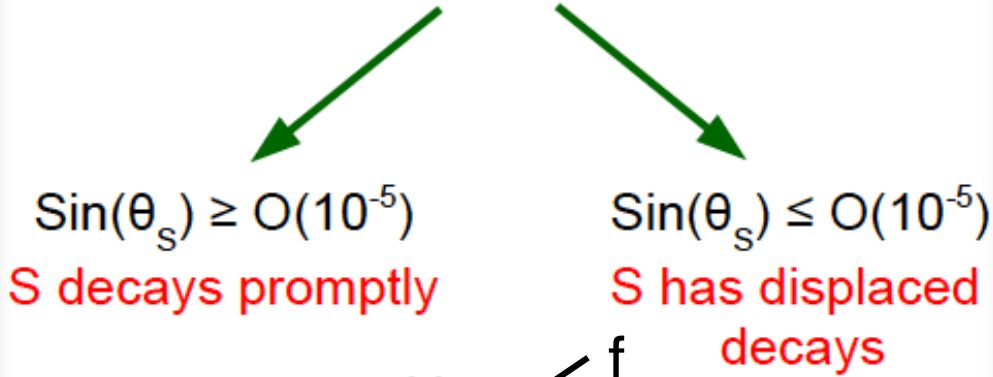
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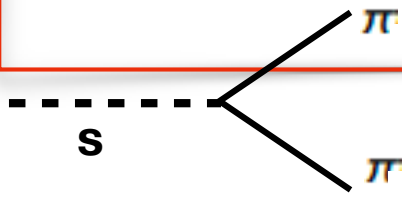
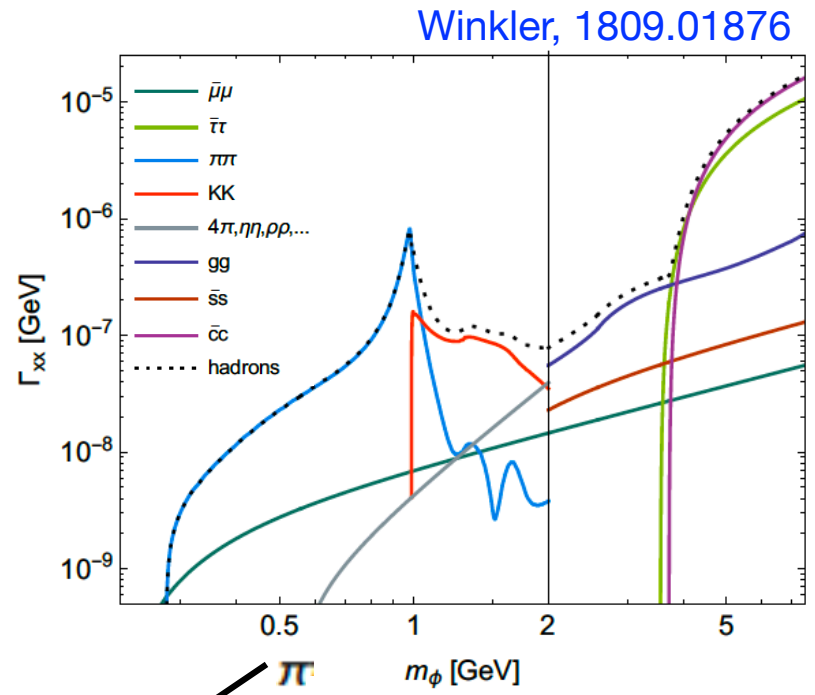
$$\Gamma(s \rightarrow f\bar{f}) = \sin^2\theta_s \frac{N_c m_s m_f^2}{8\pi v^2} \beta_f^3$$

Main BRs: $bb, \pi\pi, cc, \dots$



If the scalar is lighter than ~10 GeV:

The calculation of the widths has large theoretical errors
 Several results in the literature show disagreement



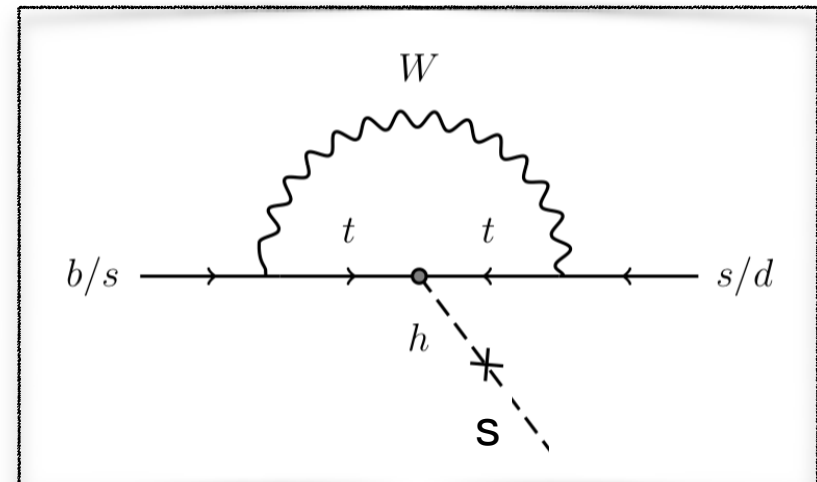
Dark scalars at fixed target experiments

At proton fixed target experiments, large production from B and K meson decays,

$B \rightarrow K s$, $K \rightarrow \pi s$



particularly suited for high energy fixed target (SPS CERN beam)



This will produce two classes of signatures:

$B \rightarrow K + \text{missing energy}$,
 $K \rightarrow \pi + \text{missing energy}$
 in the case of $2m_{DM} < m_s$

1.

$B \rightarrow K + 2 \text{ leptons/tracks}$,
 $K \rightarrow \pi + 2 \text{ leptons/tracks}$
 in the case of $2m_{DM} > m_s$

2.

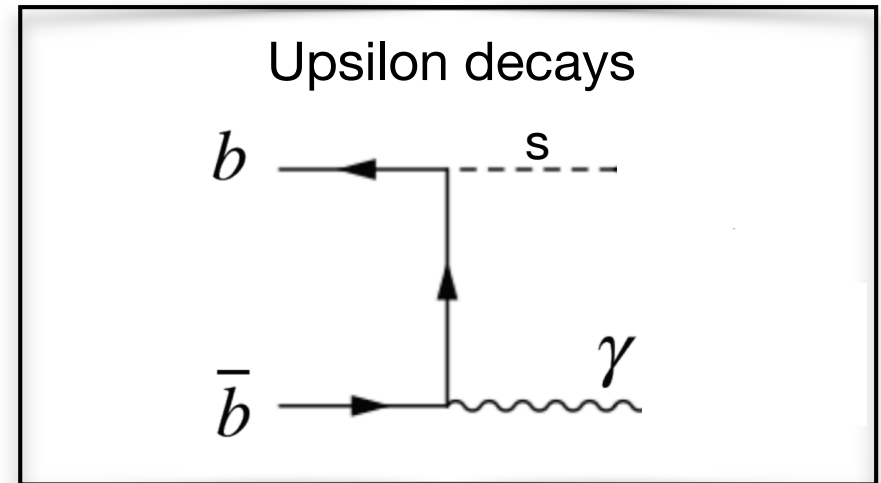
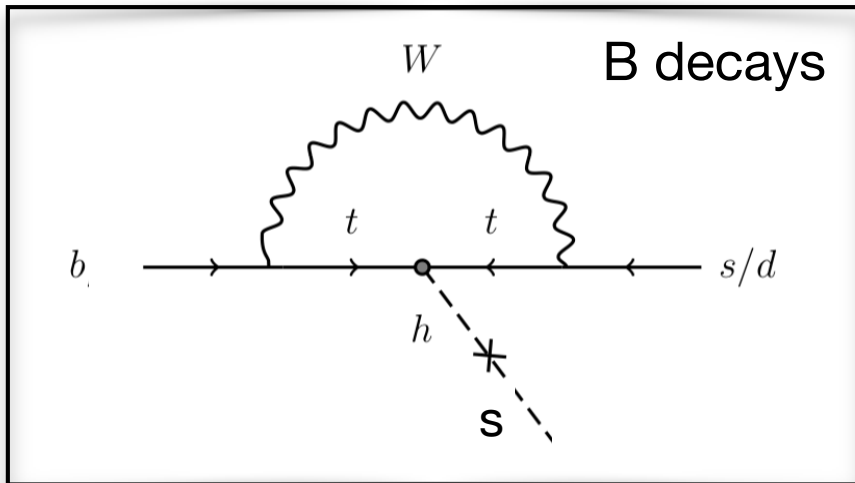
Huge meson statistics

For example, at SHiP:

	SHiP (15 year run)	Meson factory
pions	$\sim 10^{20}$	$\sim 10^{12}$
Kaons	$\sim 10^{20}$	$\sim 10^{14}$
B-mesons	$\sim 10^{13}$	$\sim 10^{10}$

Dark scalars at B factories

2 main production mechanisms:



And then again either **visible or invisible decays** of the scalar, s

Several searches have been performed at Babar

B. Echenard

- $\Upsilon(2S,3S) \rightarrow \gamma A^0, A^0 \rightarrow \mu^+\mu^-$
PRL103 (2009) 081803
- $\Upsilon(3S) \rightarrow \gamma A^0, A^0 \rightarrow \tau^+\tau^-$
PRL103 (2009) 181801
- $\Upsilon(2S,3S) \rightarrow \gamma A^0, A^0 \rightarrow \text{hadrons}$
PRL107 (2011) 221803
- $\Upsilon(2S,3S) \rightarrow \gamma A^0, A^0 \rightarrow \text{invisible}$
arXiv: 0808.0017

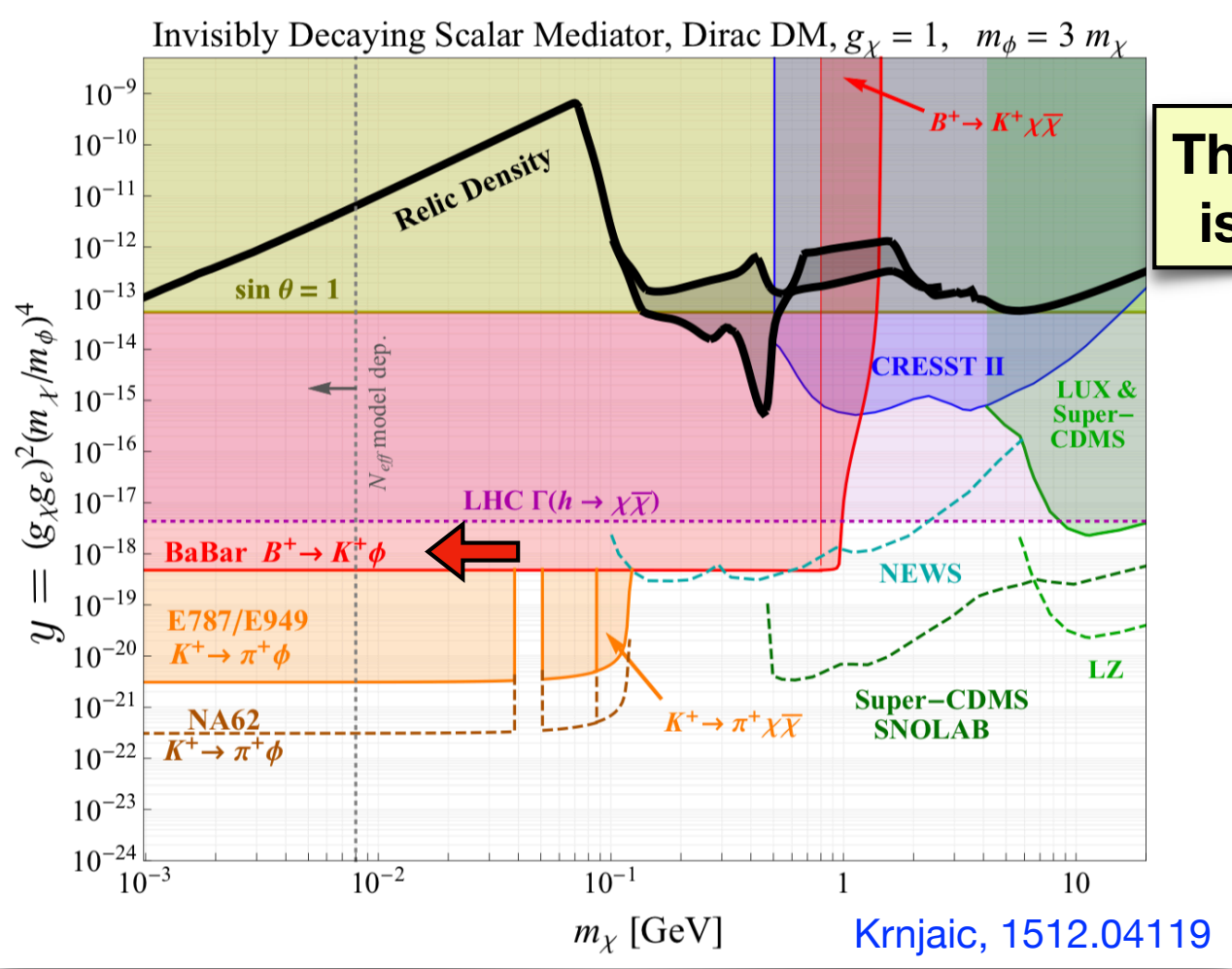
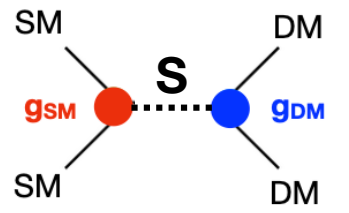
- $\Upsilon(1S) \rightarrow \gamma A^0, A^0 \rightarrow \text{invisible}$
PRL107 (2011) 021804
- $\Upsilon(1S) \rightarrow \gamma A^0, A^0 \rightarrow \mu^+\mu^-$
PRD 87 (2013) 031102
- $\Upsilon(1S) \rightarrow \gamma A^0, A^0 \rightarrow \tau^+\tau^-$
PRD 88 (2013) 071102
- $\Upsilon(1S) \rightarrow A^0, A^0 \rightarrow gg \text{ or } s\bar{s}$
PRD 88 (2013) 031701

1.

Summary plot: the invisible dark scalar

$$\frac{\xi}{2} |H|^2 s^2$$

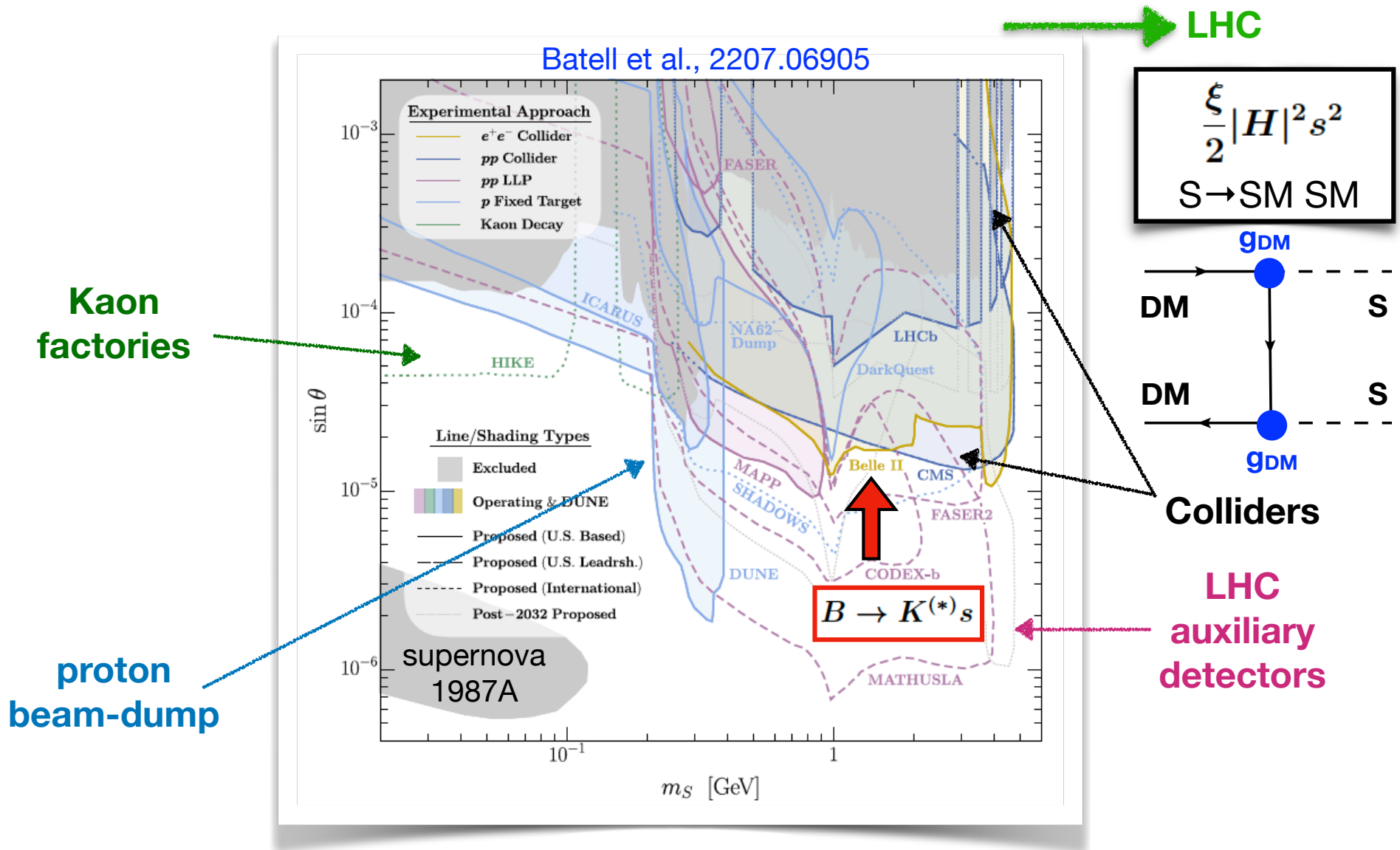
S → DM DM



The thermal line is fully probed!

Note: here we have the assumption, $m_s = 3m_\chi$

Summary plot: the visible dark scalar

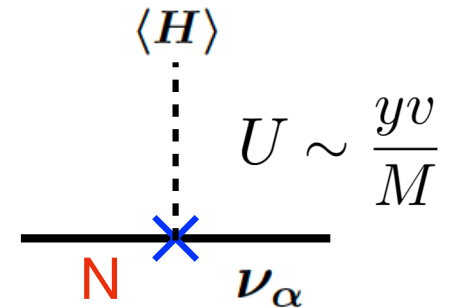


This entire parameter space predicts a **dark sector in thermal equilibrium** with the SM

The neutrino portal

$$\mathcal{L} \supset yHLN + \frac{M}{2}N^2 + \text{interactions in the dark sector}$$

This operator will generically induce some mixing of N with the three active SM neutrinos.



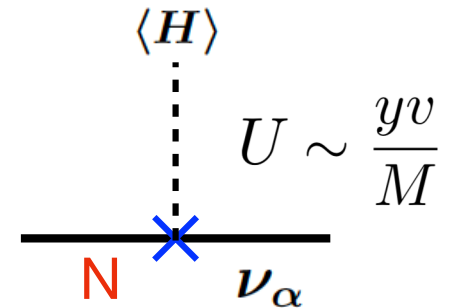
Sterile neutrinos appear in e.g., models that explain the origin of neutrino masses: [seesaw mechanism](#). Sterile neutrinos can have a mass below the few GeV scale in inverse or linear seesaw models.

$$m_\nu^{\text{SM}} \sim \frac{y^2 v^2}{M}$$

The neutrino portal

$$\mathcal{L} \supset yHLN + \frac{M}{2}N^2 + \text{interactions in the dark sector}$$

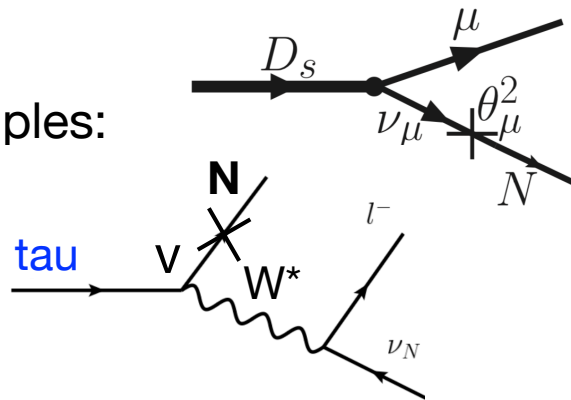
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Production

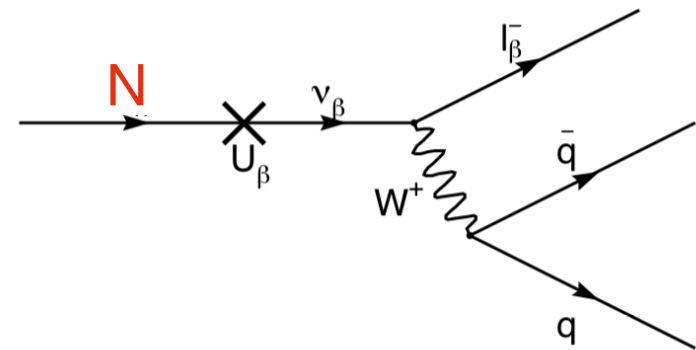
from meson or tau decays:

examples:



at higher energies, decays of W, Z, H
 $W \rightarrow lN, Z/H \rightarrow NN, Z/H \rightarrow N\nu$

Decay



$N \rightarrow Z^{(*)}\nu, N \rightarrow W^{(*)}l, N \rightarrow H^{(*)}\nu$

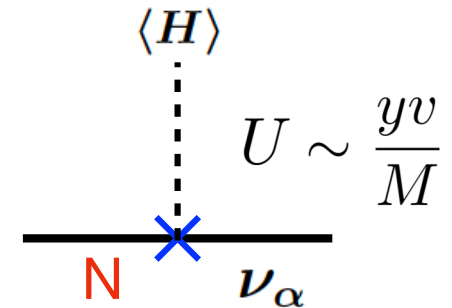
2.

If lighter than other dark sector/dark matter states

The neutrino portal

$$\mathcal{L} \supset yHLN + \frac{M}{2}N^2 + \text{interactions in the dark sector}$$

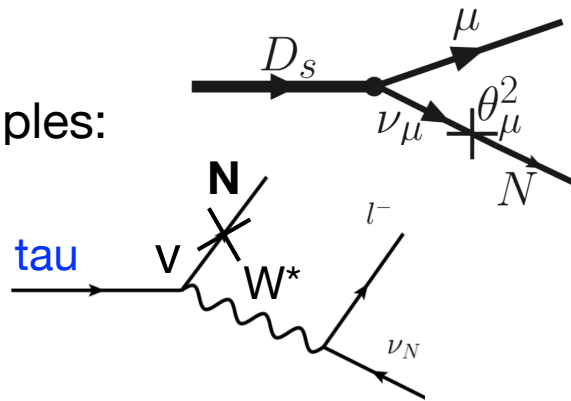
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Production

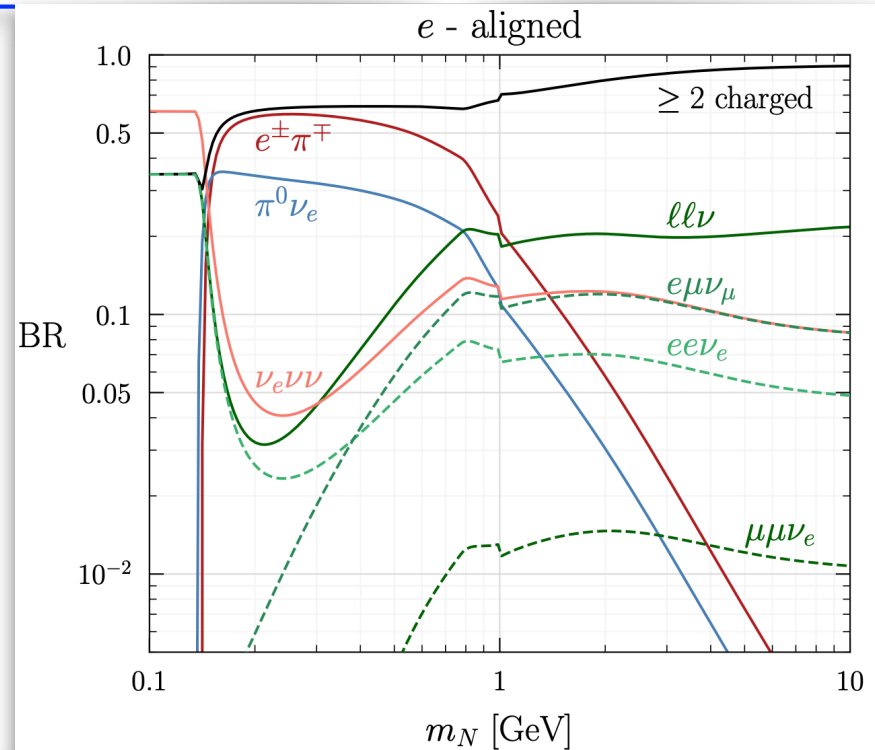
from meson or tau decays:

examples:



at higher energies, decays of W, Z, H

$$W \rightarrow lN, Z/H \rightarrow NN, Z/H \rightarrow N\nu$$

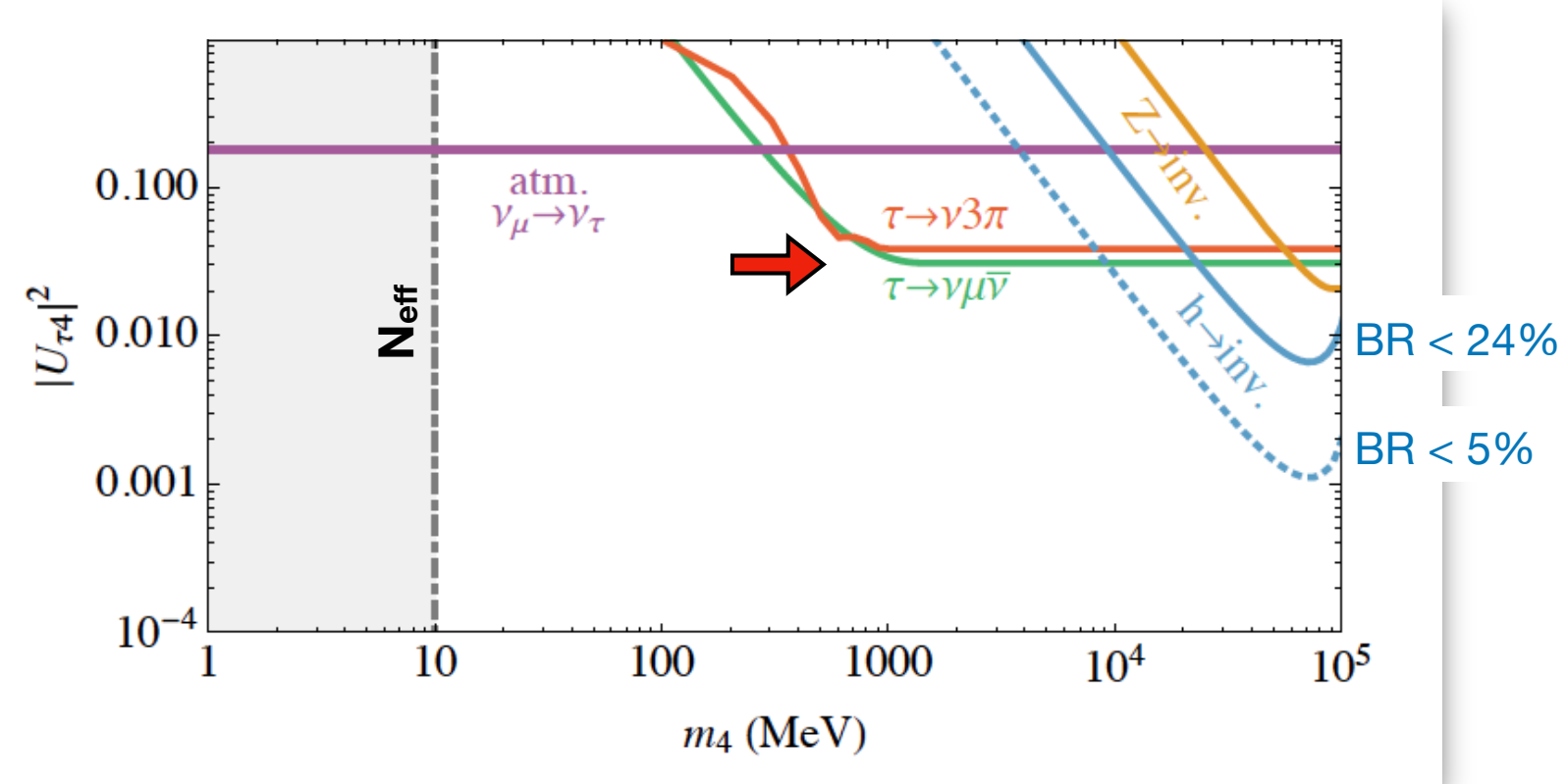


Giffin, SG, Tsai, Tuckler, 2206.13745

1.

Summary plot: the invisible sterile neutrino

Batell et al., 1709.07001

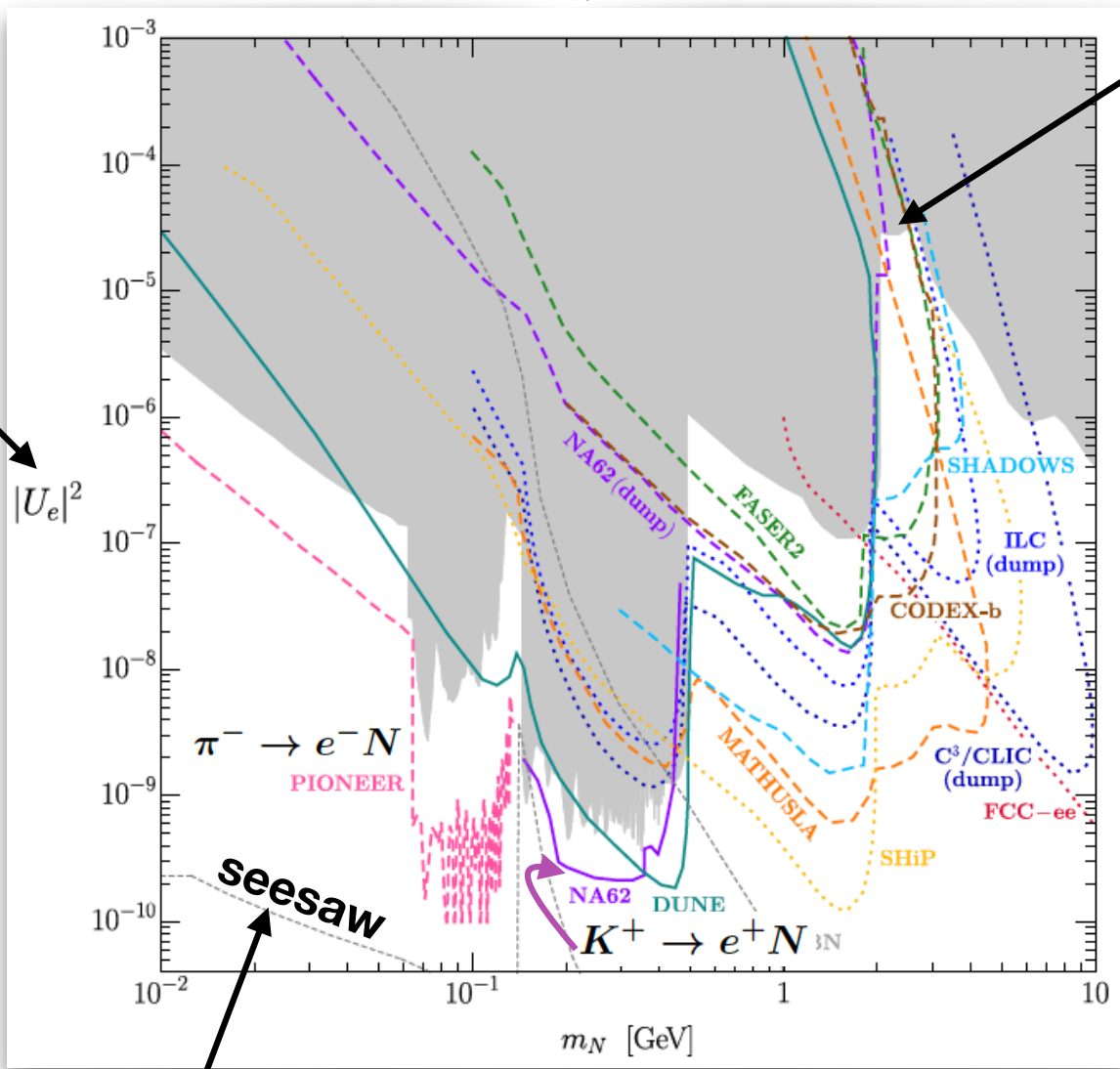


Very interesting opportunity for Belle II
Sterile neutrinos from tau decays.
Need to measure tau decay rates and their kinematics with as much precision as possible.

Summary plot: the visible sterile neutrino

Batell et al., 2207.06905

Example:
mixing with
the electron
neutrino



Belle analysis
with 711 fb⁻¹
1301.1105

$B \rightarrow X l N,$
 $N \rightarrow l \pi (l = e, \mu)$

“vanilla”
model: $|U_e| \simeq \sqrt{\frac{m_{\nu e}}{M_e}}$

Prospects for Belle II?

Chapter 3

Axions & axion-like particles (ALPs)

- * The effective field theory
- * Signatures



How do ALPs couple to the SM?

At dimension 5, the most general Lagrangian for a spin 0, CP-odd particle with an approximate shift symmetry, $a \rightarrow a+c$:

Georgi, Kaplan, Randall 1986

$$\mathcal{L} \supset -\frac{g_{ag}}{4} a G_{\mu\nu}^a \tilde{G}^{a\mu\nu} - \frac{g_{aW}}{4} a W_{\mu\nu}^a \tilde{W}^{a\mu\nu} - \frac{g_{aB}}{4} a B_{\mu\nu} \tilde{B}^{\mu\nu} + ig_{af} (\partial_\mu a) (\bar{f} \gamma^\mu \gamma_5 f)$$

For the complete one-loop analysis, see e.g.
Bonilla et al, 2107.11392
Bauer et al, 2012.12272

$$g_i \propto \frac{1}{f_a}$$

ALP effective field theory (EFT)

At dimension 5, also the operator $ig_{aH} (\partial_\mu a) (H^\dagger D_\mu H + \text{h.c.})$ exists.

However, it can be reabsorbed in the definition of the fermion coupling.
In fact, if we redefine

Redundant operator

$$H \rightarrow e^{ig_{aH} a} H, \quad f \rightarrow e^{-i\beta_f g_{aH} a} f$$

$$\beta_u - \beta_Q = -1, \quad \beta_d - \beta_Q = 1, \quad \beta_e - \beta_L = 1, \quad 3\beta_Q + \beta_L = 0$$

the Higgs operator disappears and the fermion couplings get a shift:

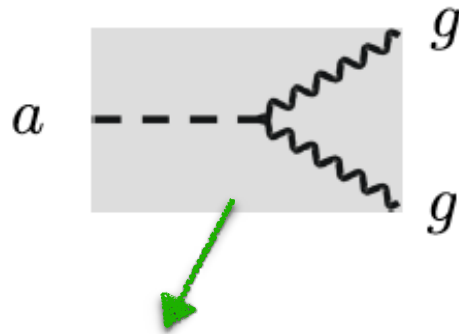
$$g_{af} \rightarrow g_{af} + \beta_f g_{aH} I_{3 \times 3}$$

How do ALPs couple to the SM?

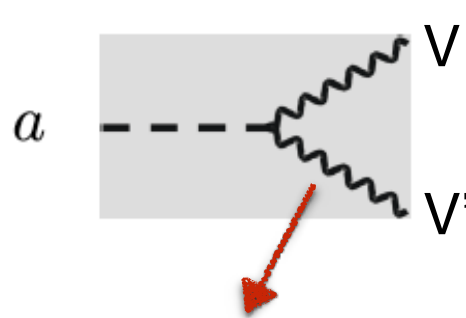
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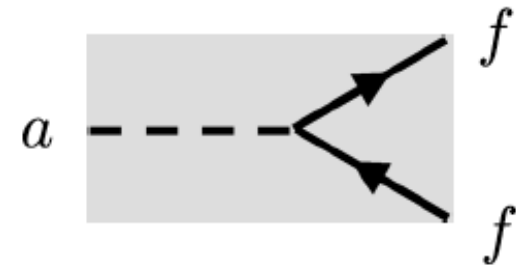
Minimal coupling expected if connection to the strong CP problem.



A ALP-photon coupling is generated in the broken phase

$$g_{aB} \cos^2 \theta + g_{aW} \sin^2 \theta$$

This is the main coupling that has been considered for phenomenological studies of ALPs in the sub-GeV scale.



ALP EFTs at high intensity/energy experiments

see e.g. Brivio et al, 1701.05379
Bauer et al, 1708.00443, 1808.10323, ...

High energy colliders
(LHC, Tevatron, LEP, future colliders)

see e.g. Calibbi et al, 2006.04795
Panci et al, 2209.03371, ...

Low energy flavor experiments
(Mu3e, MEG-II, ...)

$$\mathcal{L} \supset -\frac{g_{ag}}{4} a G_{\mu\nu}^a \tilde{G}^{a\mu\nu} - \frac{g_{aW}}{4} a W_{\mu\nu}^a \tilde{W}^{a\mu\nu} - \frac{g_{aB}}{4} a B_{\mu\nu} \tilde{B}^{\mu\nu} + ig_{af} (\partial_\mu a) (\bar{f} \gamma^\mu \gamma_5 f)$$

Fixed target (beam dump)
experiments
(proton, electron, photons)

see e.g., Dobrich et al, 1512.03069
Harland-Lang et al, 1902.04878, ...

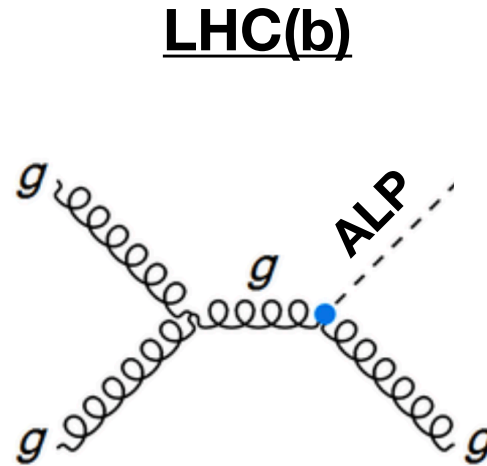
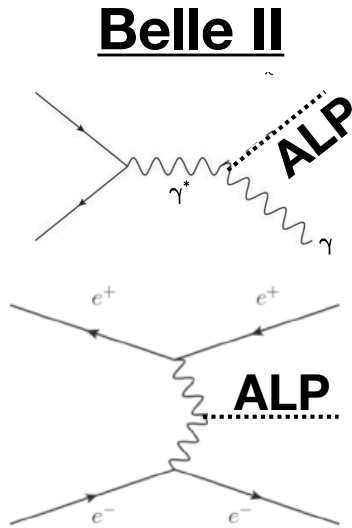
Meson factories
(pion, Kaon, and B-mesons)

see e.g., Bauer et al, 2110.10698
Altmannshofer, Dror, SG, 2209.00665, ...

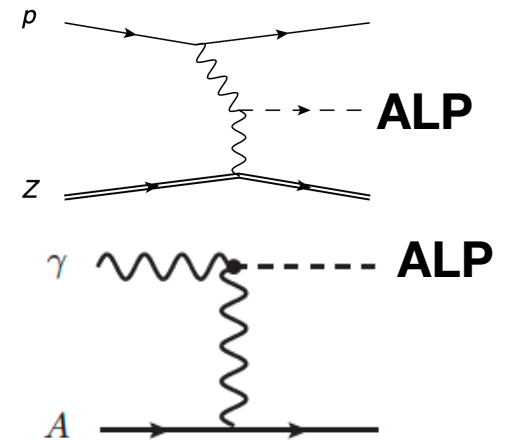
Two classes of ALP production

“Direct production” from beam collision

For example:



proton fixed target



Production from SM particle decays

Neutral current flavor changing transitions

$$K \rightarrow \pi \text{ ALP } (s \rightarrow d \text{ ALP})$$

$$\mu \rightarrow e \text{ ALP}$$

Boson decays

example: $Z \rightarrow \gamma \text{ ALP}$

Charged current transitions

example: $K^+ \rightarrow e^+ \nu \text{ ALP } (u \rightarrow s e^+ \nu \text{ ALP})$

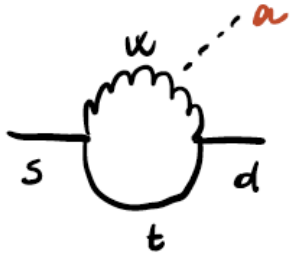
Similar transitions involving different mesons and/or leptons

Neutral & charged current meson decays to ALPs

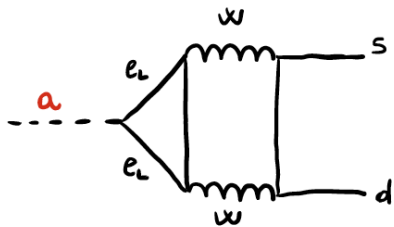
Flavor changing neutral current

They arise in models with

- * ALPs mixed with SM neutral pions (e.g. $K^+ \rightarrow \pi^+ \pi^0 \Rightarrow K^+ \rightarrow \pi^+ a$)
- * ALPs coupling to W or tops



- * ALPs coupling to leptons (higher loop)



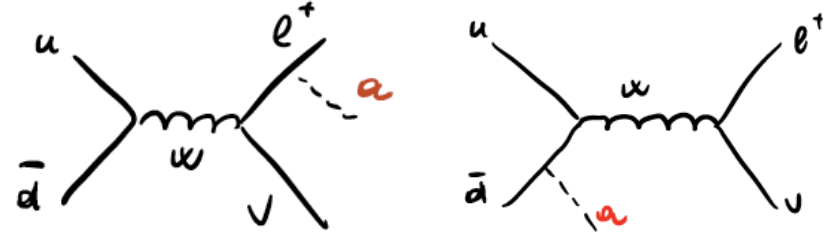
- * Flavor violating ALPs

$$\begin{aligned} K_L &\rightarrow \pi^0 a \\ K^+ &\rightarrow \pi^+ a \\ B &\rightarrow K a \end{aligned}$$

Charged current

They arise in models with

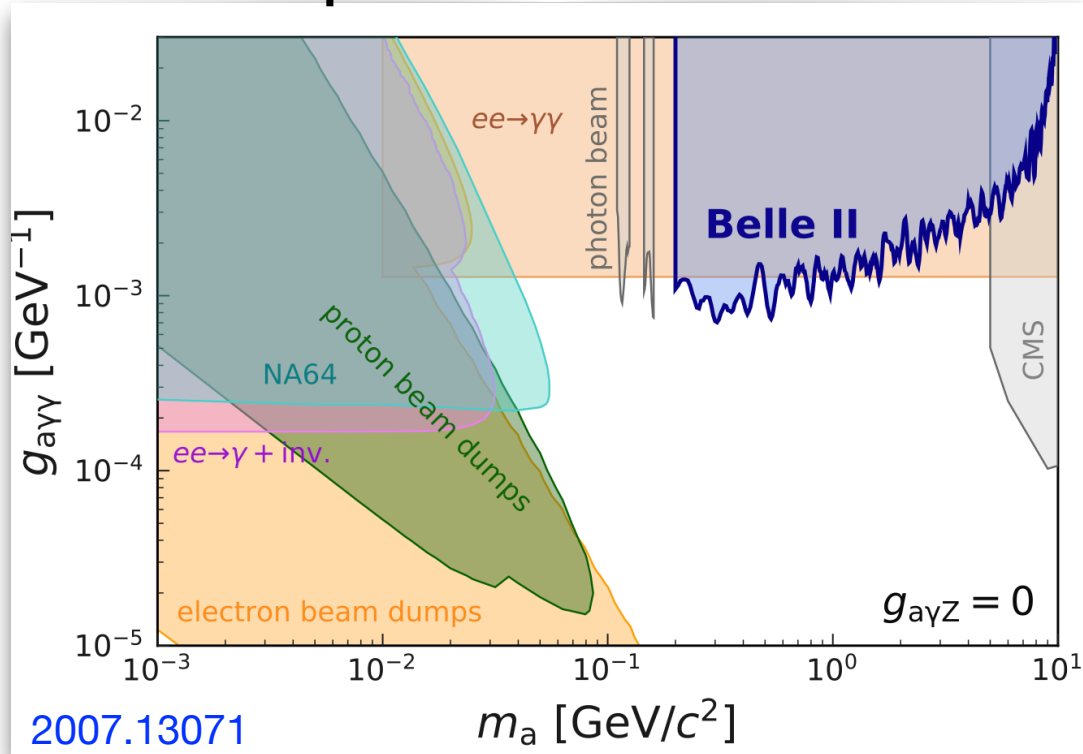
- ALPs mixed with SM neutral pions (e.g. $\pi^+ \rightarrow l^+ \nu \pi^0 \Rightarrow \pi^+ \rightarrow l^+ \nu a$)
- ALP coupling to leptons or quarks



$$\begin{aligned} \pi^+ &\rightarrow a l^+ \nu \\ K^+ &\rightarrow a l^+ \nu \\ B^+ &\rightarrow a l^+ \nu \end{aligned}$$

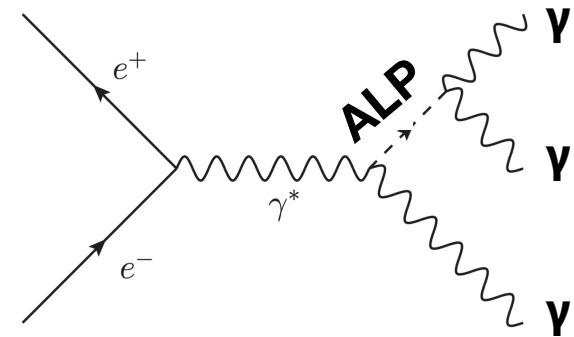
ALPs coupled to photons

496/pb! 10^{-5} of the full data set



Fit of either the recoil mass
or the $\gamma\gamma$ invariant mass

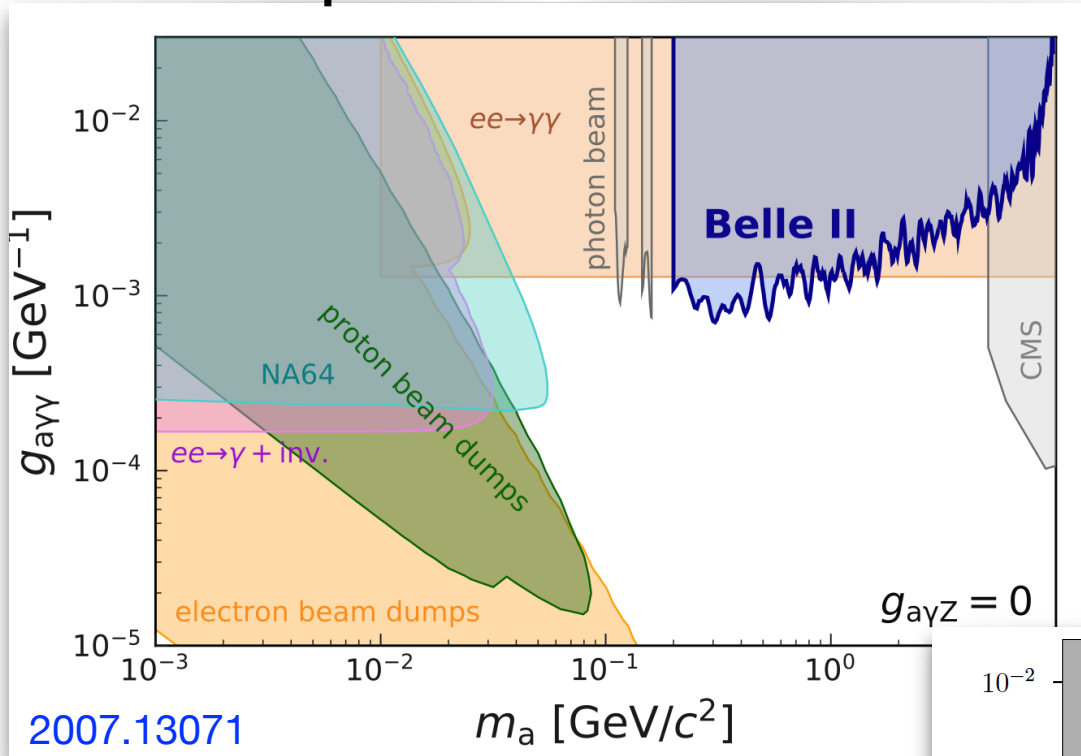
$$M_{\text{recoil}}^2 = s - 2\sqrt{s}E_{\text{recoil}}^{\text{c.m.}}$$



$$\frac{g_{a\gamma\gamma}}{4} a F_{\mu\nu} \tilde{F}^{\mu\nu}$$

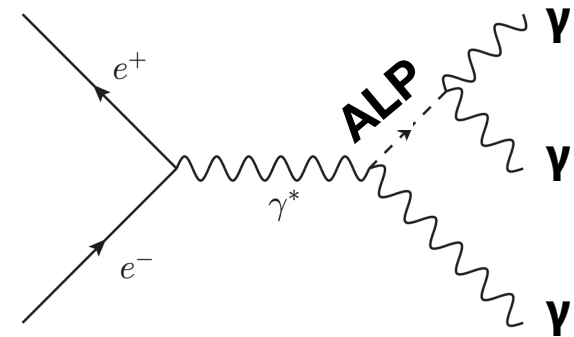
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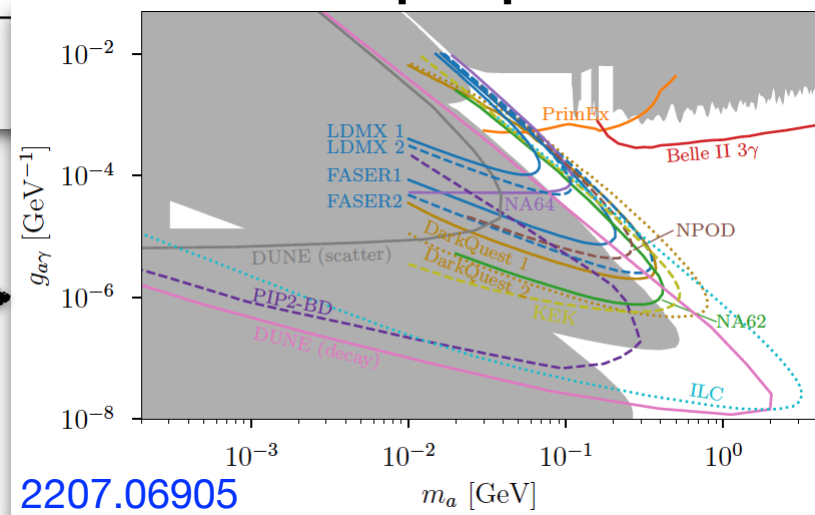


Fit of either the recoil mass
or the $\gamma\gamma$ invariant mass

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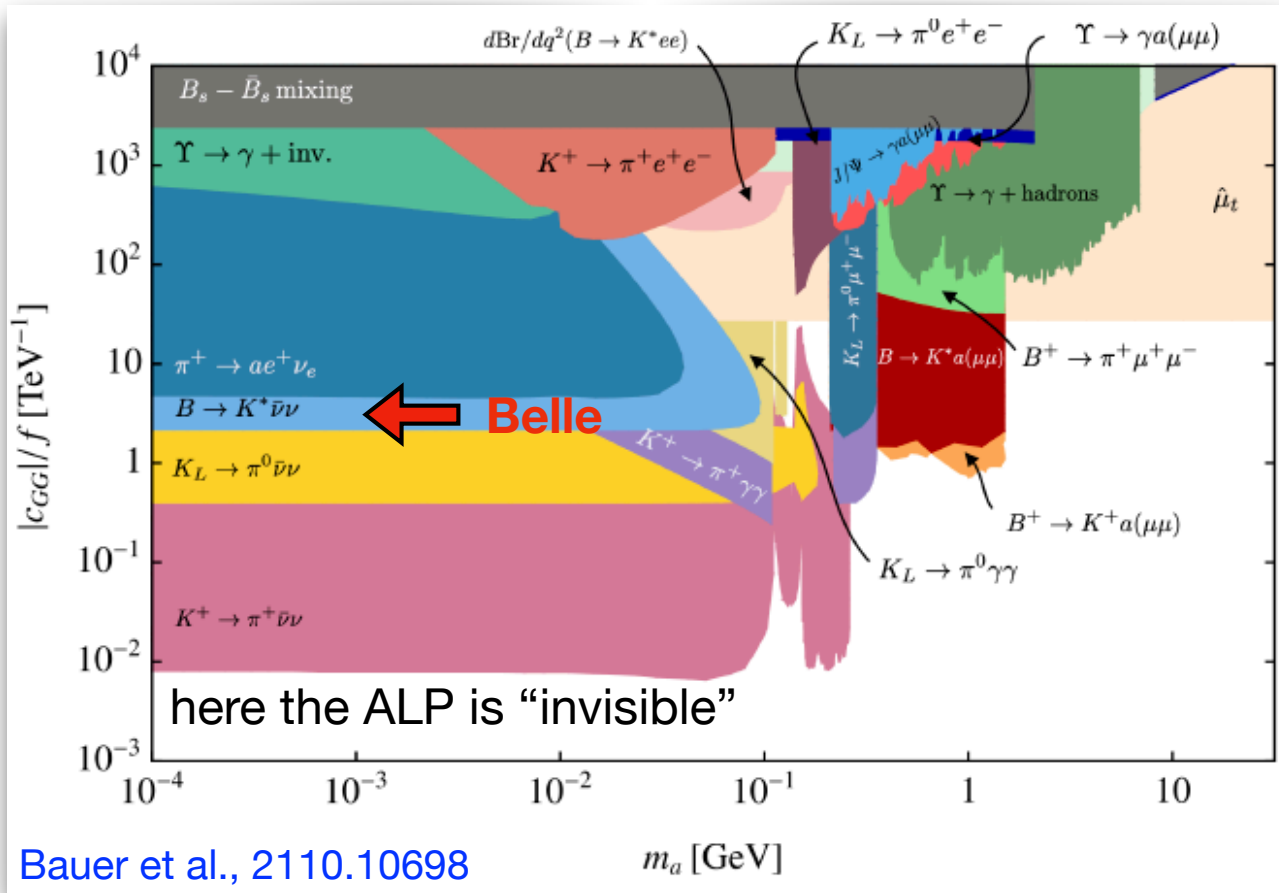
future prospects



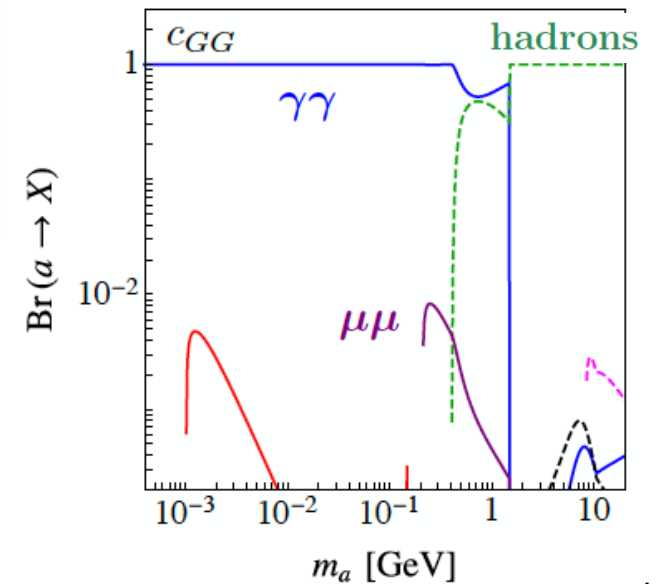
with 20/fb,
Dolan et al.,
1709.00009

$$\frac{g_{a\gamma\gamma}}{4} a F_{\mu\nu} \tilde{F}^{\mu\nu}$$

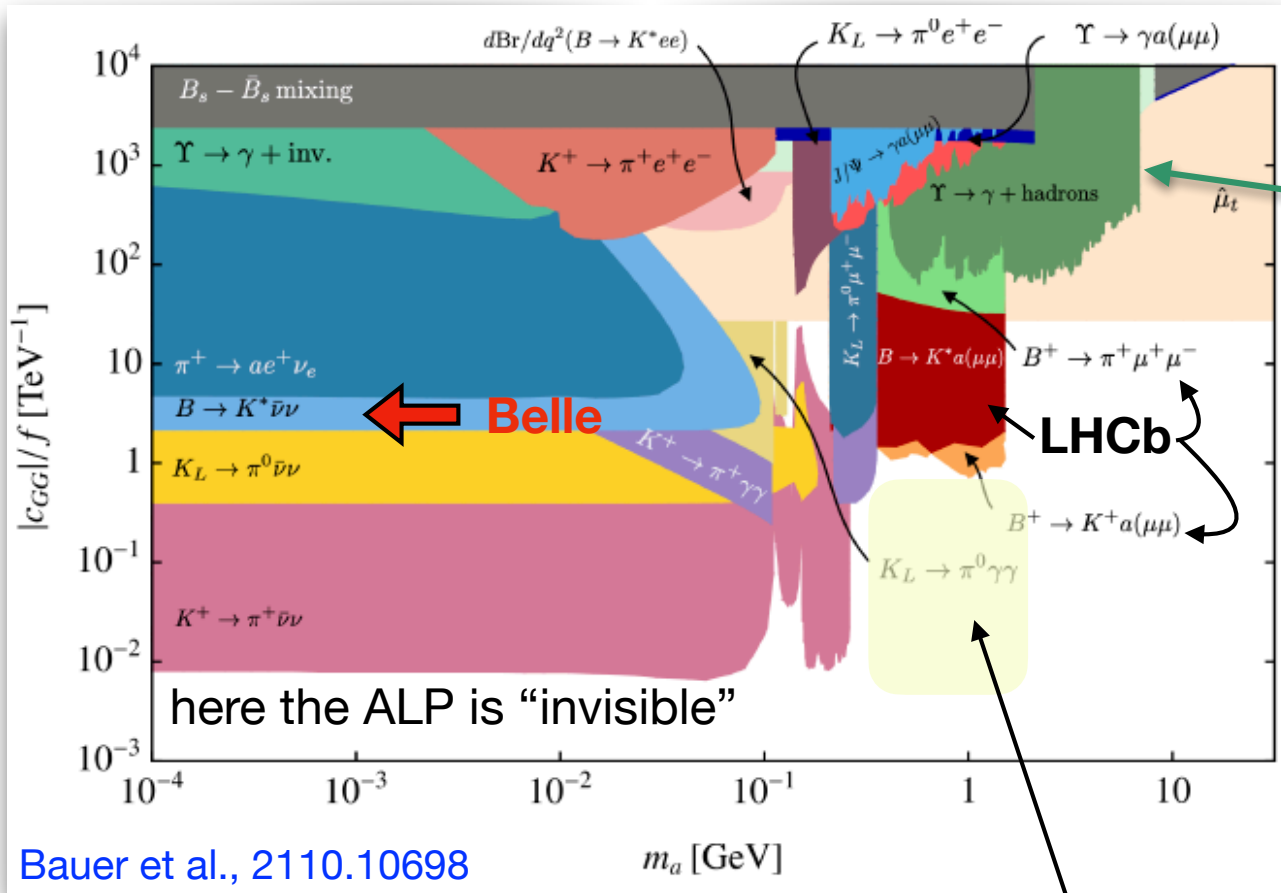
ALPs coupled to gluons



$$c_{GG} \frac{\alpha_S a}{4\pi f} G_{\mu\nu}^a \tilde{G}^{\mu\nu,a}$$



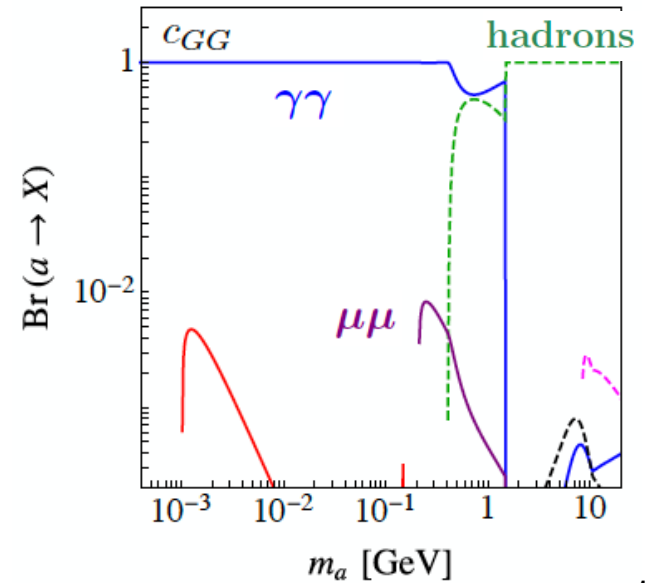
ALPs coupled to gluons



Babar, 1108.3549
 What about Belle II?
 $\Upsilon \rightarrow \gamma a, a \rightarrow \text{hadrons}$

What about Belle II,
 $B \rightarrow Ka, a \rightarrow \gamma\gamma$?

$$c_{GG} \frac{\alpha_S a}{4\pi f} G_{\mu\nu}^a \tilde{G}^{\mu\nu,a}$$



Chapter 4

Non-minimal models

1. Inelastic Dark Matter (non-minimal freeze out)
2. $L_\mu - L_\tau$ theories (“flavor specific” theories)

1.

Inelastic Dark Matter

Inelastic DM (IDM) models were initially proposed to explain the DAMA anomaly, while being consistent with Dark Matter direct detection bounds from CDMS

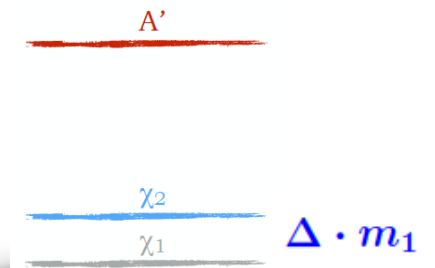
Tucker-Smith, Weiner, 0101138

$$-\mathcal{L} \supset m_D \eta \xi + \frac{1}{2} \delta_\eta \eta^2 + \frac{1}{2} \delta_\xi \xi^2 + \text{h.c.}$$

2-component Weyl spinors
with opposite charge under U(1)'

$$\chi_1 = i(\eta - \xi)\sqrt{2},$$

$$\chi_2 = (\eta + \xi)\sqrt{2}$$



Two states close in mass: $\Delta \equiv \frac{m_2 - m_1}{m_1} \sim \frac{\delta_\xi + \delta_\eta}{m_D} \ll 1$

Easy to get it small
since it is a U(1)'
breaking effect

1.

Inelastic Dark Matter

Inelastic DM (IDM) models were initially proposed to explain the DAMA anomaly, while being consistent with Dark Matter direct detection bounds from CDMS

Tucker-Smith, Weiner, 0101138

$$-\mathcal{L} \supset m_D \eta \xi + \frac{1}{2} \delta_\eta \eta^2 + \frac{1}{2} \delta_\xi \xi^2 + \text{h.c.}$$

2-component Weyl spinors
with opposite charge under U(1)'

The only relevant interaction is inelastic:

$$\mathcal{L} \supset \frac{ie_D m_D}{\sqrt{m_D^2 + (\delta_\xi - \delta_\eta)^2/4}} A'_\mu (\bar{\chi}_1 \gamma^\mu \chi_2 - \bar{\chi}_2 \gamma^\mu \chi_1)$$

$$\begin{aligned} \chi_1 &= i(\eta - \xi)\sqrt{2}, \\ \chi_2 &= (\eta + \xi)\sqrt{2} \end{aligned}$$

A'

The elastic piece is very small ($\delta_{\eta,\xi} \ll m_D$):

$$\mathcal{L} \supset \frac{e_D (\delta_\xi - \delta_\eta)}{\sqrt{4m_D^2 + (\delta_\xi - \delta_\eta)^2}} A'_\mu (\bar{\chi}_2 \gamma^\mu \chi_2 - \bar{\chi}_1 \gamma^\mu \chi_1)$$

χ_2

χ_1

$\Delta \cdot m_1$

Two states close in mass: $\Delta \equiv \frac{m_2 - m_1}{m_1} \sim \frac{\delta_\xi + \delta_\eta}{m_D} \ll 1$

Easy to get it small
since it is a U(1)'
breaking effect

1.

Inelastic Dark Matter

Inelastic DM (IDM) models were initially proposed to explain the DAMA anomaly, while being consistent with Dark Matter direct detection bounds from CDMS

Tucker-Smith, Weiner, 0101138

$$-\mathcal{L} \supset m_D \eta \xi + \frac{1}{2} \delta_\eta \eta^2 + \frac{1}{2} \delta_\xi \xi^2 + \text{h.c.}$$

2-component Weyl spinors with opposite charge under U(1)'

The only relevant interaction is inelastic:

$$\mathcal{L} \supset \frac{ie_D m_D}{\sqrt{m_D^2 + (\delta_\xi - \delta_\eta)^2/4}} A'_\mu (\bar{\chi}_1 \gamma^\mu \chi_2 - \bar{\chi}_2 \gamma^\mu \chi_1)$$

$$\begin{aligned} \chi_1 &= i(\eta - \xi)\sqrt{2}, \\ \chi_2 &= (\eta + \xi)\sqrt{2} \end{aligned}$$

A'

The elastic piece is very small ($\delta_{\eta,\xi} \ll m_D$):

$$\mathcal{L} \supset \frac{e_D (\delta_\xi - \delta_\eta)}{\sqrt{4m_D^2 + (\delta_\xi - \delta_\eta)^2}} A'_\mu (\bar{\chi}_2 \gamma^\mu \chi_2 - \bar{\chi}_1 \gamma^\mu \chi_1)$$

χ_2

χ_1

$\Delta \cdot m_1$

Two states close in mass: $\Delta \equiv \frac{m_2 - m_1}{m_1} \sim \frac{\delta_\xi + \delta_\eta}{m_D} \ll 1$

Easy to get it small since it is a U(1)' breaking effect

Abundance of χ_1 and χ_2 is determined by two coupled Boltzmann equations, that keep into account:

- * $\chi_1 \chi_2$ co-annihilation,
- * $\chi_2 f \rightarrow \chi_1 f$ inelastic scattering,
- * $\chi_2 \rightarrow \chi_1 + \text{SM decays}$

1.

IDM displaced signatures

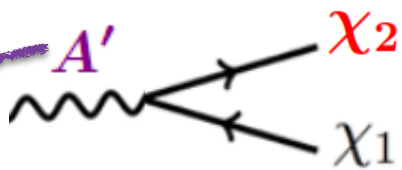
IDMs are rather hidden to direct detection experiments

Also CMB constraints are relaxed

The prime avenue to probe IDM is at high intensity experiments

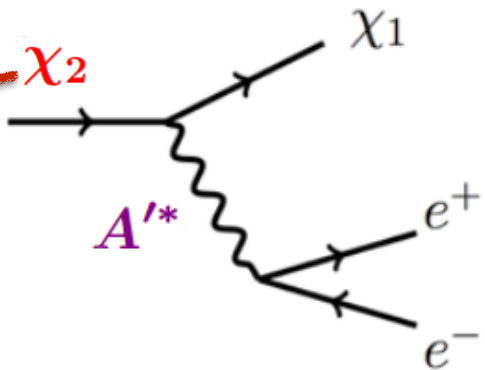
(see, however, Bramante et al., 1608.02662)

$$m_{\chi} < m_{A'}$$



Copiously produced at high intensity experiments (see yesterday's lecture)

with



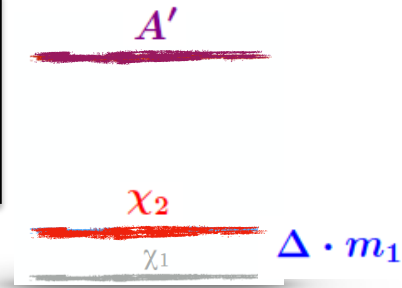
$$\Gamma(\chi_2 \rightarrow \chi_1 e^+ e^-) \simeq \frac{4\epsilon^2 \alpha_{\text{em}} \alpha_D \Delta^5 m_1^5}{15\pi m_{A'}^4}$$

Non-resonant decays

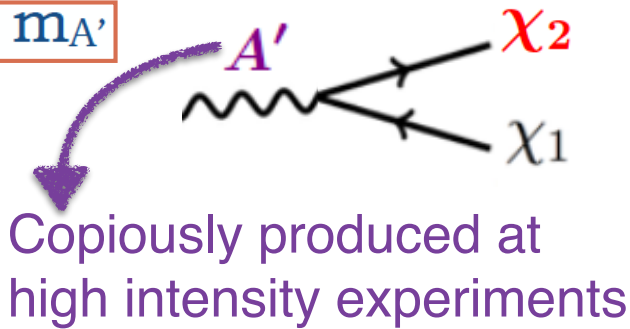
1.

IDM displaced signatures

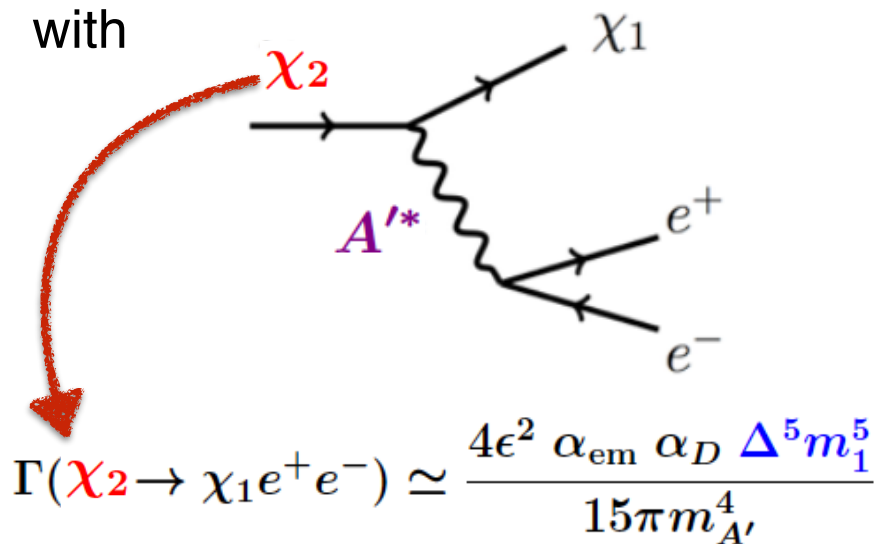
IDMs are rather hidden to direct detection experiments
 Also CMB constraints are relaxed
 The prime avenue to probe IDM is at high intensity experiments



$m_X < m_{A'}$



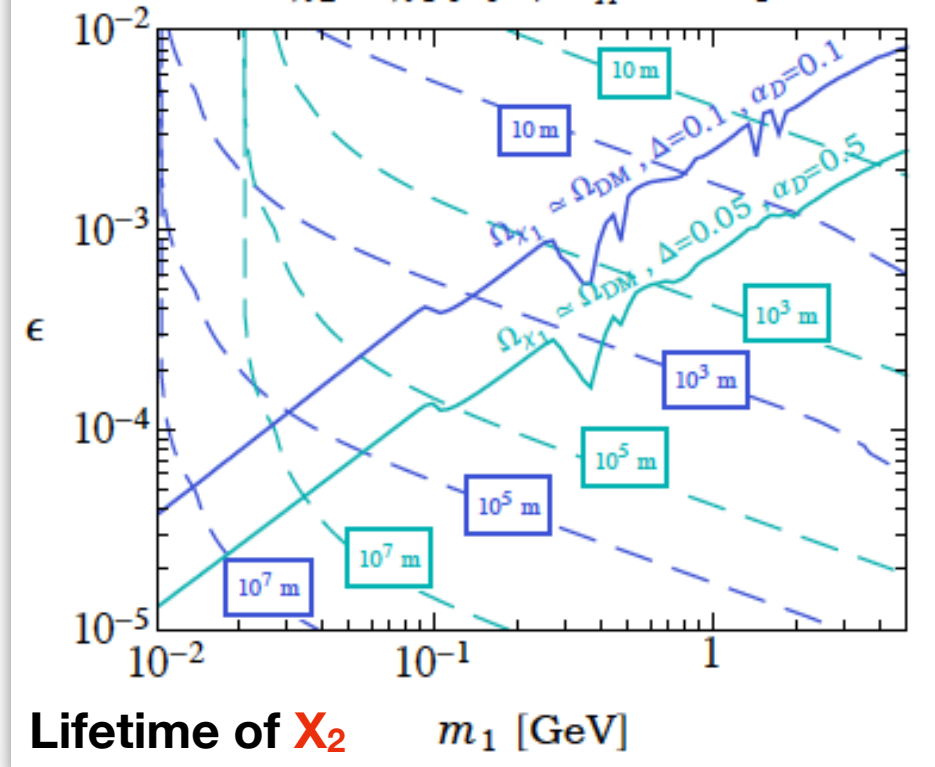
with



Non-resonant decays

Berlin, SG, Schuster, Toro, 1804.00661

$\chi_2 \rightarrow \chi_1 f \bar{f}, m_{A'} = 3 m_1$



Displaced decays

1.

New opportunities for B-factories

New proposed search for Belle-II:

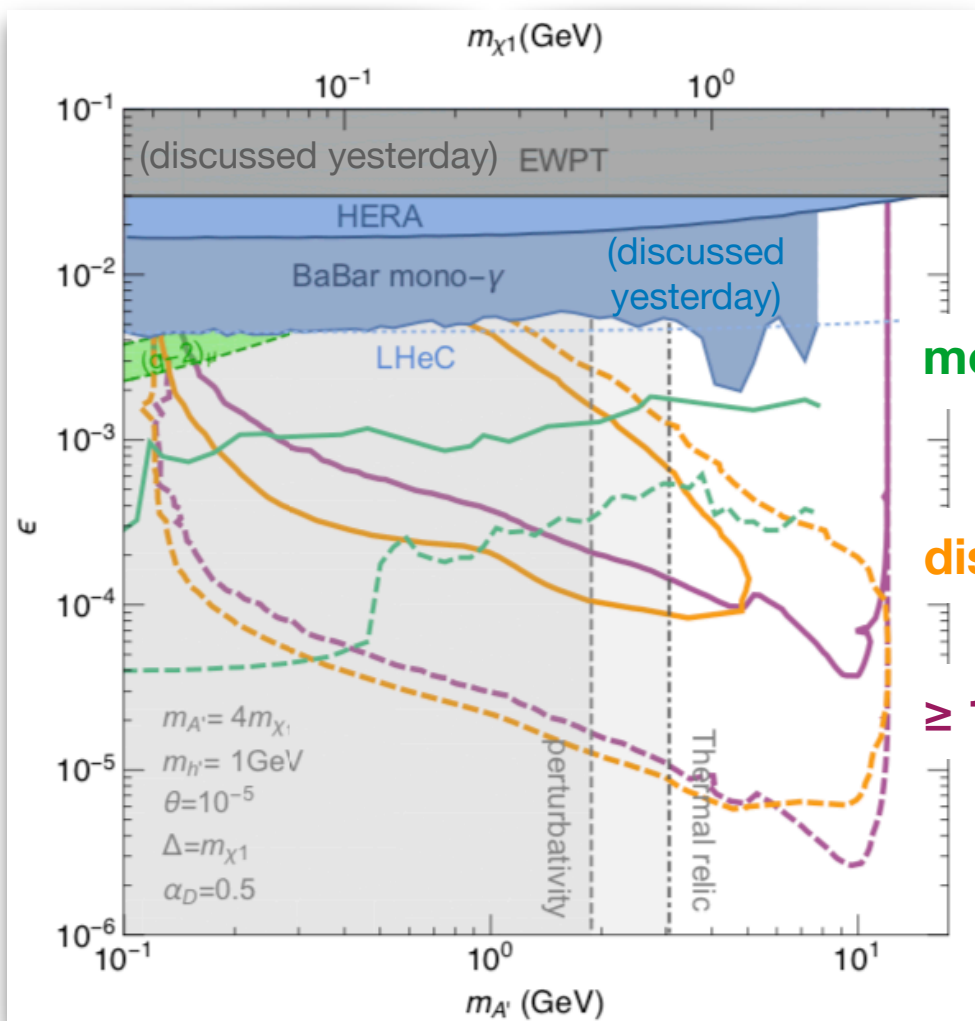
(Photon) + displaced tracks + missing energy

1.

New opportunities for B-factories

New proposed search for Belle-II:

(Photon) + displaced tracks + missing energy



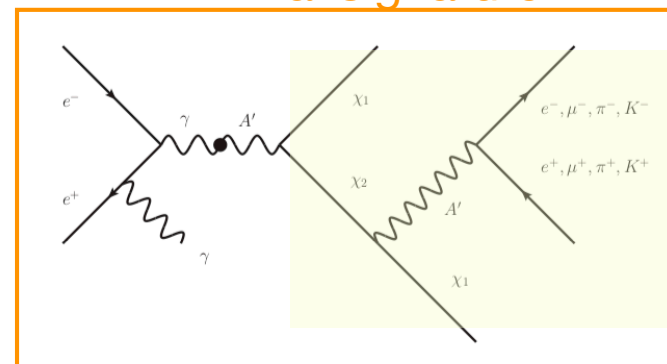
— 100/fb
 50/ab

mono-photon

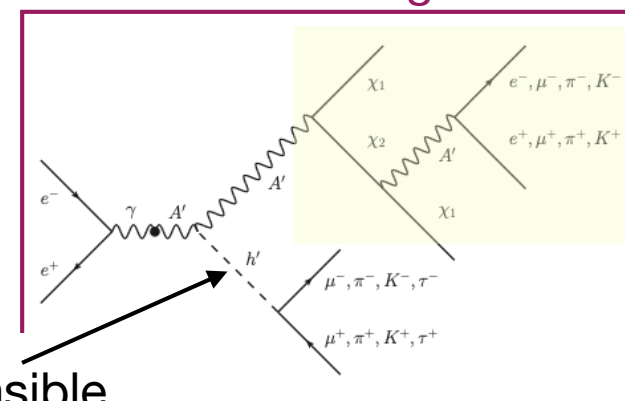
displaced+photon

≥ 1 displaced

“minimal signature”



“non-minimal signature”



Higgs responsible
 for the A' mass

Higgs-strahlung

Duerr et al. 2012.08595

2.

A new gauge symmetry for DM?

$$L_\mu - L_\tau$$

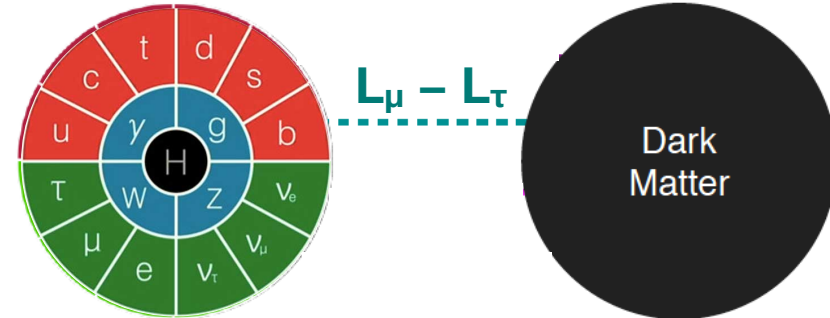
Simple example:

dark matter is a Dirac fermion

charged under $L_\mu - L_\tau$, $q_\chi g' \bar{\chi} \gamma^\mu \chi Z'_\mu$

Altmannshofer, SG, Profumo, Queiroz 1609.04026

(see also Kile et al. 1411.1407; Kim et al. 1505.04620; Baek 1510.02168 ...)



This symmetry is also motivated by
neutrino mass model building
+ anomalies in data ($b \rightarrow sll$, $(g-2)_\mu$)

If $m_{Z'} < 2m_{DM}$, the Z' will decay exclusively
to muons, taus, and neutrinos

2.

A new gauge symmetry for DM?

$$L_\mu - L_\tau$$

Simple example:

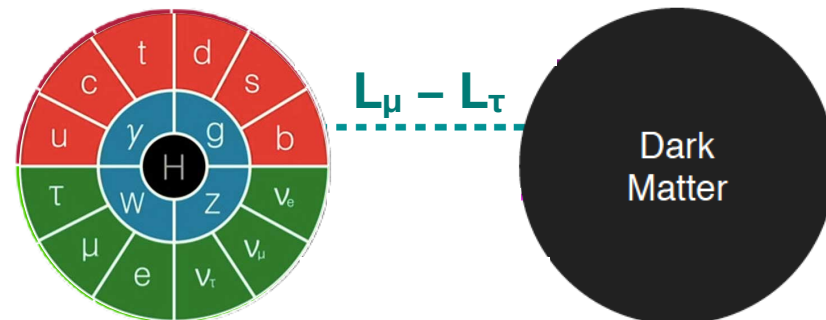
dark matter is a Dirac fermion

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Altmannshofer, SG, Profumo, Queiroz 1609.04026

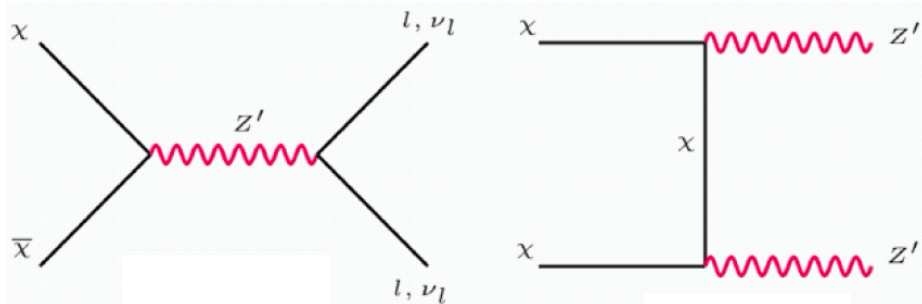
(see also Kile et al. 1411.1407; Kim et al. 1505.04620;

Baek 1510.02168 ...)



Possible signals at:

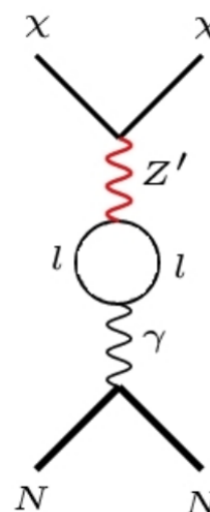
1. DM indirect detection experiments



Main constraints from CMB

If $m_{Z'} < 2m_{DM}$, the Z' will decay exclusively to muons, taus, and neutrinos

2. DM direct detection experiments



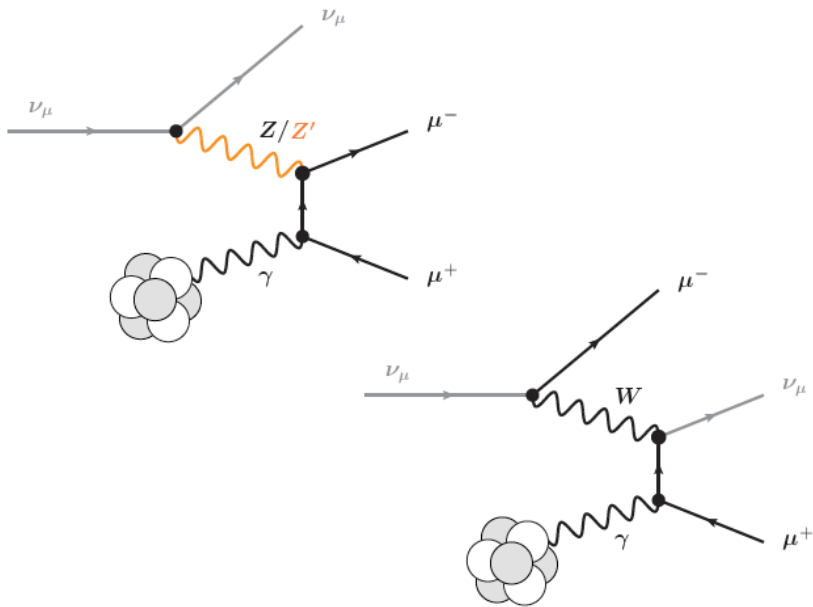
It **can be sizable**, despite the loop suppression

Constraints from LUX, Xenon, and PandaX

2.

Tests at neutrino experiments: CCFR, CHARM experiments

Neutrino induced $\mu^+\mu^-$ production in the Coulomb field of a heavy nucleus:
“neutrino trident production”



Altmannshofer, SG, Pospelov, Yavin, 1406.2332

Z' contribution to the cross section:

$$\frac{\sigma}{\sigma_{\text{SM}}} \simeq \frac{1 + \left(1 + 4s_W^2 + \frac{2v^2(g')^2}{M_{Z'}^2}\right)^2}{1 + (1 + 4s_W^2)^2}$$

(in the approximation of heavy Z')

Measurements in the early '90s
by CCFR and CHARM:

$$\sigma/\sigma_{\text{SM}} = 0.82 \pm 0.28$$

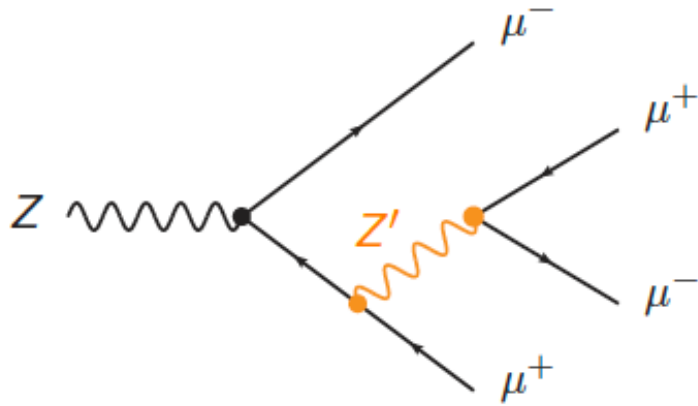
(CCFR, PRL66 (1991) 3117)

Not 100% clear
how solid this
measurement is

2.

Searches at the LHC

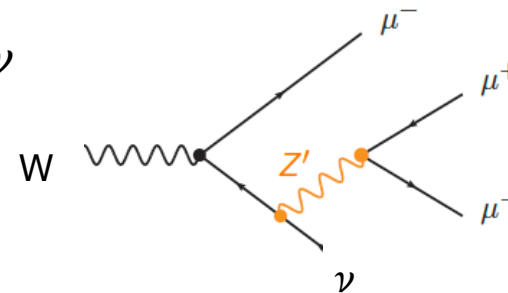
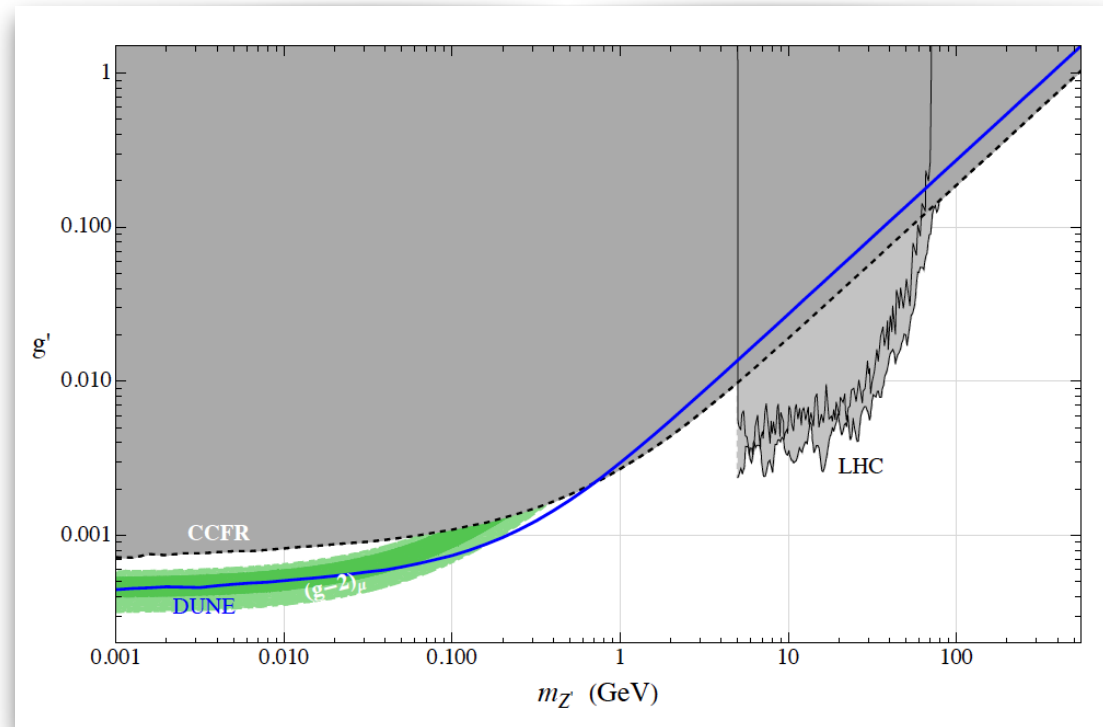
Bounds from the measured $Z \rightarrow 4\mu$ branching ratio



Recent dedicated CMS + ATLAS searches for the $L_\mu - L_\tau$ gauge boson, (1808.03684, 2402.15212)

Combination with $W \rightarrow Z' \mu \nu \rightarrow (\mu \mu) \mu \nu$

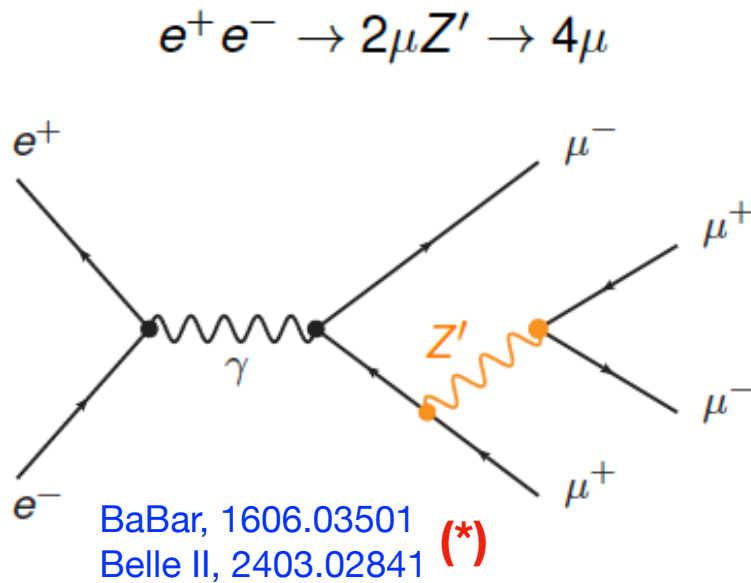
(Updated) Altmannshofer, SG, Martin-Albo, Sousa, Wallbank, 1902.06765



2.

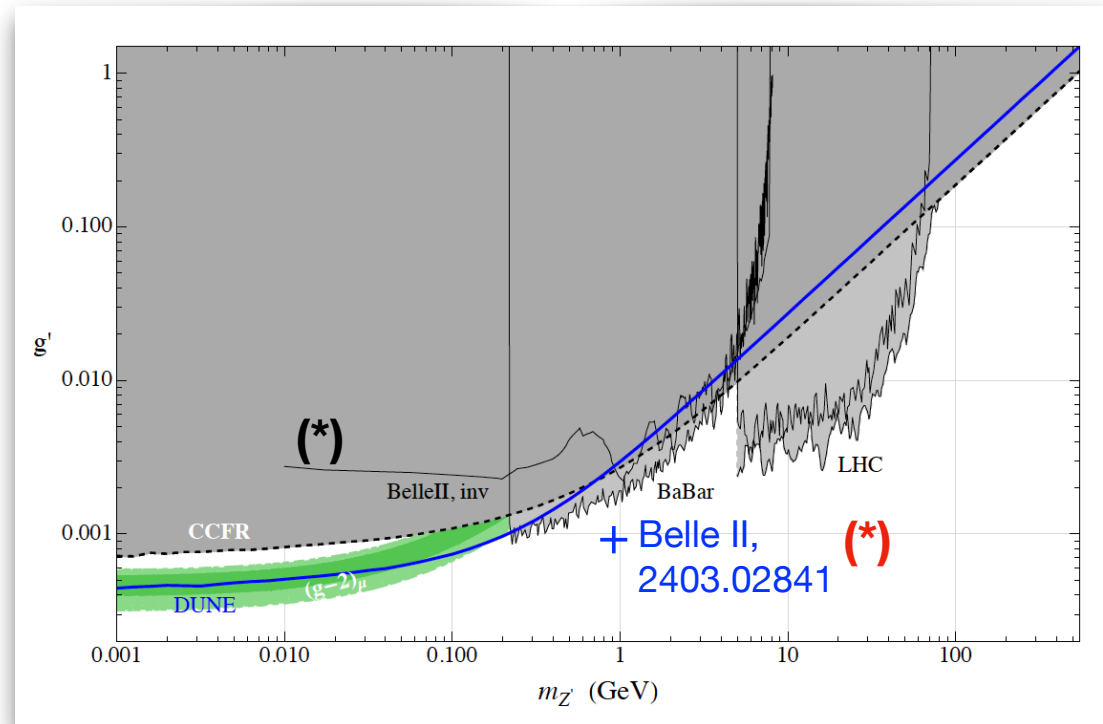
Searches at B-factories

B-factories can search for the light Z' produced with muons



Additional Belle II search [2212.03066](#)
for $e^+e^- \rightarrow \mu^+\mu^- + Z'$, $Z' \rightarrow \nu\nu$ (*)
(particularly relevant for $m_{Z'} < 2m_\mu$)

(Updated) Altmannshofer, SG, Martin-Albo,
Sousa, Wallbank, 1902.06765

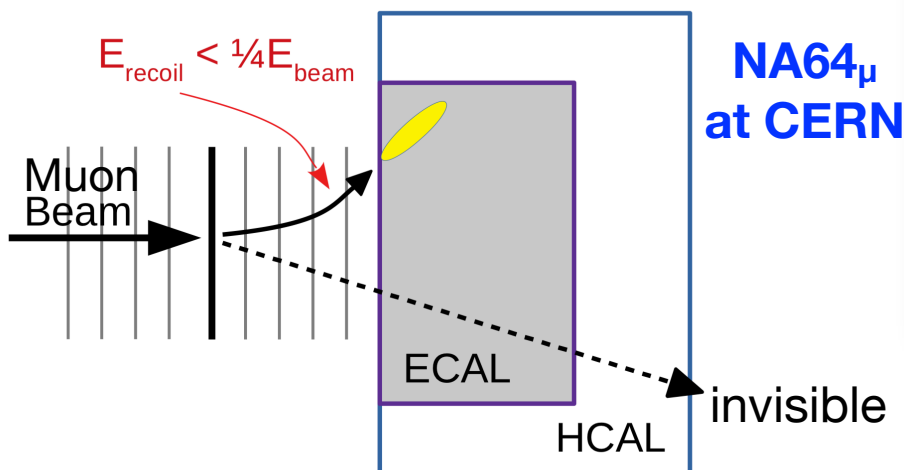
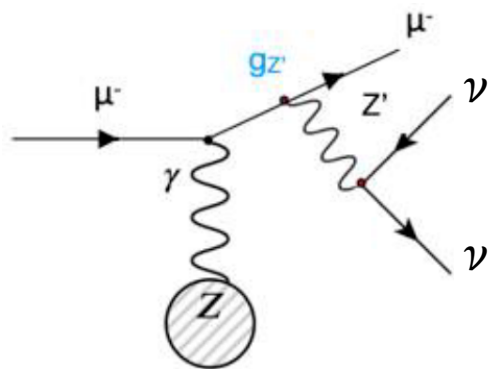


2.

Searches at fixed target experiments

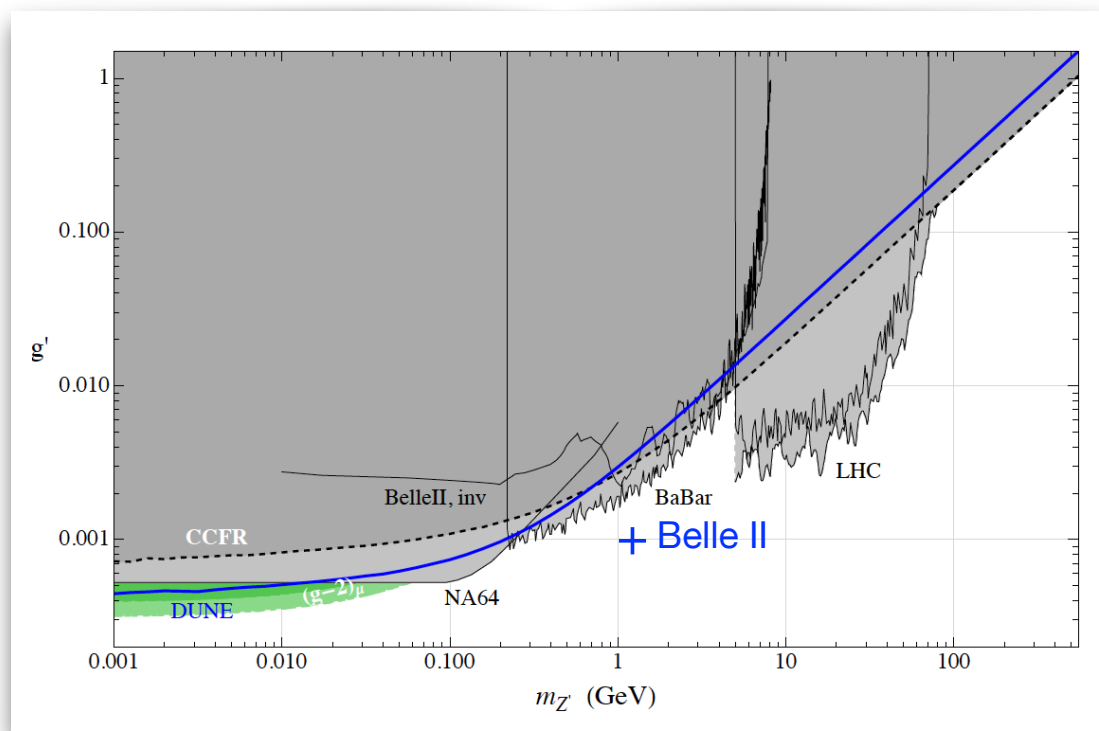
High intensity fixed target experiments can produce an invisible Z'

(Updated) Altmannshofer, SG, Martin-Albo, Sousa, Wallbank, 1902.06765



missing energy-momentum search

2401.01708

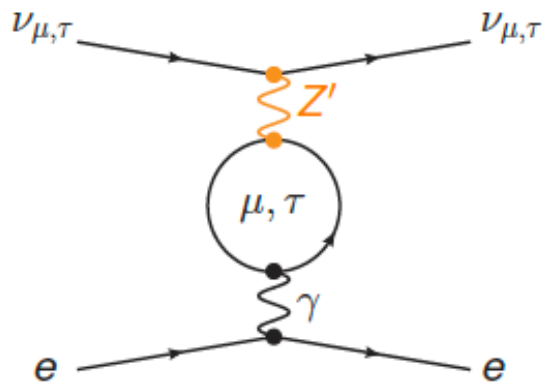


2.

Tests at neutrino experiments

Borexino

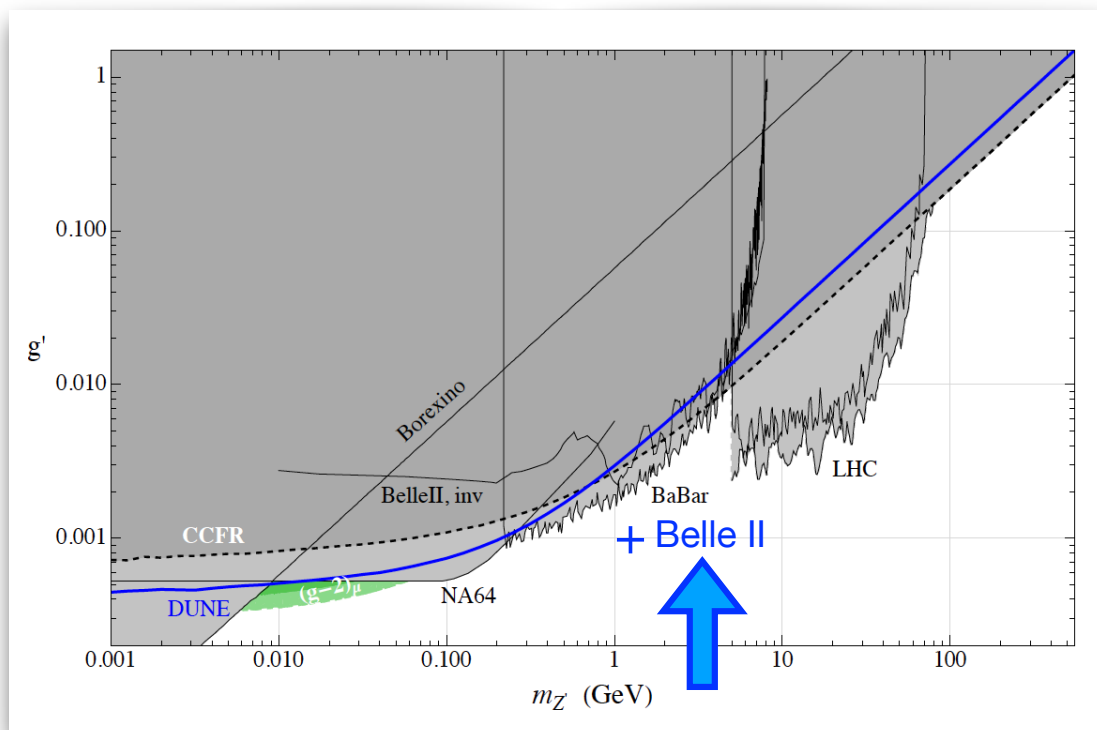
Bounds from measurements of solar neutrino-electron scattering



tiny momentum transfer
⇒ Z' can mix with the SM photon

relevant constraint at low masses from the Borexino experiment

(Updated) Altmannshofer, SG, Martin-Albo, Sousa, Wallbank, 1902.06765



Belle II probes a mass range that is not probed otherwise

What we have learnt

Dark sectors are ubiquitous.

DM and the strong CP problem are two of the several motivations.

Experimental targets arise in these models (i.e. how small the couplings we should aim to probe are)

Minimal models

- ▶ dark photon
- ▶ dark scalar
- ▶ sterile neutrino

- visible
- invisible

- * axion-like-particles
(several independent couplings with the SM)

Beyond the minimal models

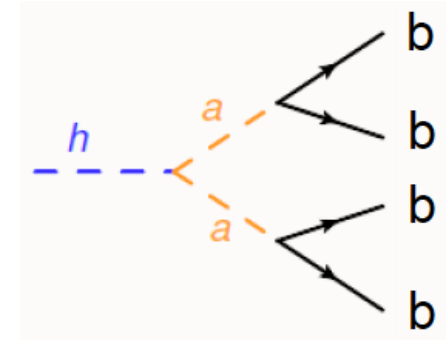
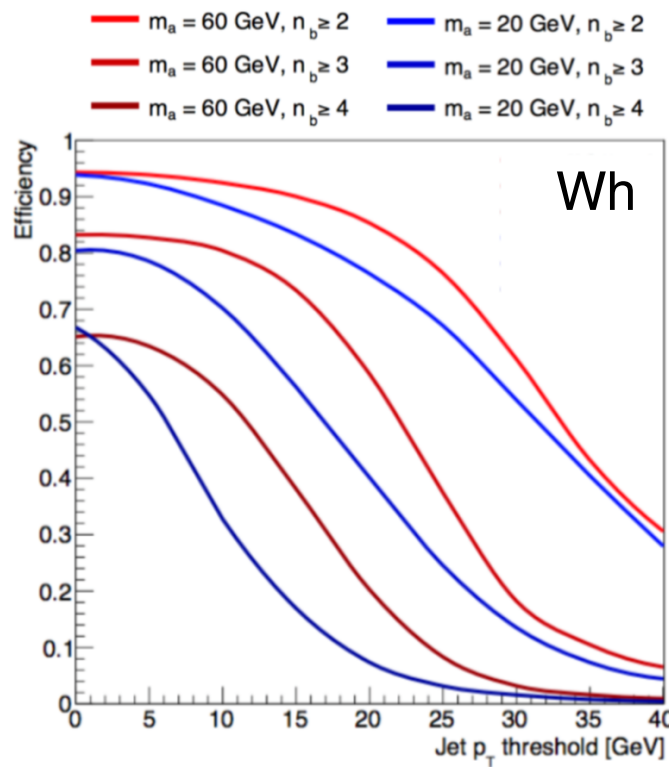
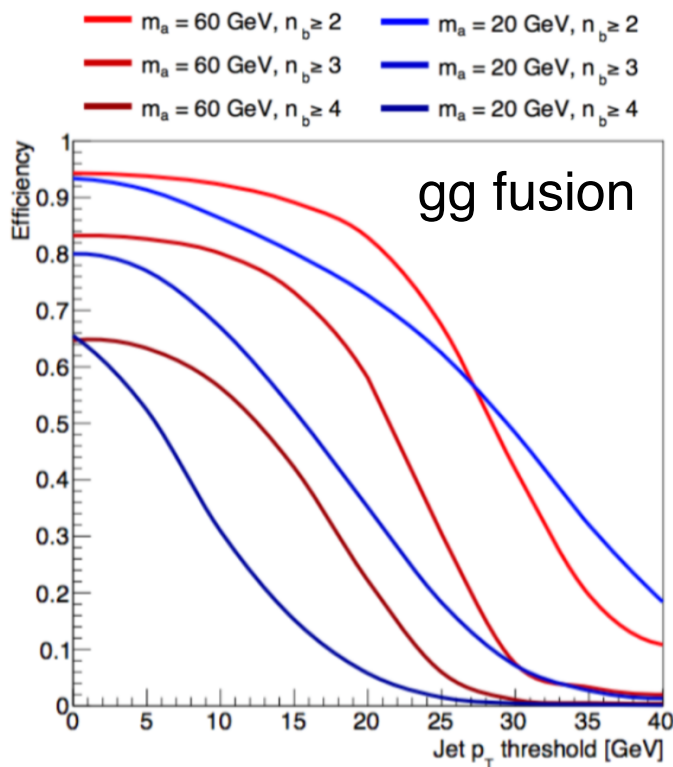
- ▶ Inelastic Dark Matter
- * $L_\mu - L_\tau$

- semivisible

The challenge for the LHC

To be sensitive to light scalars and Higgs exotic decays, dedicated studies of **trigger strategies** are needed!

Let us take, for example, the challenging decay mode $h \rightarrow 4b$



Risk of losing the signal already at the trigger level



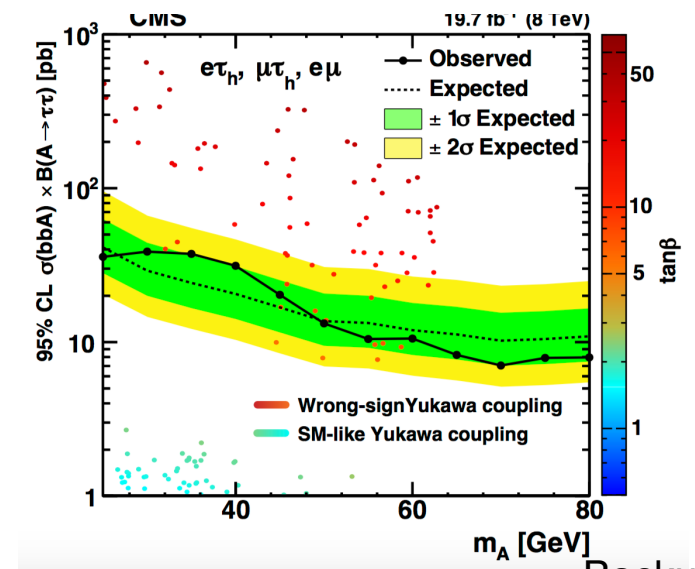
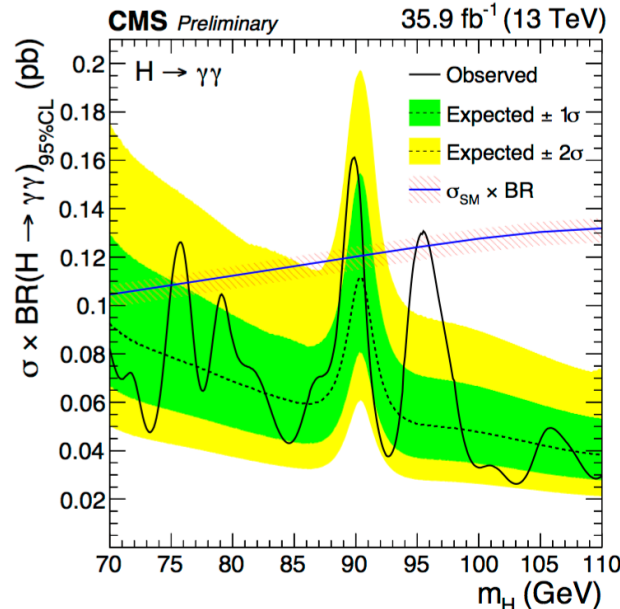
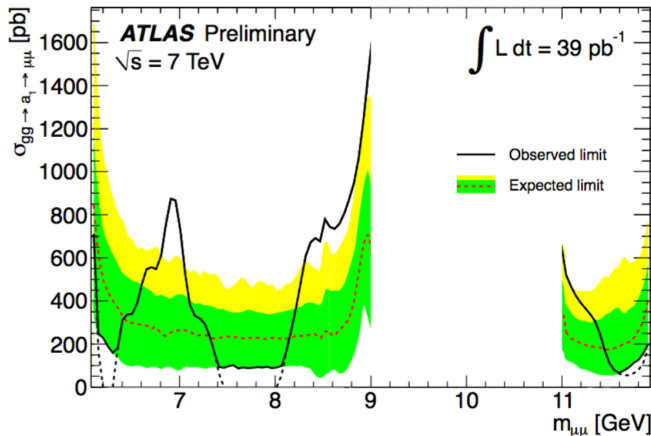
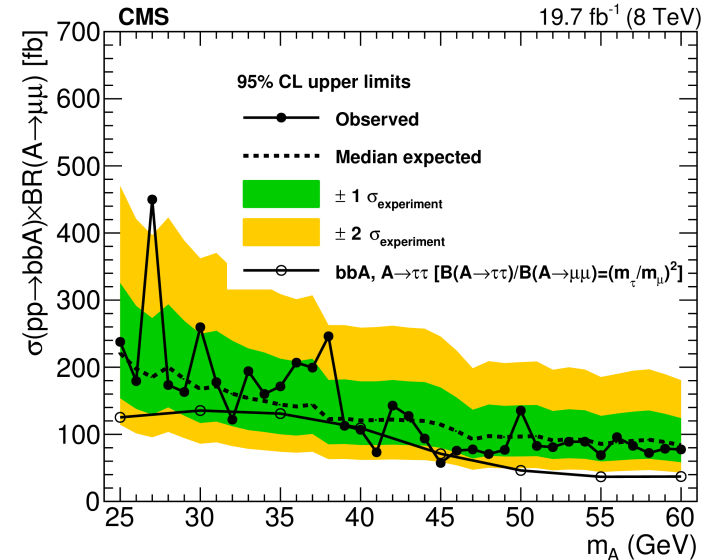
From the LHC Higgs cross section working group, Yellow report 4, 1610.07922

LHC production of the dark scalar

The production cross section $\sigma_s = \text{production_SM Higgs} \times \sin^2(\theta_s)$
 Same production modes as for the SM Higgs

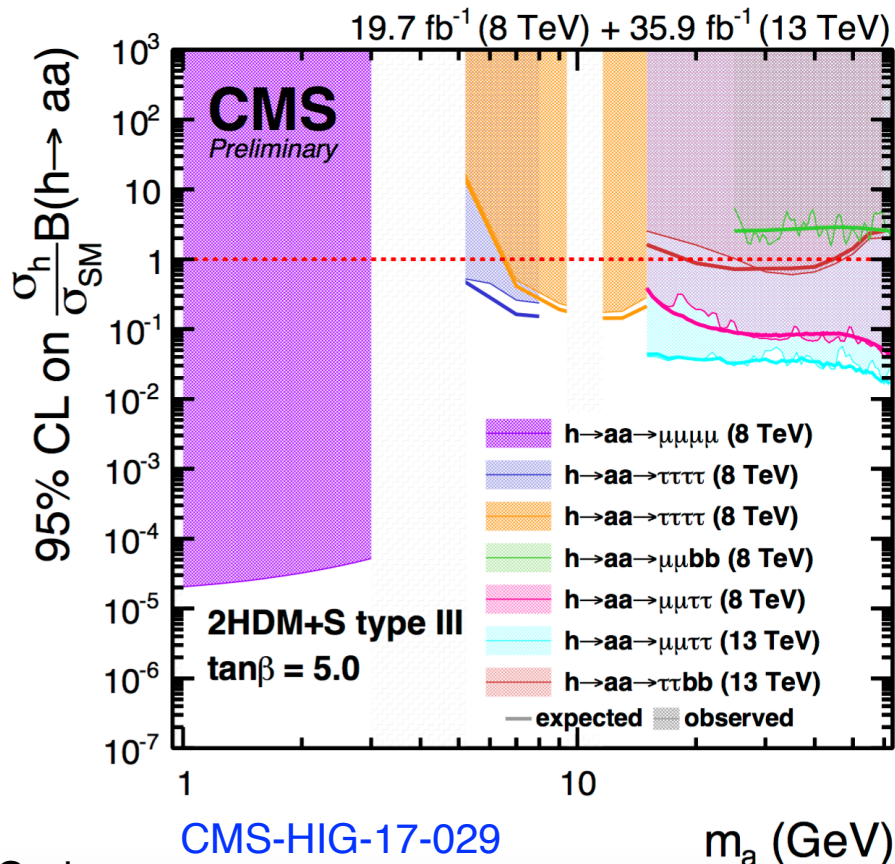
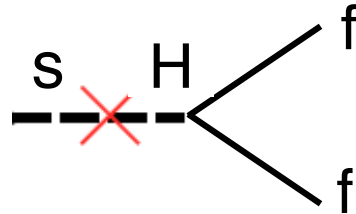
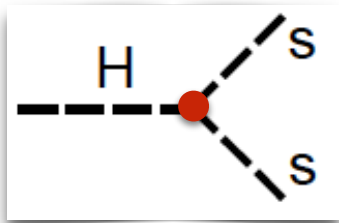
Only a few LHC searches have been performed.

- Examples are $bb\bar{s}, s \rightarrow \mu\mu$
- $ggs, s \rightarrow \mu\mu$,
- $ggs, s \rightarrow \gamma\gamma$
- $bb\bar{s}, s \rightarrow \text{tautau}$



Higgs exotic decays to dark scalars

The scalar can decay thanks to its mixing with the Higgs



Many possible signatures to look for

- $gg \rightarrow h \rightarrow ss \rightarrow 4b$
- $gg \rightarrow h \rightarrow ss \rightarrow 2b \ 2\tau$
- $gg \rightarrow h \rightarrow ss \rightarrow 2b \ 2\mu$
- $gg \rightarrow h \rightarrow ss \rightarrow 4\tau$
- $gg \rightarrow h \rightarrow ss \rightarrow 4\mu$
- $gg \rightarrow h \rightarrow ss \rightarrow 2\tau \ 2\mu$
- ...

both **prompt** & **displaced**

+ sub-leading production modes of the Higgs boson

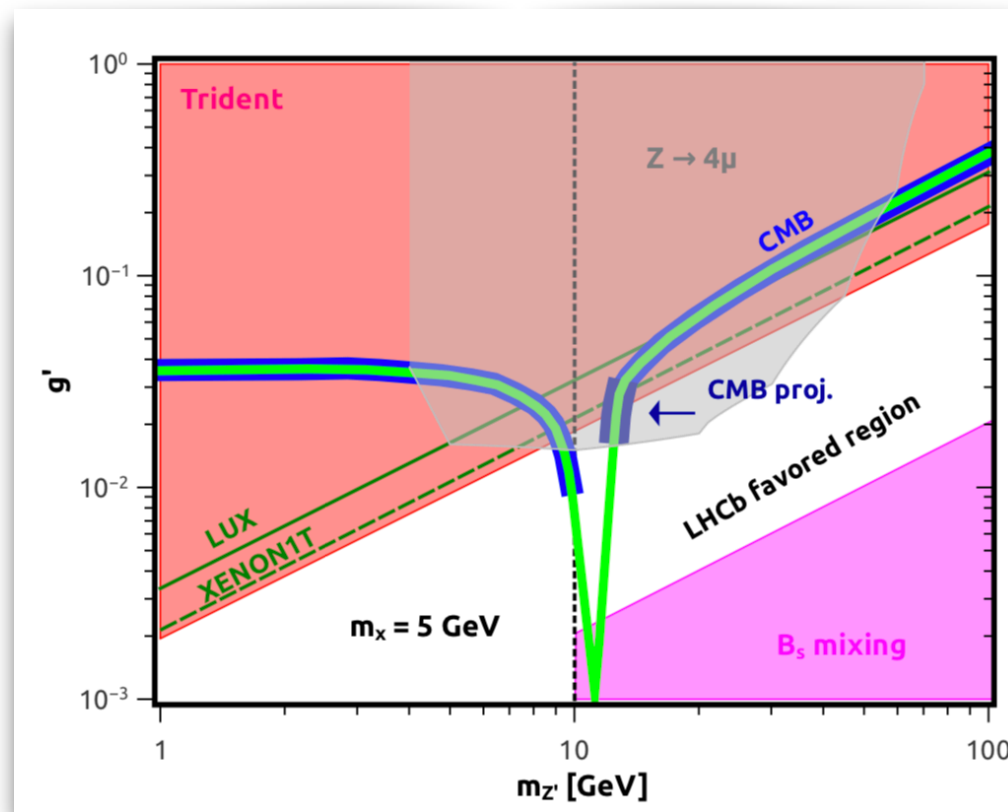
e.g. $qq \rightarrow Zh \rightarrow Z(ss) \rightarrow Z(4b)$

...

$L_\mu - L_\tau$ and dark matter

$L_\mu - L_\tau$ model + Dirac fermion DM

Altmannshofer, SG, Profumo, Queiroz, 1609.04026



Towards fully probing this model.

Only DM in $\sim(5-20)$ GeV is still viable