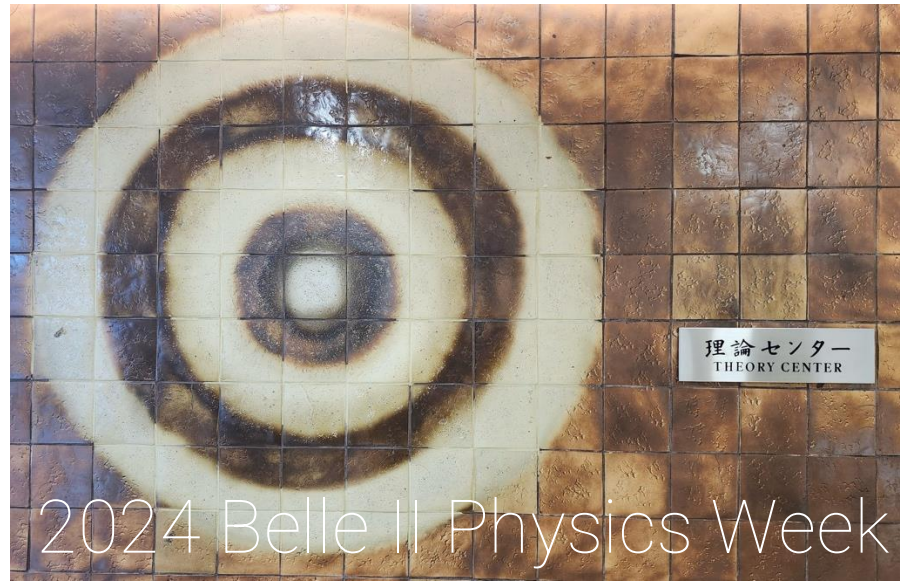


Dark Matter, Baryogenesis, & Naturalness @ Belle II

Robert McGehee

UNIVERSITY OF MINNESOTA



2024 Belle II Physics Week

What am I talking about?

New Signatures @ Belle II?

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New Signatures @ Belle II?

→ Nope. Cf Gilly's talk, including our "Mesogenesis" papers

[G. Elor, **RM PRD** 2021]

[F. Elahi, G. Elor, **RM PRD** 2022]

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Models Discoverable @ Belle II!

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Dark Matter Baryogenesis Naturalness

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[P. Bittar, S. Koren, **RM** WIP]

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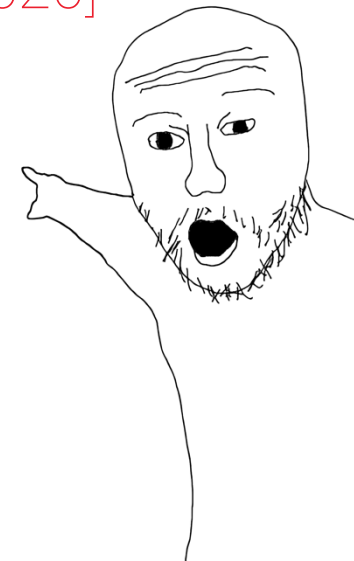
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R McGee

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Take-Home Message

The theory motivation for the BSM searches at Belle II is robust.

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[K. Harigaya, RM, H. Murayama, K. Schutz JHEP 2020]

Strongly interacting massive particles through the axion portal

[Y. Hochberg, E. Kuflik, RM, H. Murayama, K. Schutz PRD 2018]

Strongly interacting massive particles through the axion portal

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Complementary model to Stefania's!
message = SIMPs are cool for Belle II !!!

Strongly Interacting Massive Particles (SIMP)s

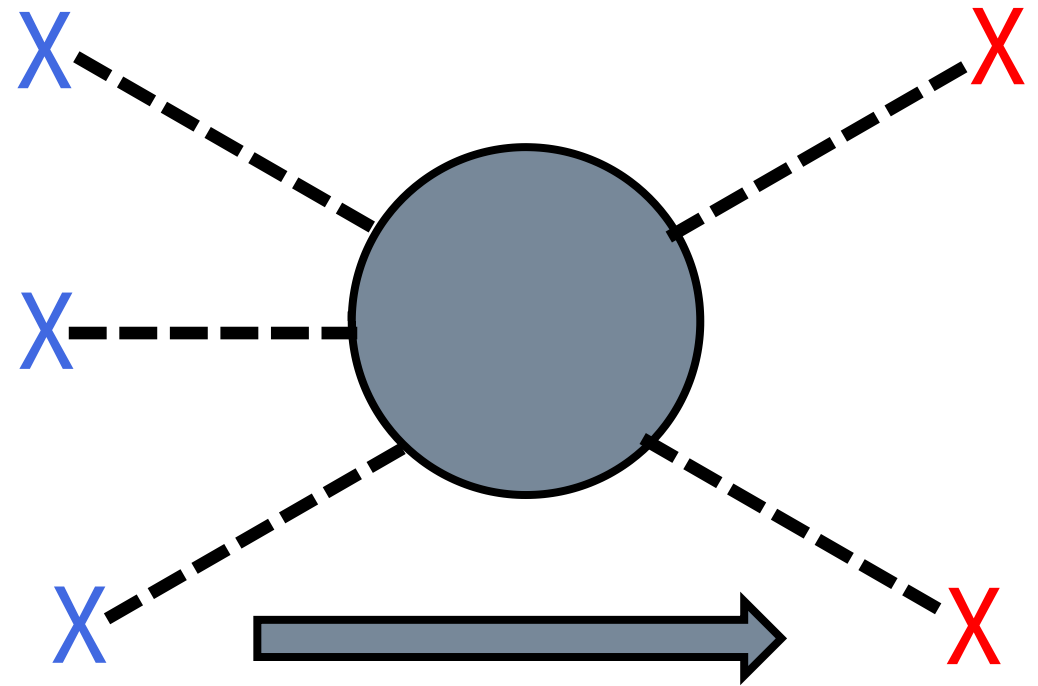
[Y. Hochberg et al. PRL 2014]

Relic density set by $3 \rightarrow 2$

Equilibrium with SM

Strong self interactions

MeV to GeV masses



The SIMPlEst Miracle

[Y. Hochberg et al. PRL 2015]

$\text{Sp}(\mathbf{N}_c)$ gauge theory

$2 \mathbf{N}_f$ Weyl fermions

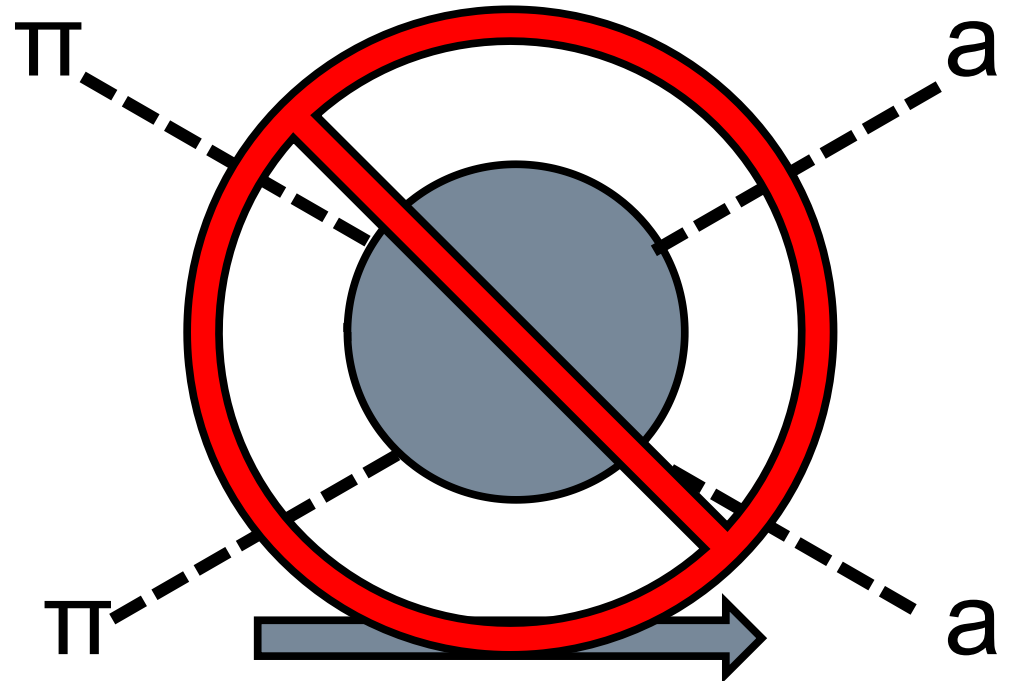
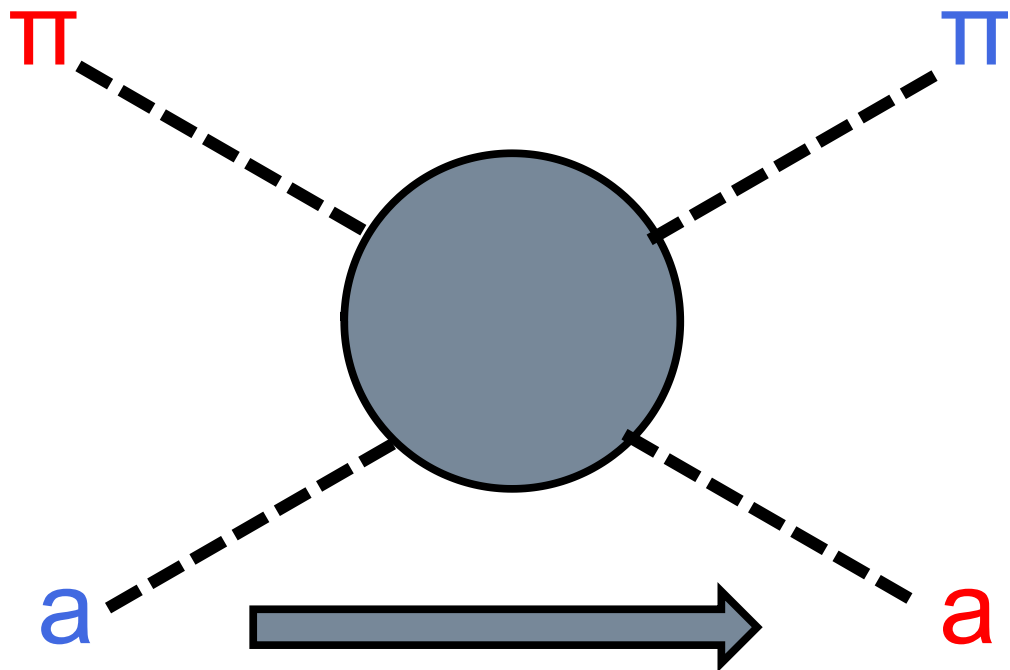
Chiral symmetry breaking

Wess-Zumino-Witten (WZW)

$$\mathcal{L}_{\text{WZW}} = \frac{8N_c}{15\pi^2 f_\pi^5} \epsilon^{\mu\nu\rho\sigma} \epsilon_{abcde} \pi^a \partial_\mu \pi^b \partial_\nu \pi^c \partial_\rho \pi^d \partial_\sigma \pi^e$$

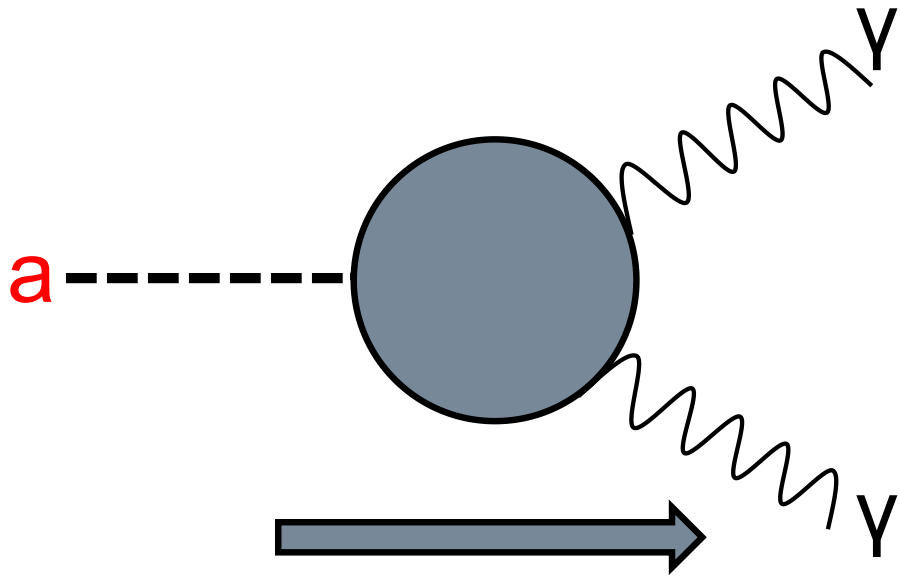
SIMP-Axion Coupling

$$\mathcal{L}_{aq} = -\frac{1}{2}m_Q e^{ia/f_{a\pi}} J^{ij} q_i q_j + \text{h.c.} \xrightarrow{\text{Chiral SB}} \mathcal{L}_{aq} = \frac{m_\pi^2}{4f_{a\pi}^2} a^2 \pi^2 + \dots$$

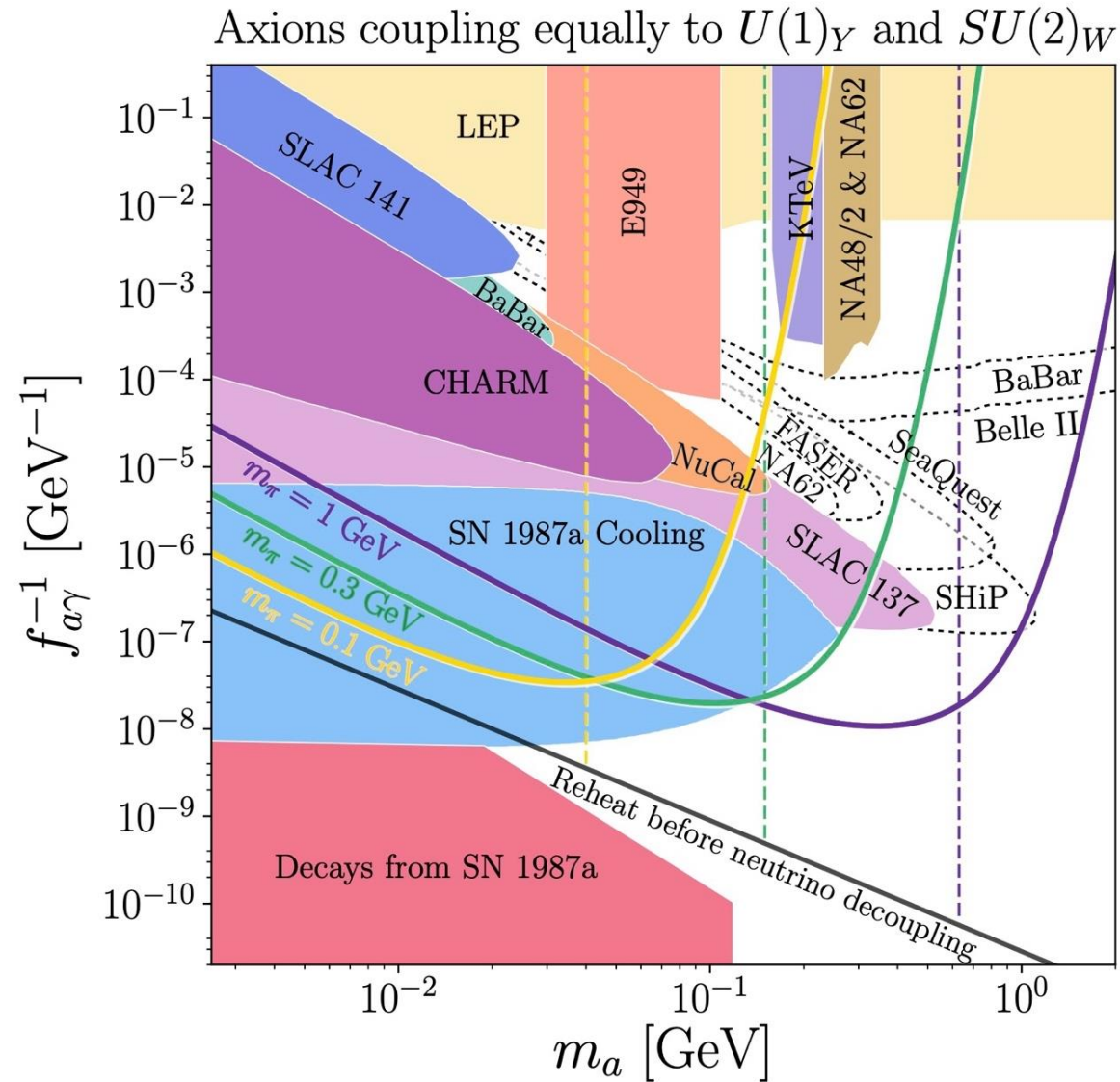


Axion-Photon Coupling

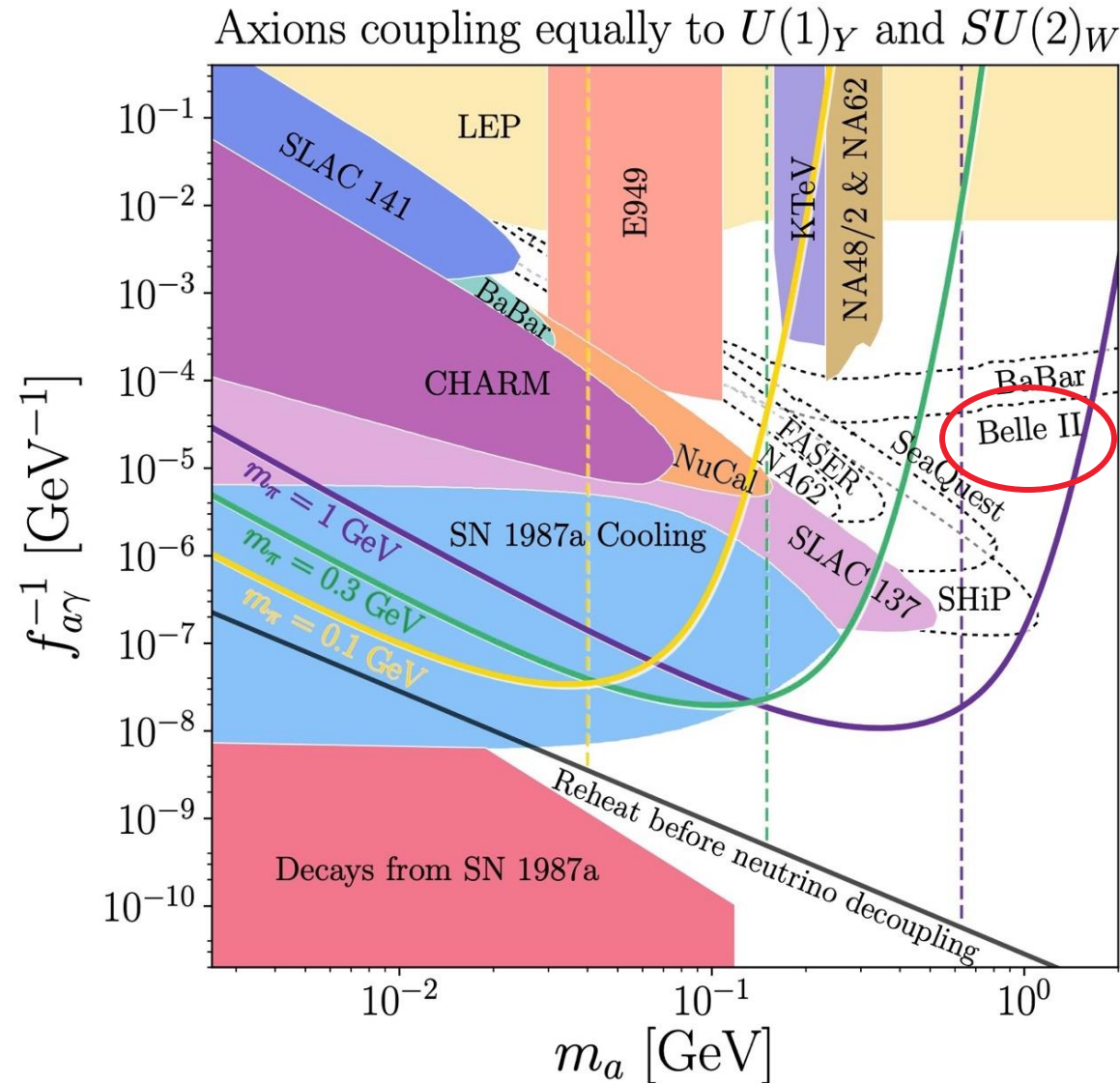
$$\mathcal{L}_{a\gamma} = \frac{1}{4f_{a\gamma}} a F^{\mu\nu} \tilde{F}_{\mu\nu}$$



Viability Axion-Photon Parameter Space



Viability Axion-Photon Parameter Space



Belle II will hopefully probe the majority of the viable parameter space!

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[K. Harigaya, RM, H. Murayama, K. Schutz JHEP 2020]

Asymmetric matter from a dark first-order phase transition

[E. Hall, T. Konstandin, **RM**, H. Murayama, PRD 2023]

Baryogenesis & Asymmetric DM

Ω_b requires explanation

$\Omega_c \approx 5\Omega_b$ suggests DM may also be **asymmetric**

Common origin?

ADM + SM gauge copy \rightarrow

Mirror world models

See e.g. hep-ph/0312335, astro-ph/0407623,
1805.06876, 1811.10232, 1907.03404

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$$SU(3)' \times SU(2)' \times U(1)'$$

$$\Phi_1, \Phi_2$$

$$Q', u'_R, d'_R$$

$$L', e'_R, N'_R \longleftarrow Y_e \text{ large}$$

Dark-Sector Baryogenesis

1. $SFOPT$ + CP -violating potential + sphalerons in DS
2. 2 Higgs doublets, EW-like baryogenesis in DS $\rightarrow B'+L'$

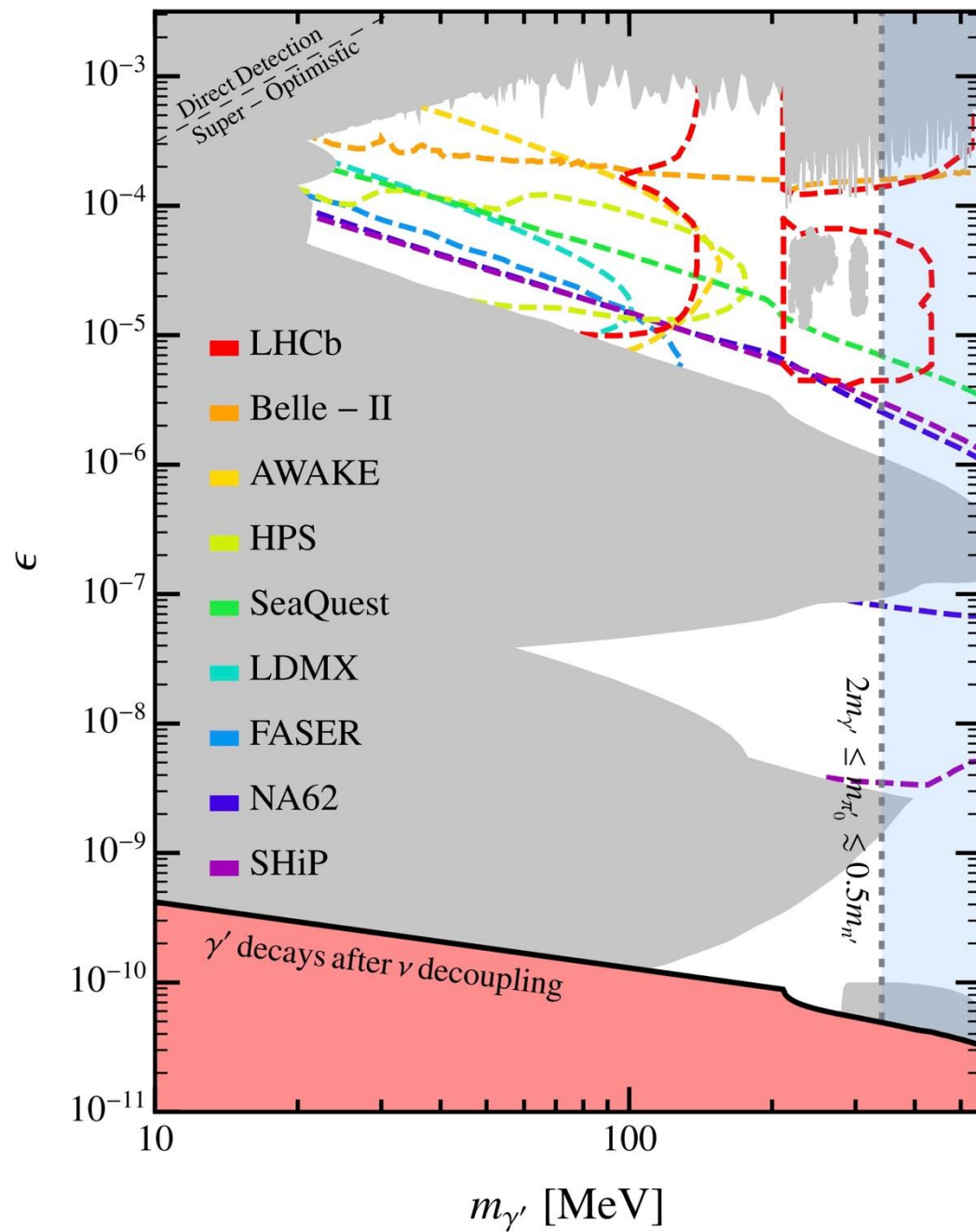
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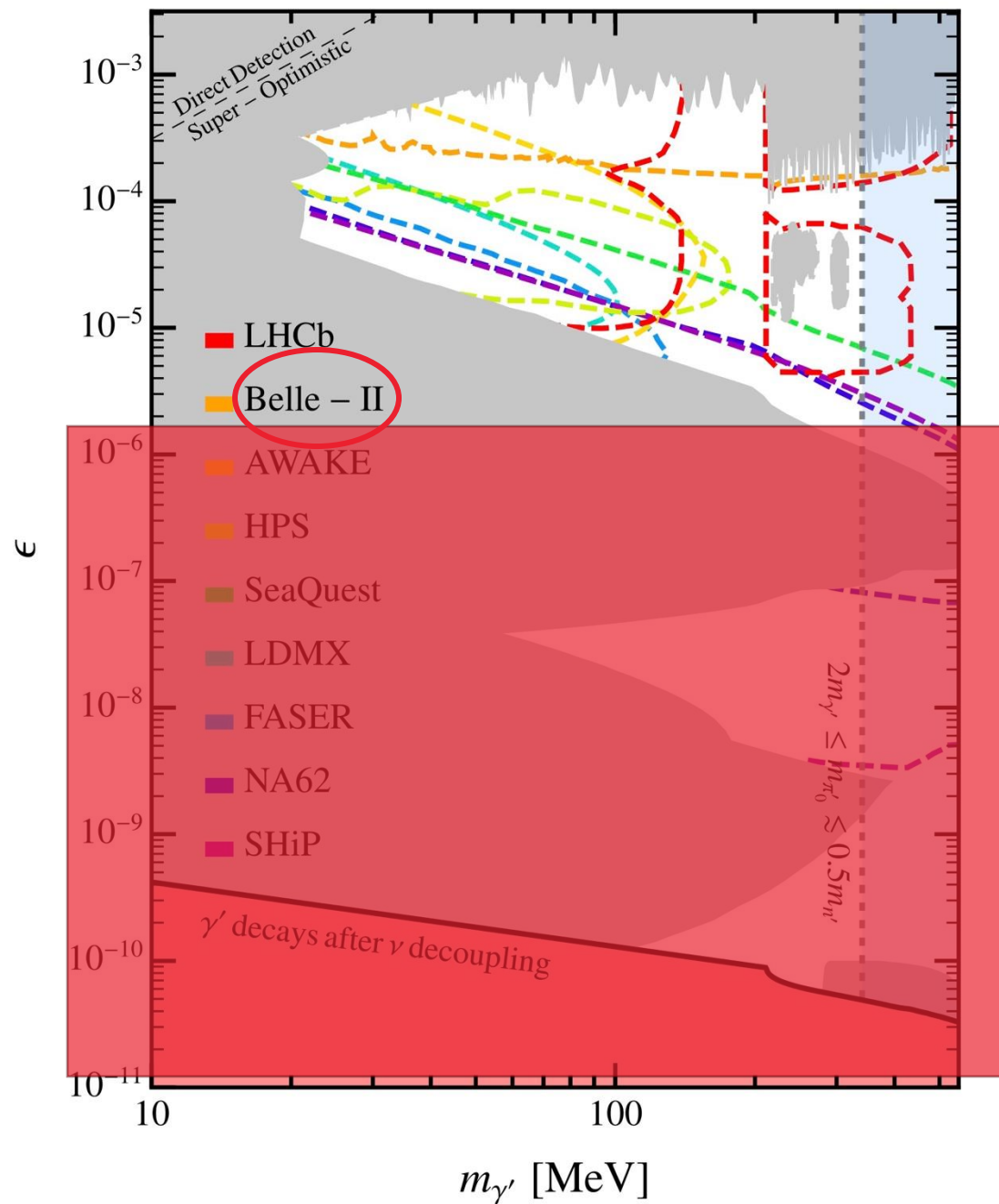
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3. Dark leptons in equilibrium via $neutrino$ portal
4. SM $sphalerons$ then generate some B
5. Symmetric part of dark hadrons $annihilate$ to dark photons, which transfer excess entropy to SM
6. Remaining B' forms (part of) ADM

Dark Neutron Dark Matter



Dark Neutron Dark Matter



Theory bias:
kinetic mixing
generated @ 1
loop

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Naturalness

[K. Harigaya, **RM**, H.
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JHEP 2020]

A predictive mirror twin Higgs with small Z_2 breaking

[K. Harigaya, RM, H. Murayama, K. Schutz JHEP 2020]

Mirror Twin Higgs



[Z. Chacko, H.-S. Goh, R. Harnik PRL 2006]

\mathbf{Z}_2 mirror copy of SM introduced (twin sector)

SU(4) global symmetry permits SM Higgs to be PNGB

\mathbf{Z}_2 protects Higgs mass up to ~ 5 -10 TeV cutoff

The N_{eff} Problem

$$\rho_r = \left(1 + \frac{7}{8} \left(\frac{4}{11} \right)^{4/3} N_{\text{eff}} \right) \rho_\gamma$$

$$N_{\text{eff}}^{\text{SM}} = 3.046$$

$$N_{\text{eff}}^{\text{Pl}} = 2.99^{+0.34}_{-0.33}$$

$$\Delta N_{\text{eff}} = N_{\text{eff}}^{\text{Pl}} - N_{\text{eff}}^{\text{SM}} < 0.284$$

[Planck Collaboration A&A 2020]

$$\Delta N_{\text{eff}}^{\text{MTH}} \sim 5.6$$

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[Planck Collaboration A&A 2020]

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Fraternal Twin Higgs: only 3rd gen. twin fermions & no U(1)'

[N. Craig, A. Katz, M. Strassler, and R. Sundrum, JHEP 2015]

Raise twin neutrino masses,
raise the twin photon mass,
inject entropy asymmetrically
after decoupling, or adjust twin
Yukawas

See 1601.07181, 1609.05589, 1611.07975,
1611.07977, 1703.06884, 1706.05548, 1803.03263,
1805.09345, 1805.12139, 1904.10468, 1905.00861

A New Solution

Raise (non-top) twin Yukawas

Twin neutrinos decouple earlier,
as high as a few GeV

Early decoupling dilutes
contribution to N_{eff}

Stuckelberg mass for dark
photon allows it to transfer
excess entropy to the SM

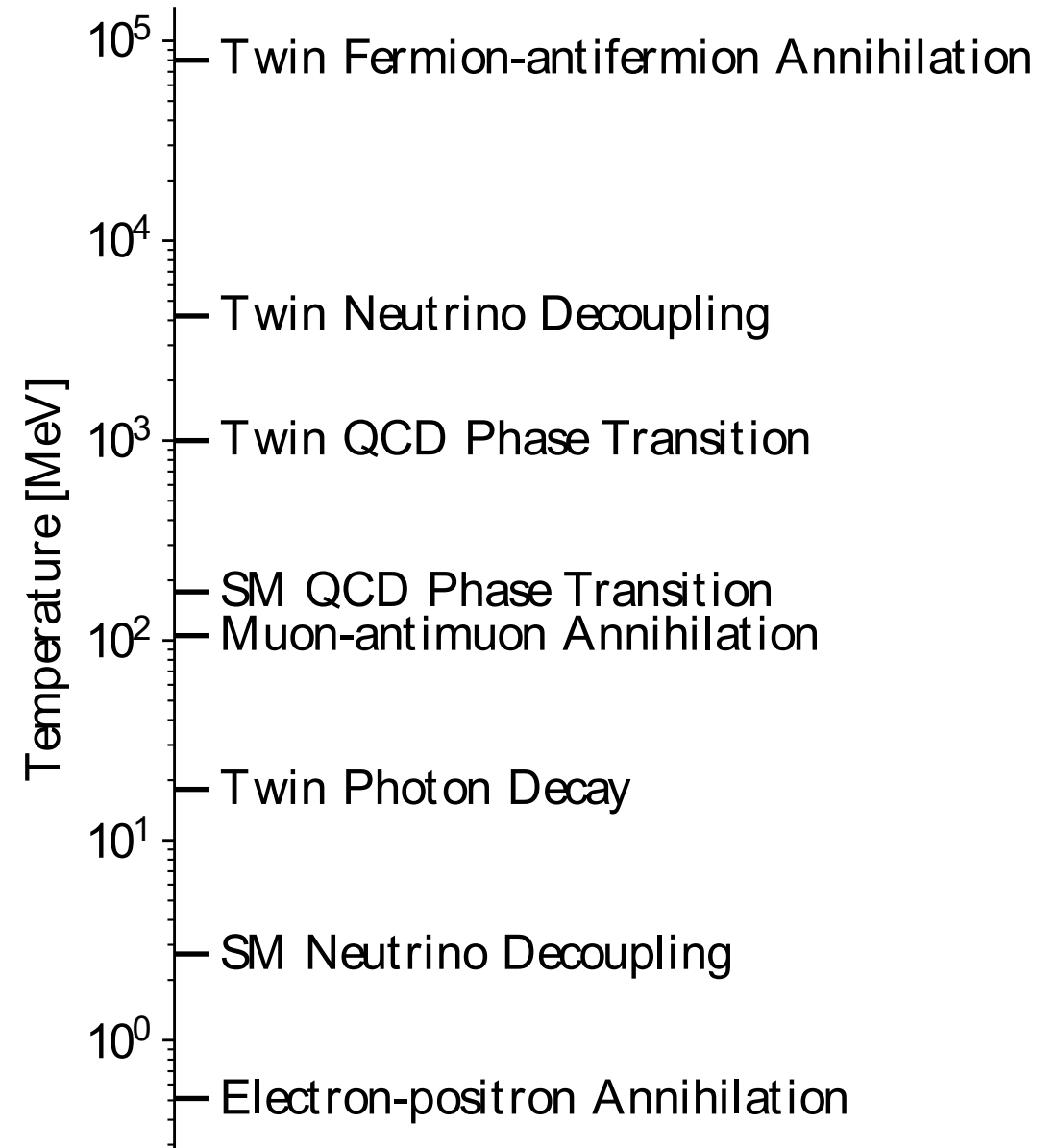
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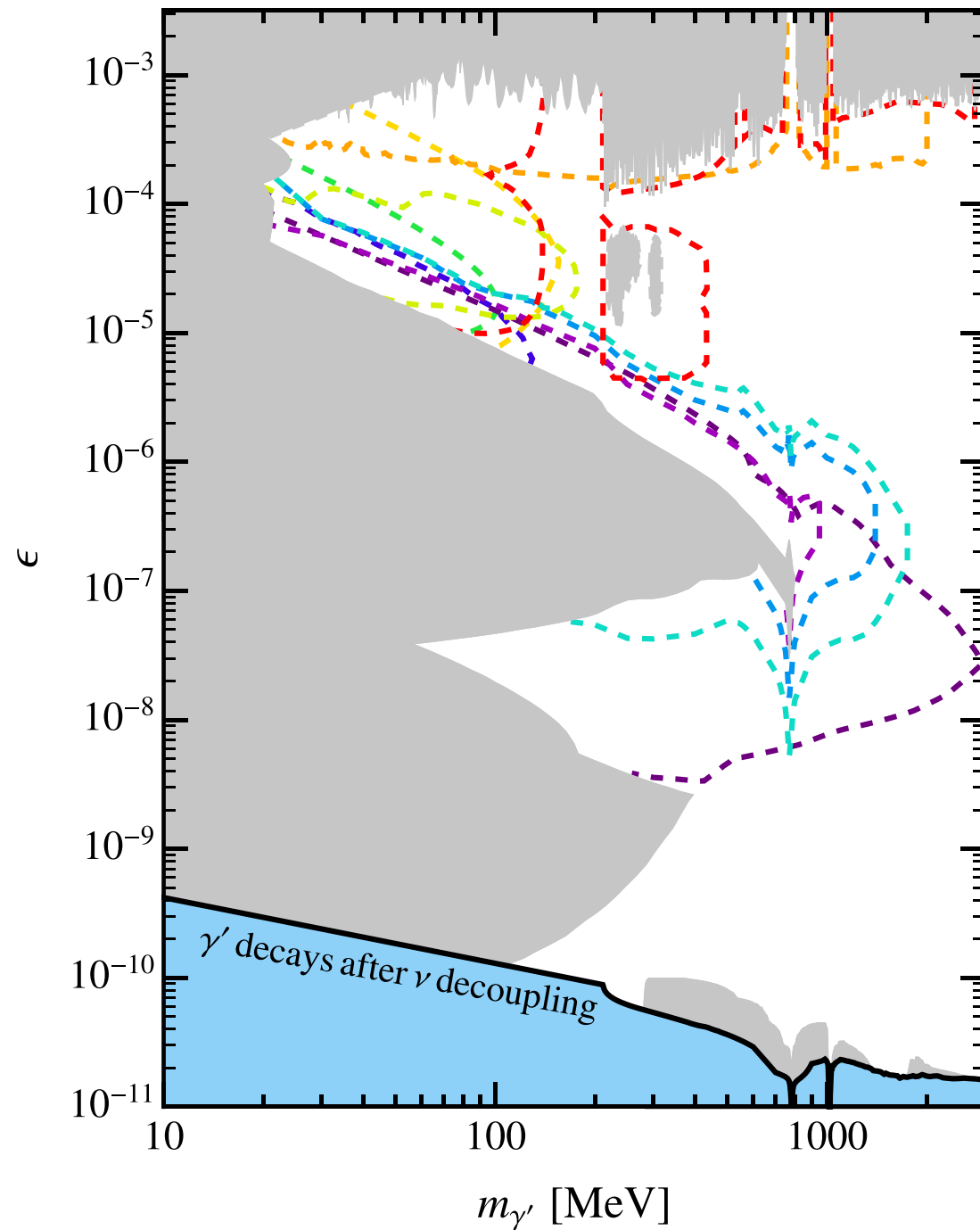
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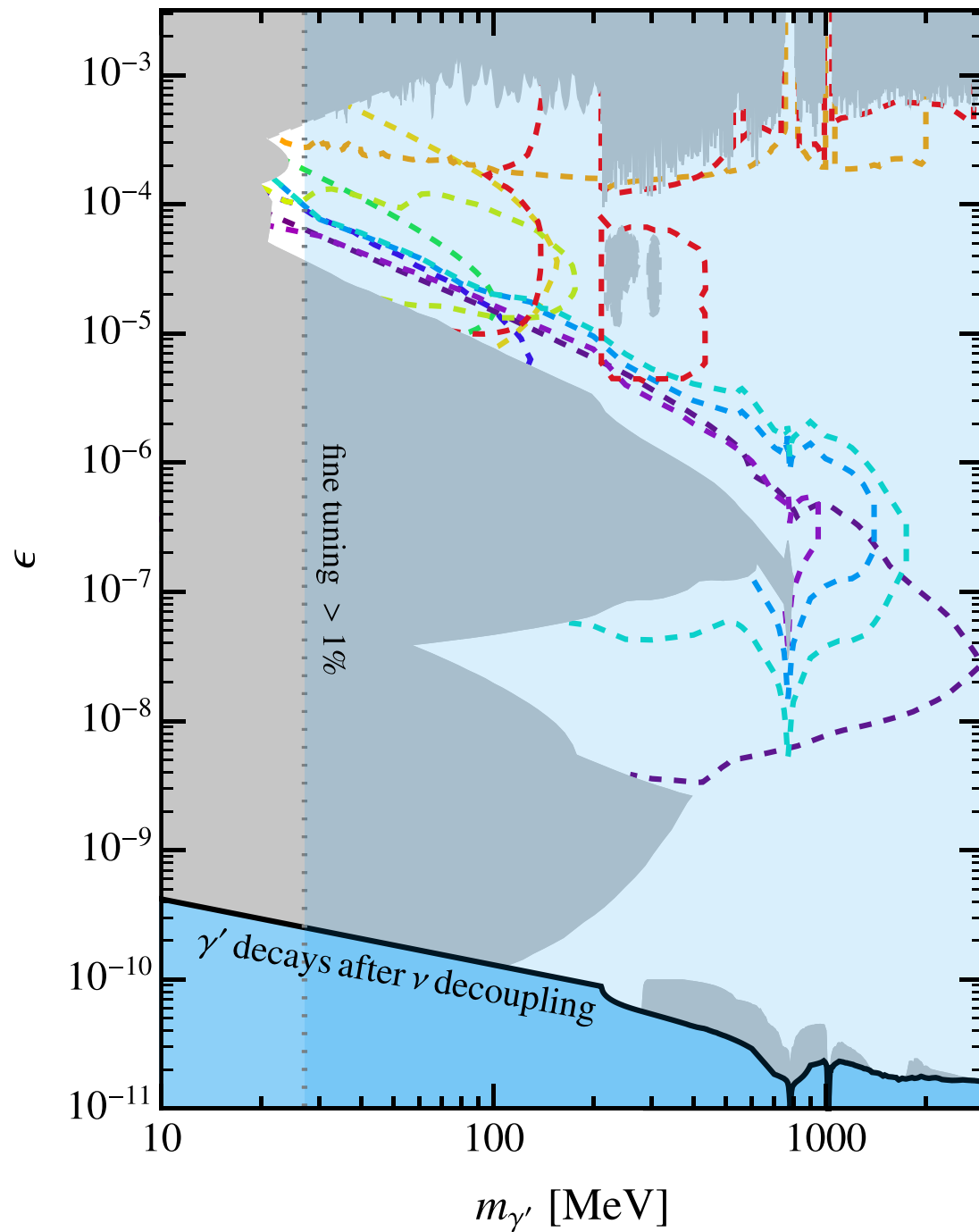
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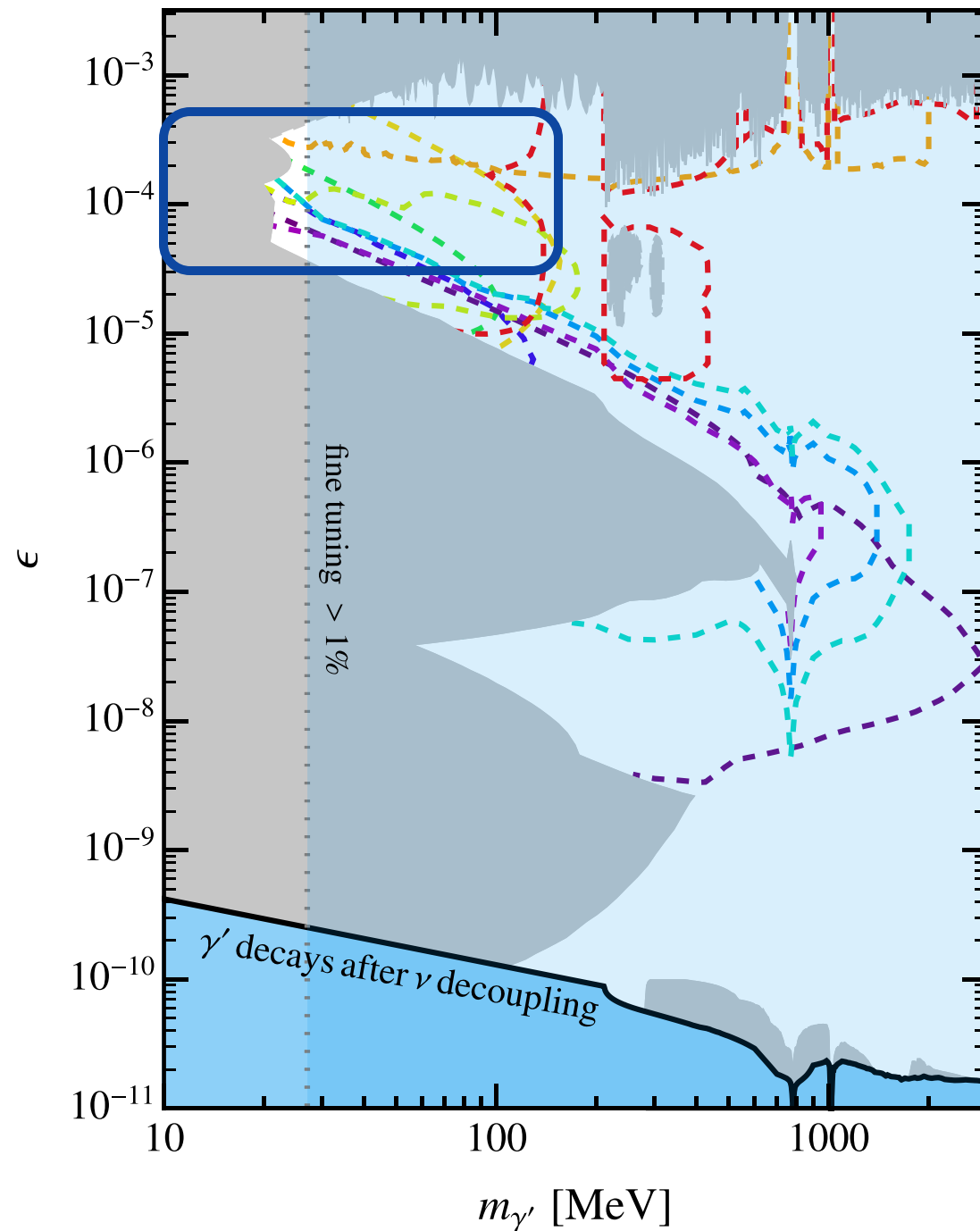
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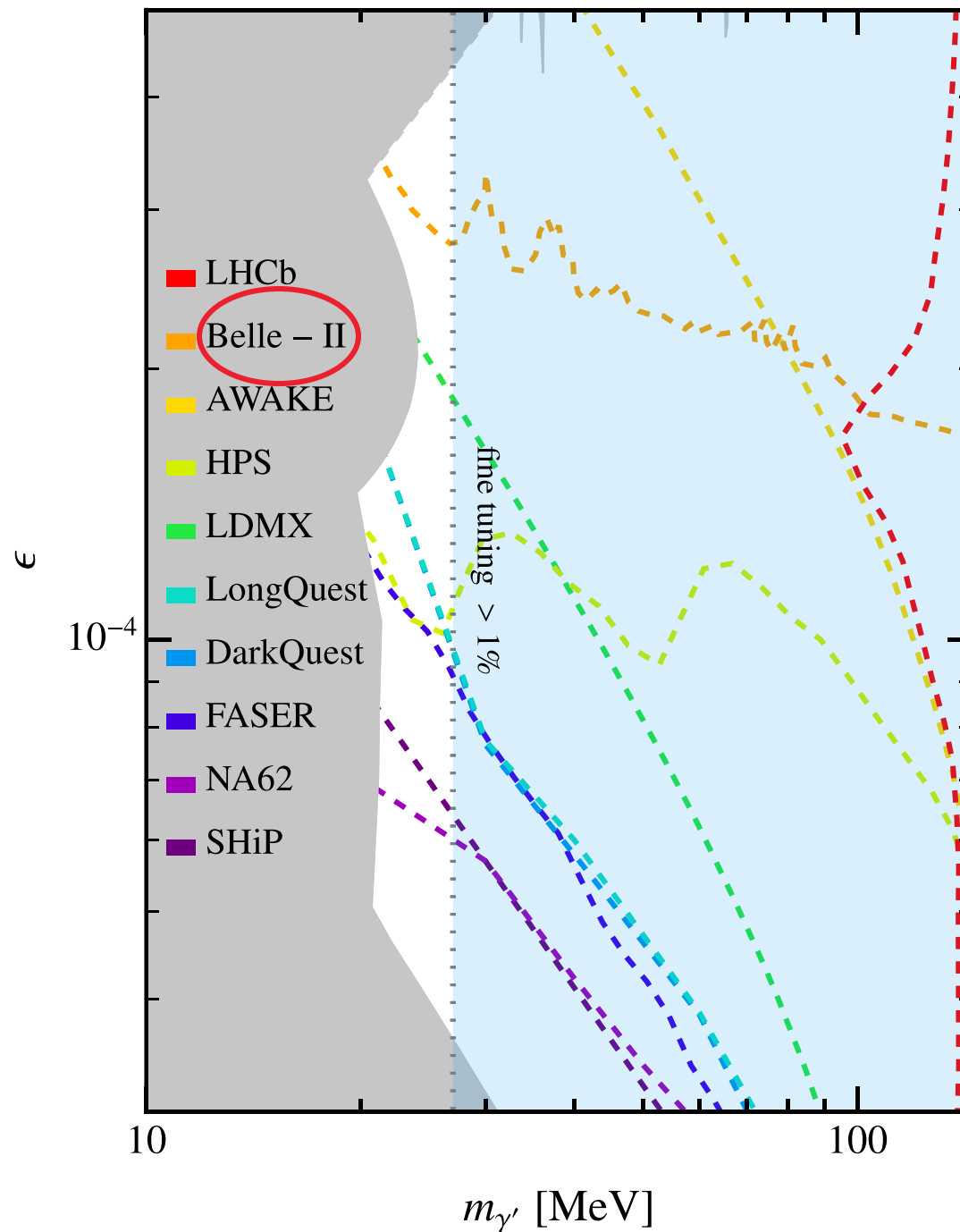
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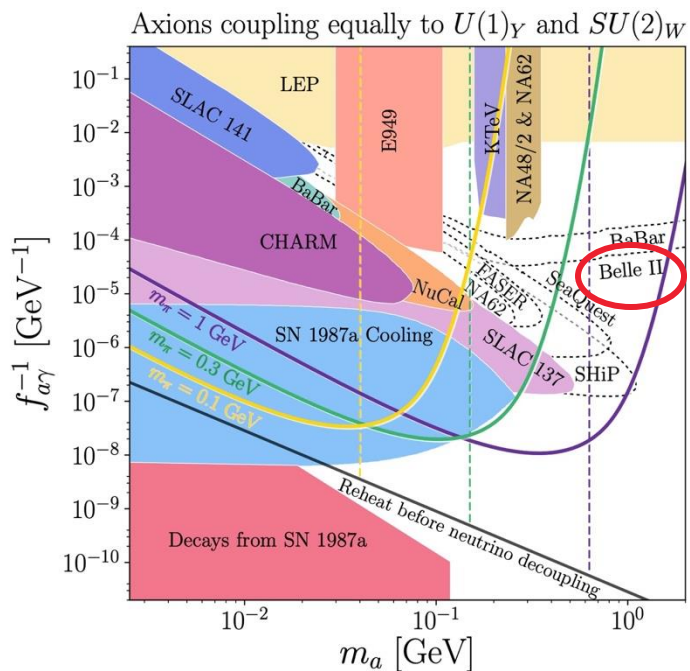
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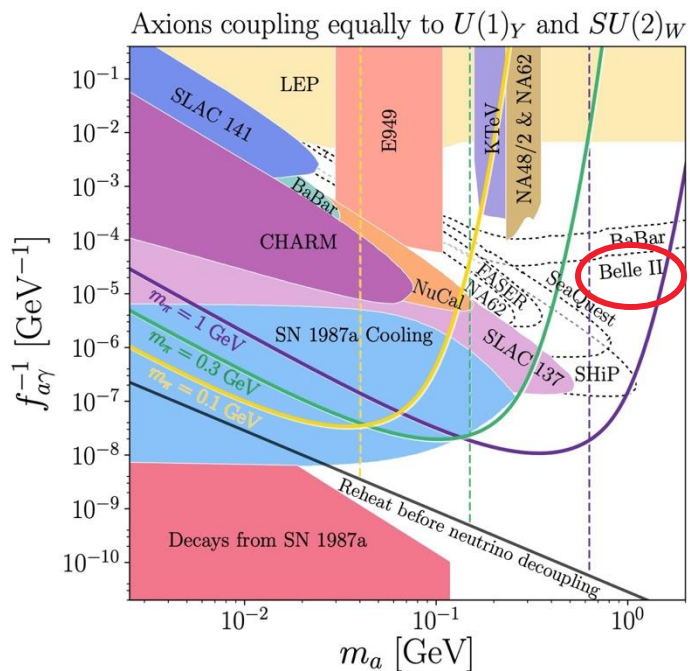
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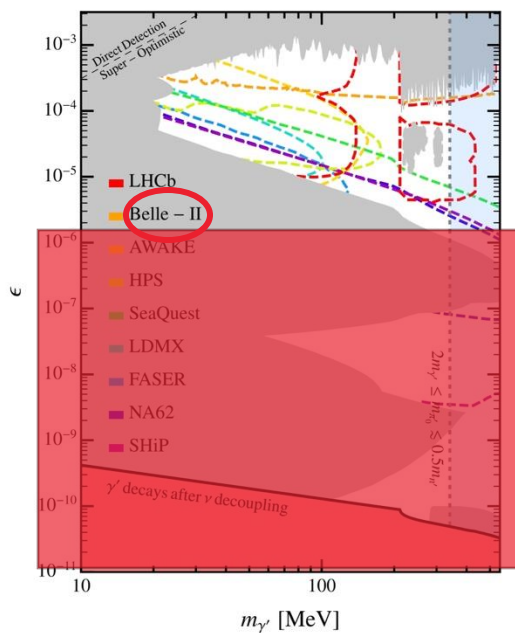
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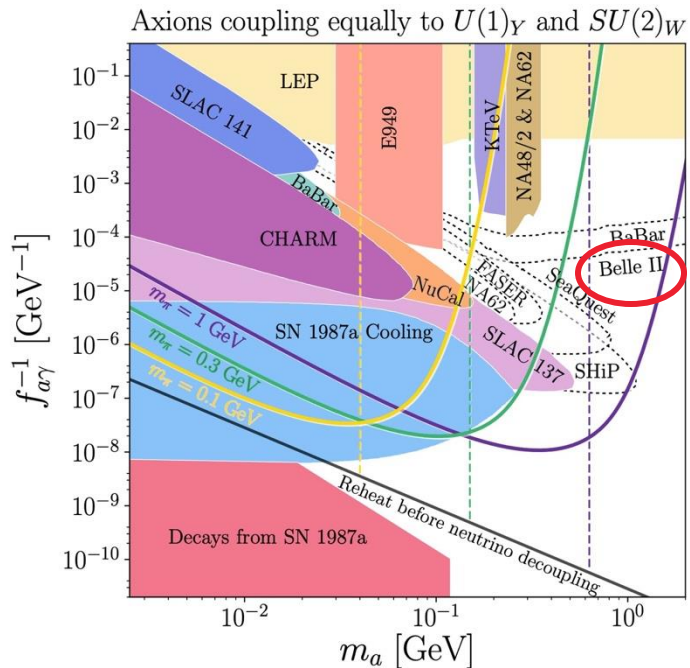
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