

# CHALLENGING BSM SEARCHES WITH TAUS

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based on work with I. Bigaran, P. Fox, Y. Gouttenoire,  
R. Harnik, J. Kopp, G. Krnjaic, T.Menzo, 2412.nnnnn

2024 Belle II Physics Week, KEK, Aug 16 2024

# GOAL

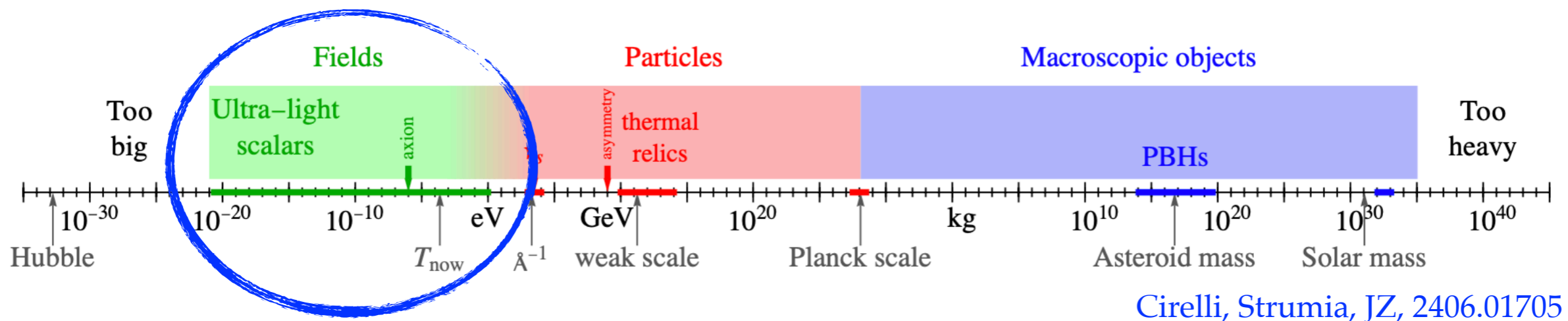
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see, e.g., review by Cirelli, Strumia, JZ, 2406.01705

- diagonal couplings to light DM  $\bar{\psi}_i \psi_i \phi_{\text{DM}}$   
 $\Rightarrow$  time dependent signals
- time-varying constants of nature,  $\alpha_{\text{em}}$
- time-varying electron, proton, neutron masses
- any implications for flavor violating transitions?

# GOAL

- yes! DM could be discovered in FCNCs
  - for instance, in  $\tau \rightarrow \mu + \text{inv}$
  - here, interested in very light DM
    - time dependent  $\tau \rightarrow \mu + \text{inv}$  rate



# OUTLINE

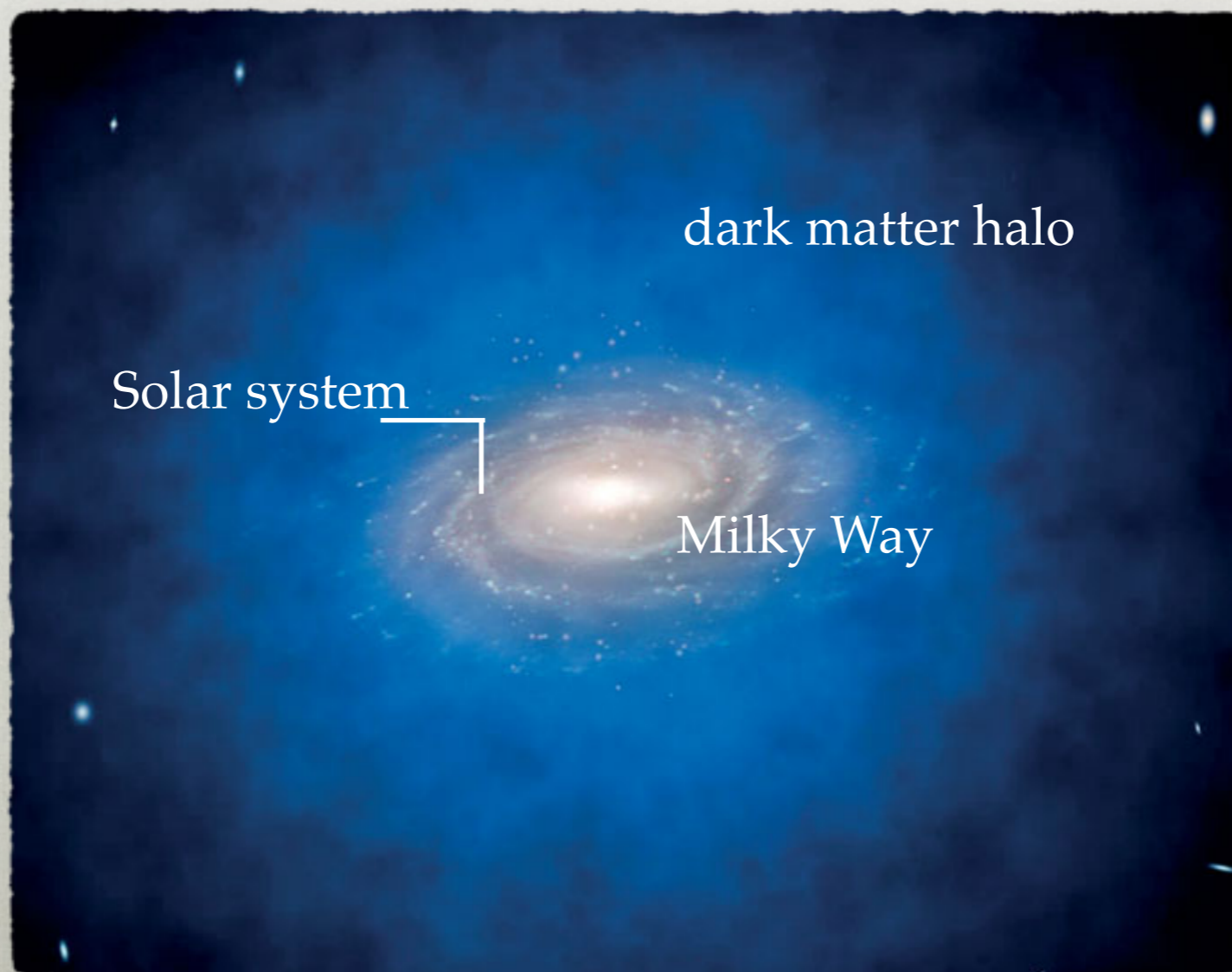
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- light DM and coherent background field oscillations
- light scalars
  - single axion-like particle
  - non-Abelian pNGBs
- time dependent FCNCs

# COHERENT OSCILLATIONS

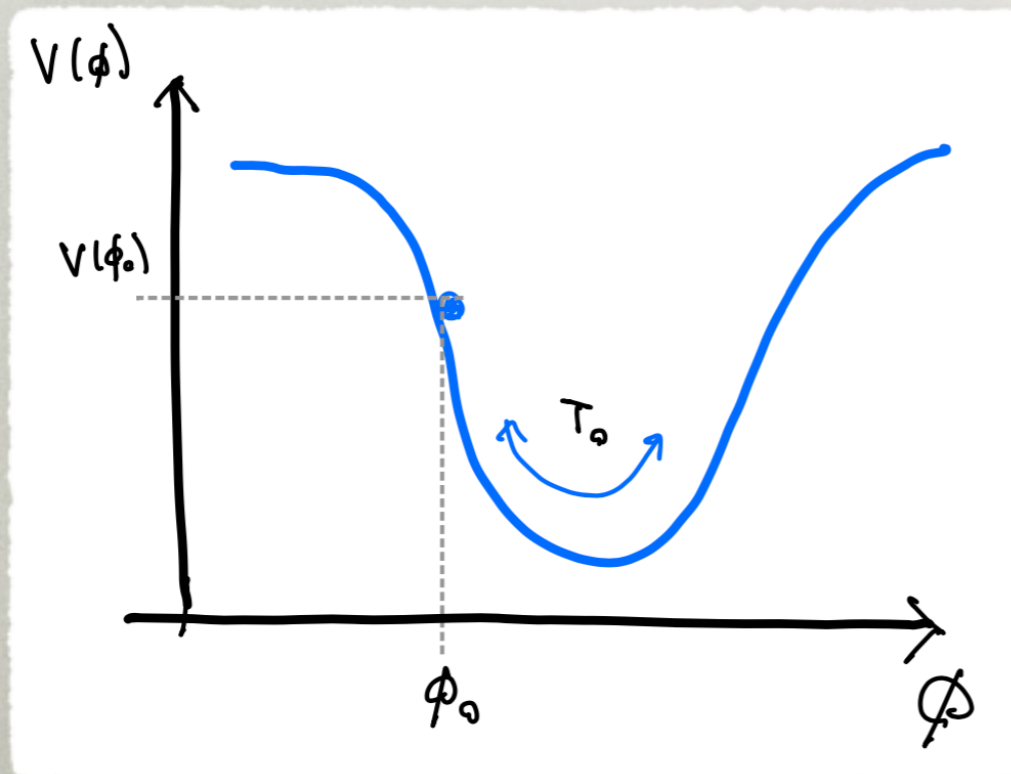
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- we are immersed in DM halo
  - non-relativistic DM particles  $v \sim \mathcal{O}(10^{-3})$



# COHERENT OSCILLATIONS

- local DM density  $\rho_\phi = 0.4 \text{ GeV/cm}^3 \approx 3 \times 10^{-42} \text{ GeV}^4$
- bosonic DM of mass  $m_\phi \lesssim 30 \text{ eV}$
- highly degenerate: many DM particles per de Broglie volume,  $n_{\text{DM}}(m_\phi v)^3 \gg 1$
- well approximated by oscillating wave  $\phi_{\text{cl}}(t)$



**Amplitude:**

$$\rho_\phi = V(\phi_0),$$

**Oscillation period:**

$$T_0 = 2\sqrt{2} \int_0^{\phi_0} \frac{d\phi}{\sqrt{V(\phi_0) - V(\phi)}}.$$

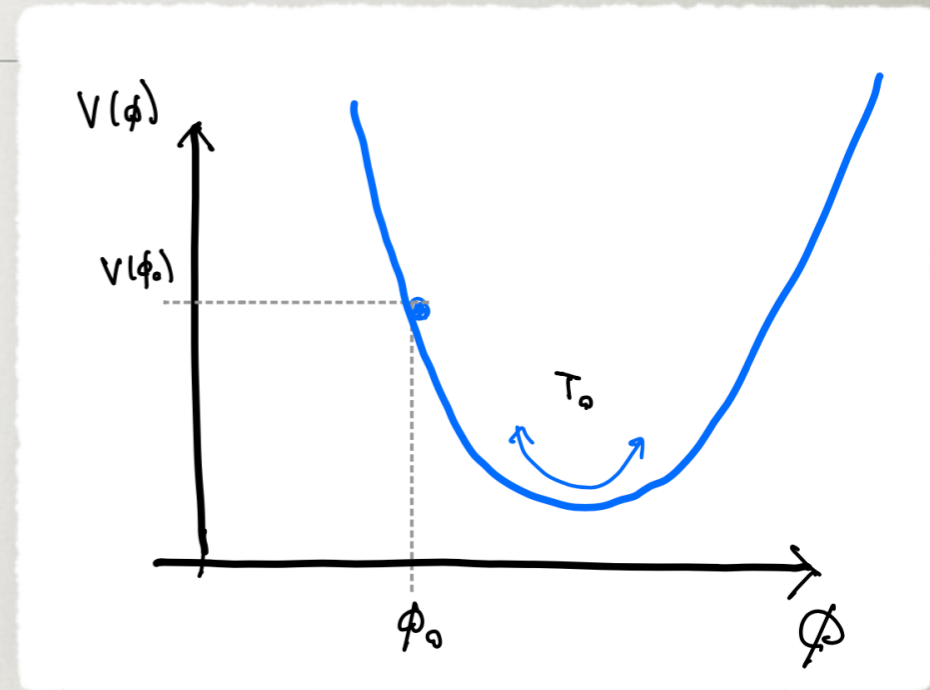
# COHERENT OSCILLATIONS

- example: quadratic potential

$$V_{\text{quad}}(\phi) = \frac{1}{2}m_\phi^2\phi^2.$$

- harmonic oscillations

$$\phi_{\text{cl}}(t) = \phi_0 \cos(m_\phi t + \delta) \quad , \quad \phi_0 = \frac{\sqrt{2\rho_\phi}}{m_\phi} \quad ,$$



- note: oscillation amplitude larger for lighter DM

$$\phi_0 \simeq 2.5 \text{ TeV} \left( \frac{10^{-15} \text{ eV}}{m_\phi} \right),$$

- oscillation frequency given by the DM mass  $T_0 = 2\pi/m_\phi$

$$T_0 \simeq 4.1 \text{ ns} \left( \frac{1 \mu\text{eV}}{m_\phi} \right) \simeq 4.1 \text{ s} \left( \frac{10^{-15} \text{ eV}}{m_\phi} \right) \simeq 16 \text{ month} \left( \frac{10^{-22} \text{ eV}}{m_\phi} \right).$$

- couplings  $\bar{\psi}_i \psi_j \phi^n \rightarrow n \bar{\psi}_i \psi_j \phi_{\text{cl}}^{n-1} \phi \Rightarrow$  time dependent  $\psi_j \rightarrow \psi_i \phi$  decays

- $\Gamma(\psi_j \rightarrow \psi_i \phi) \propto \cos^{2(n-1)}(m_\phi t)$

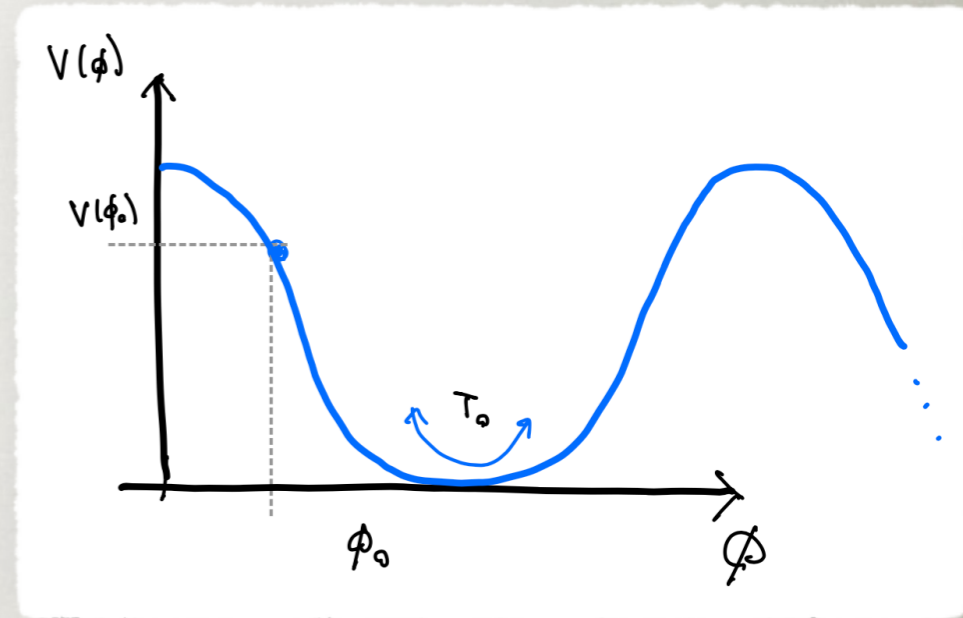
note:  $\bar{\psi}_i \psi_j \phi \Rightarrow$  no time dependence  
need at least  $\bar{\psi}_i \psi_j \phi^2$

# COHERENT OSCILLATIONS

- example: axion-like potential

$$V_{\text{ALP}}(\phi) = m_\phi^2 f^2 \left( 1 - \cos \frac{\phi}{f} \right),$$

- note: for fixed  $f \Rightarrow$   
max. DM density is  $\rho_\phi^{\text{max}} = m_\phi^2 f^2$



- anharmonic oscillations

- oscillation period  $T_0 = \frac{2\pi}{m_\phi} \left( 1 + \frac{1}{16} \phi_0^2 + \dots \right)$




- amplitude of oscil.  $\phi_0 = f \cos^{-1} \left( 1 - \frac{\rho_\phi}{m_\phi^2 f^2} \right).$

- note:  $\phi_0$  at most  $\phi_0^{\text{max}} = \frac{\pi}{2} f.$



# LIGHT SCALAR?

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- why can  $\phi$  be light?
  - if a pNGB of a global spontaneously broken symmetry
    - $\Rightarrow$  shift symmetry  $\phi \rightarrow \phi + \delta$
    - $\Rightarrow$  mass  $m_\phi$  protected
- three limits
  - a single axion-like particle 
  - non-Abelian pNGBs 
  - general light scalar 

# SINGLE ALP

- pNGB from spont. broken global U(1)
- the most general interaction at low energies starts at dim-5

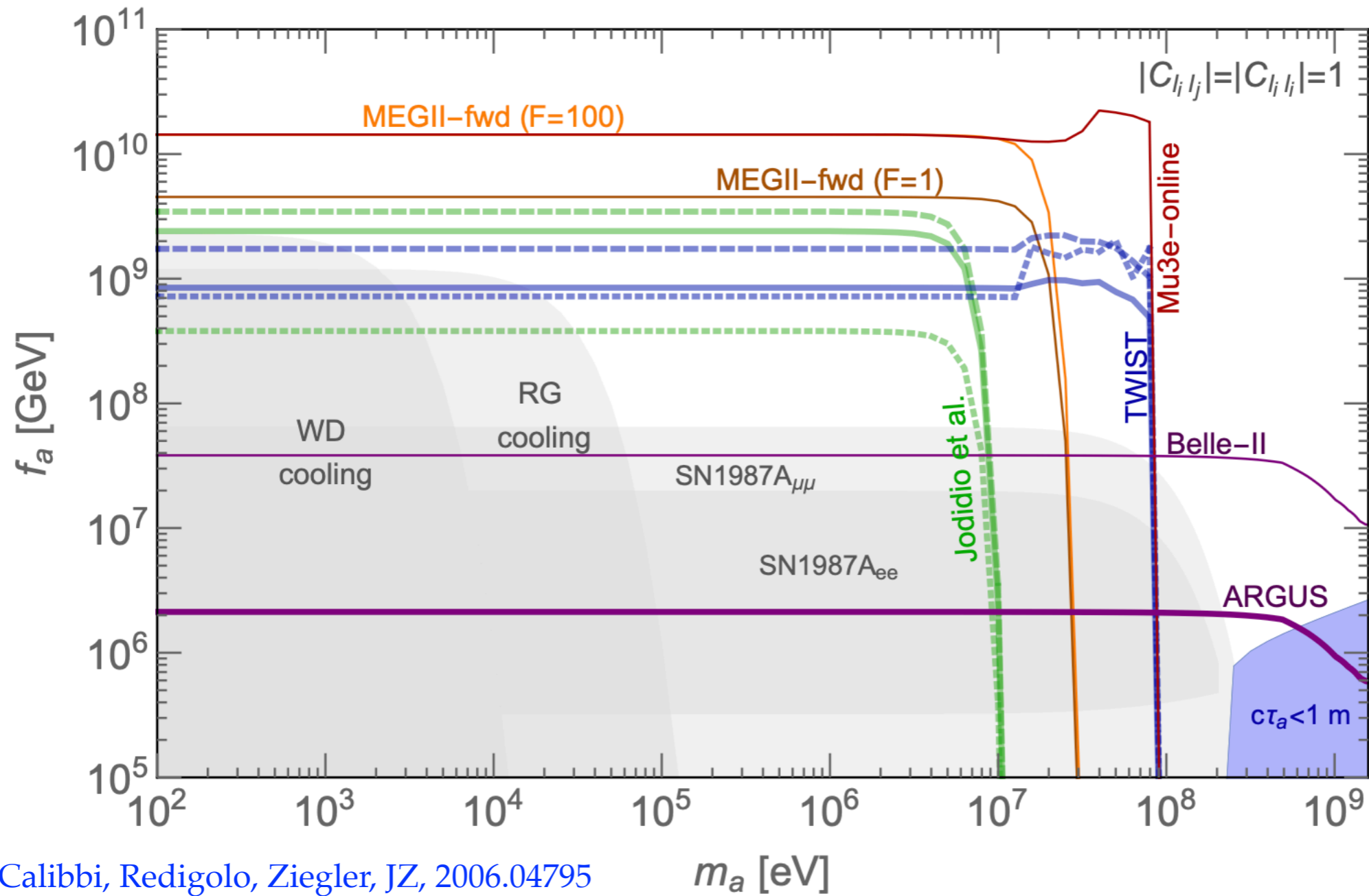
$$\mathcal{L}_{\text{int}} = \frac{\partial_\mu \phi}{2f} \bar{\psi}_i \gamma^\mu (C_{\psi_i \psi_j}^V + C_{\psi_i \psi_j}^A \gamma_5) \psi_j + c_g \frac{\phi}{f} \frac{\alpha_s}{8\pi} G_{\mu\nu}^a \tilde{G}^{a\mu\nu} + c_\gamma \frac{\phi}{f} \frac{\alpha_{\text{em}}}{8\pi} F_{\mu\nu} \tilde{F}^{\mu\nu}$$

- if  $m_\phi$  due to explicit breaking of shift symmetry also higher order terms

$$\mathcal{L}_{\text{int}} \supset \sum_n \frac{m_\phi}{f} \left[ \left( \frac{\phi}{f} \right)^n \bar{\psi}_i (C_{\psi_i \psi_j}^{S(n)} + C_{\psi_i \psi_j}^{P(n)} \gamma_5) \psi_j + c_g^{(n)} \left( \frac{\phi}{f} \right)^n \frac{\alpha_s}{8\pi} G_{\mu\nu}^a \tilde{G}^{a\mu\nu} + c_\gamma^{(n)} \left( \frac{\phi}{f} \right)^n \frac{\alpha_{\text{em}}}{8\pi} F_{\mu\nu} \tilde{F}^{\mu\nu} + \dots \right]$$

- for these to be comparable to dim-5 terms requires  $\phi_0 \gg f$ 
  - $\Leftarrow$  hard to arrange in any realistic model
  - $\Rightarrow$  time dependence of  $\psi_i \rightarrow \psi_j \phi$  decays highly suppressed

# SINGLE ALP

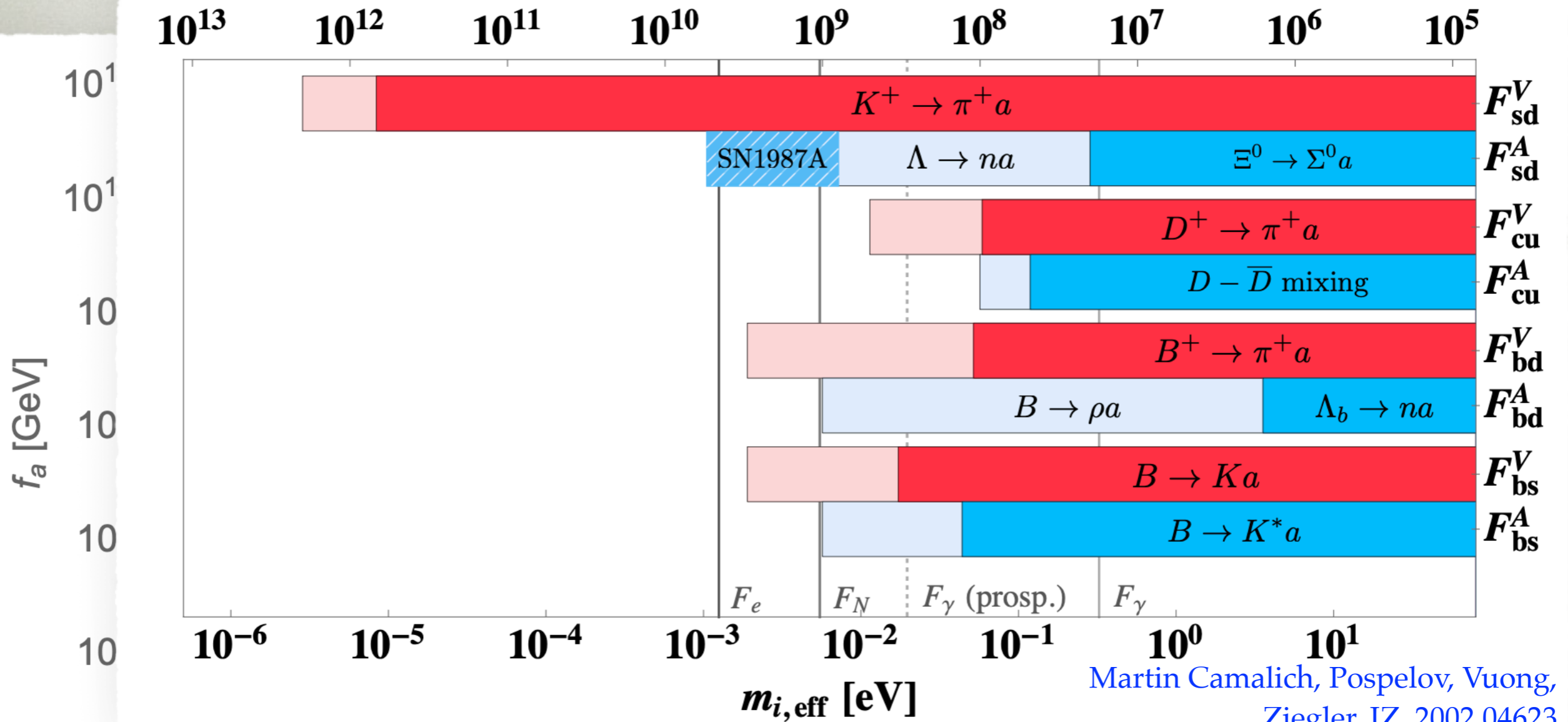


Calibbi, Redigolo, Ziegler, JZ, 2006.04795

$m_a$  [eV]

# SINGLE ALP

$F_i$  [GeV]



Calibbi, Redigolo, Ziegler, JZ, 2006.04795

$m_a$  [eV]

# NON-ABELIAN PNOGB

- consider  $G \rightarrow H$  breaking
- where pNOGBs in  $G/H$  coset  $U(\phi)$  have non-linear interactions
  - low energy interaction start as

$$\mathcal{L}_{\text{int}} \supset \text{Tr} (U^\dagger i\partial_\mu U) \bar{\psi}_i \gamma^\mu (\tilde{C}_{\psi_i \psi_j}^V + \tilde{C}_{\psi_i \psi_j}^A \gamma_5) \psi_j + \text{h.c.}$$

- if no U(1) factors, interactions start at  $\mathcal{O}(\phi^2)$

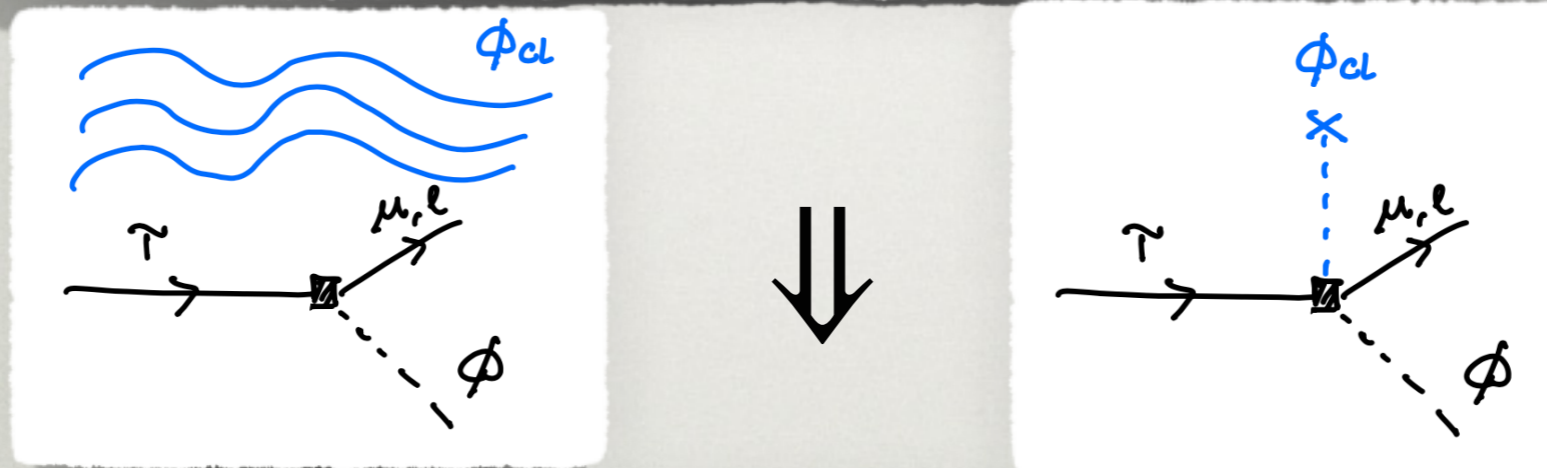
$$\mathcal{L}_{\text{int}} \supset \sum_a \frac{\phi_a}{f} \frac{i\partial_\mu \phi_a}{f} \bar{\psi}_i \gamma^\mu (C_{\psi_i \psi_j}^V + C_{\psi_i \psi_j}^A \gamma_5) \psi_j + \text{h.c.},$$

example from QCD+QED:  $\pi^+ \partial_\mu \pi^- J_{\text{em}}^\mu$

# NON-ABELIAN PNGB

- in the light DM background

$$\mathcal{L}_{\text{int}} \supset \sum_a \frac{\phi_a}{f} \frac{i\partial_\mu \phi_a}{f} \bar{\psi}_i \gamma^\mu (C_{\psi_i \psi_j}^V + C_{\psi_i \psi_j}^A \gamma_5) \psi_j + \text{h.c.},$$

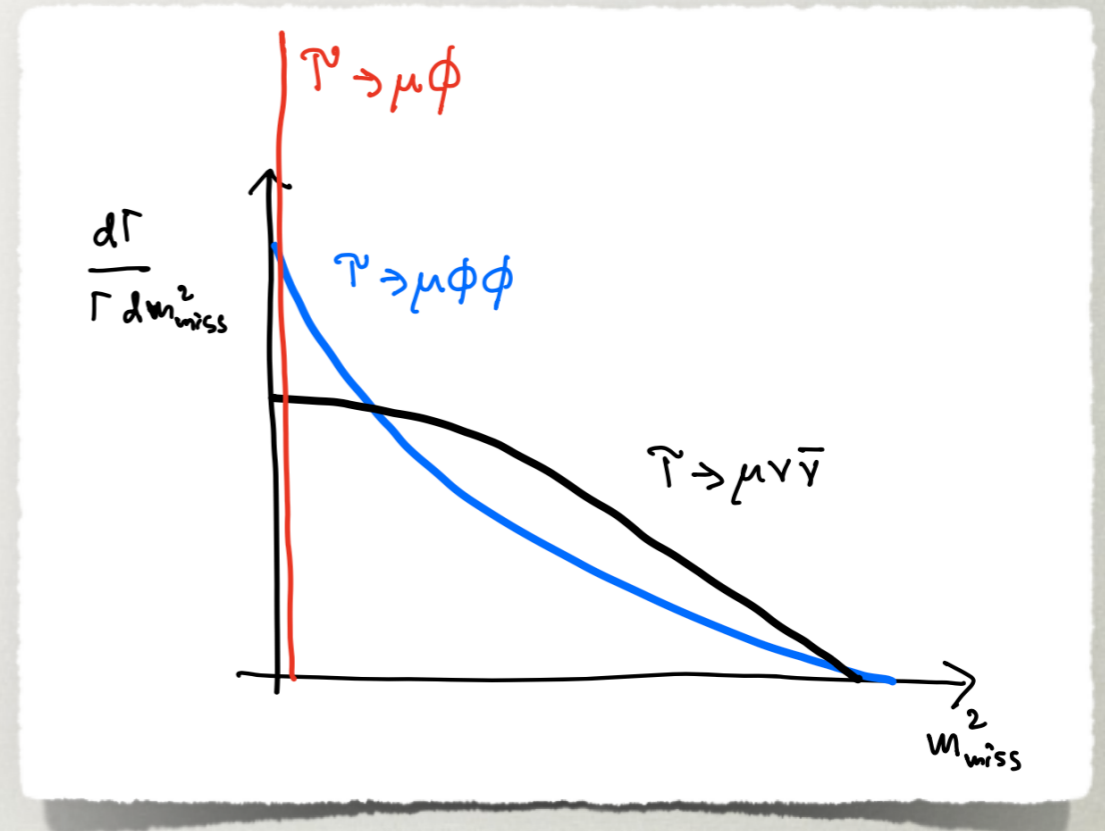


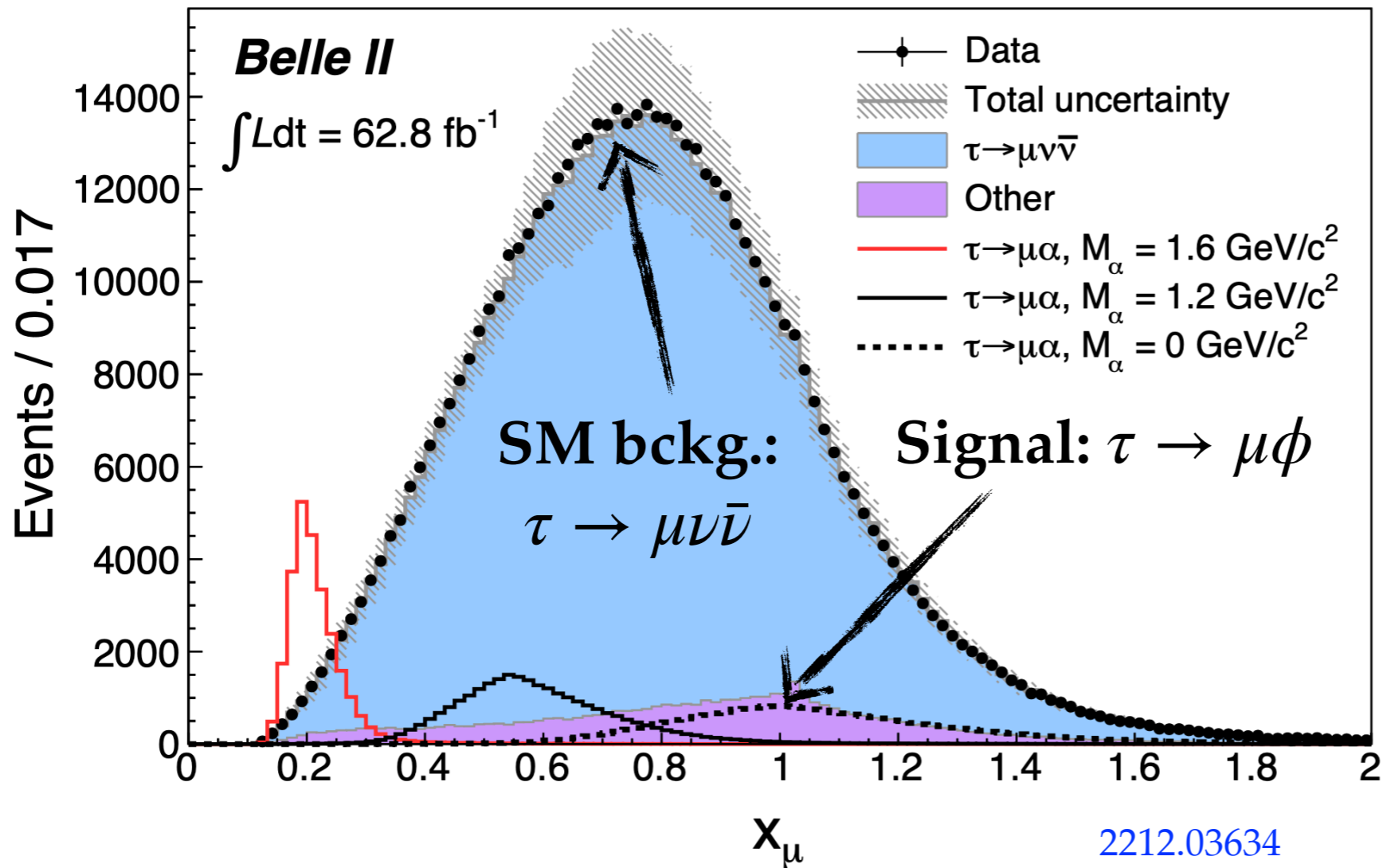
$$\mathcal{L}_{\text{int}} \supset \frac{\phi_{a,\text{cl}}}{f} \frac{i\partial_\mu \phi_a}{2f} \bar{\psi}_i \gamma^\mu (C_{\psi_i \psi_j}^V + C_{\psi_i \psi_j}^A \gamma_5) \psi_j + \text{h.c.}$$

- induces time dependent FCNCs
- example:  $Br(\tau \rightarrow \mu\phi) \propto \cos^2(m_\phi t)$

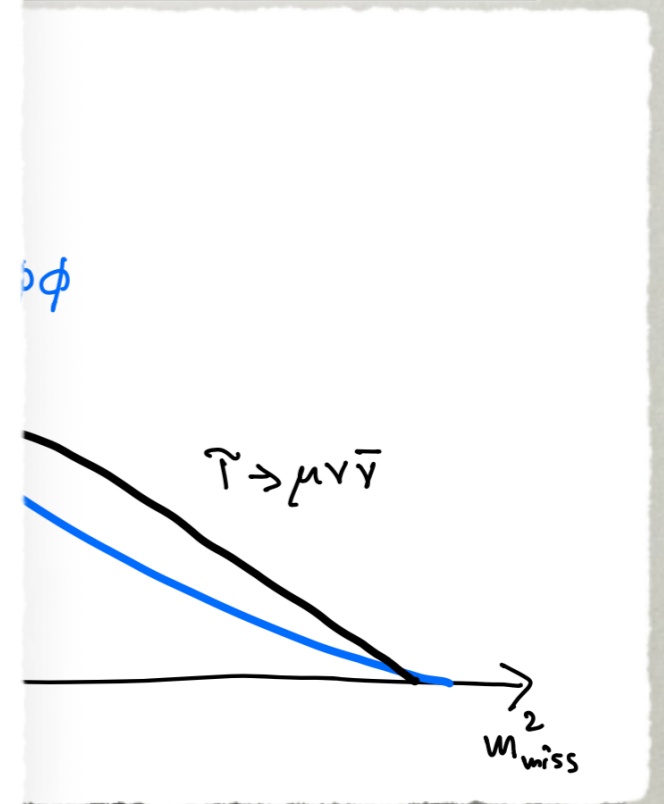
# TIME DEPENDENT $\tau \rightarrow \mu\phi$

- interaction:  $\phi\partial_\alpha\phi\bar{\tau}\gamma^\alpha\mu$ 
  - induces  $\tau \rightarrow \mu\phi\phi$ 
    - three body decay, large background from  $\tau \rightarrow \mu\nu\bar{\nu}$
    - very poor bound on  $f$
  - DM background induces time dependent  $\tau \rightarrow \mu\phi$ 
    - two body decay: mono-energetic  $\mu$  in tau rest- frame
- tau decays additional complication
  - $e^+e^- \rightarrow \tau^+\tau^-$ , at least one neutrino on tag side
  - not possible to reconstruct tau rest frame  $\Rightarrow$  use pseudo rest-frame
  - time dependence of the signal helps
- same for  $\tau \rightarrow e\phi$





$\mu \phi$



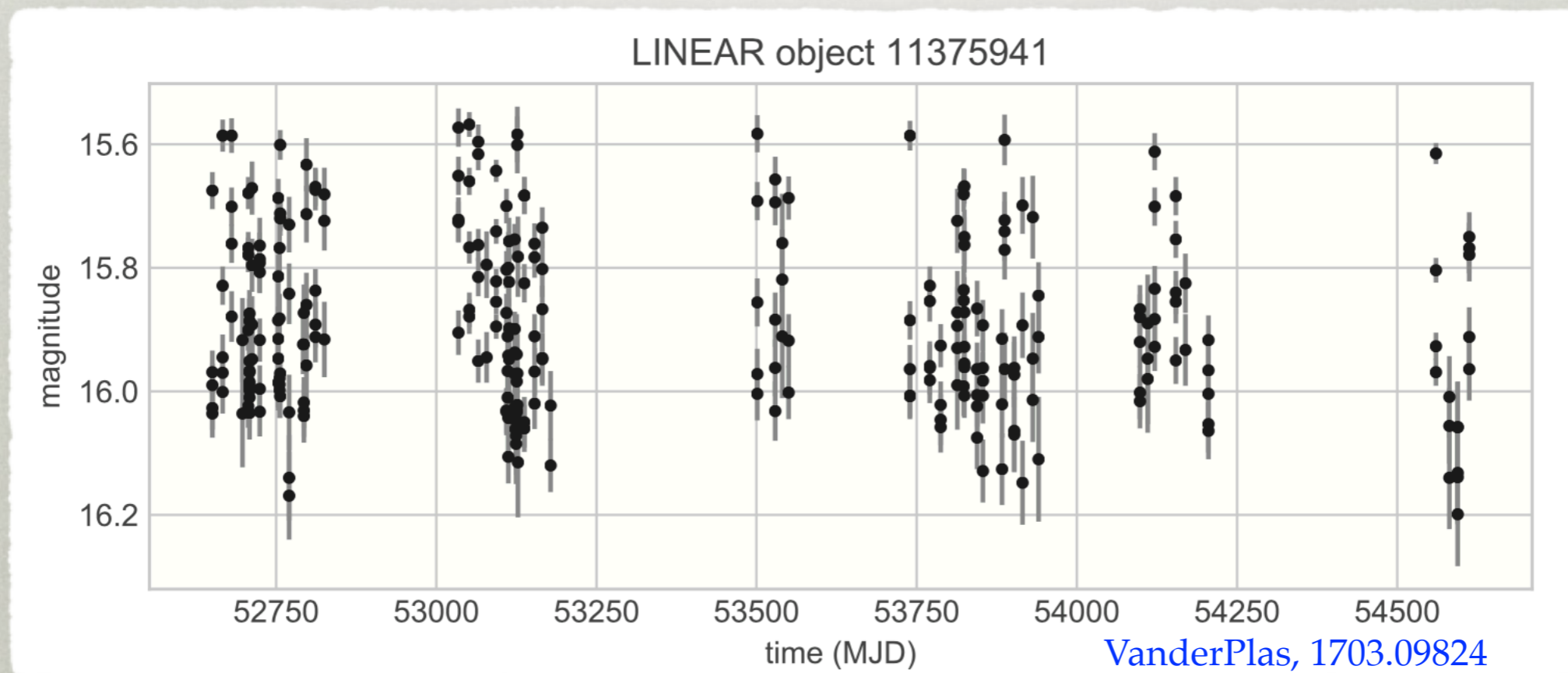
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# HOW TO SEARCH FOR PERIODIC SIGNALS

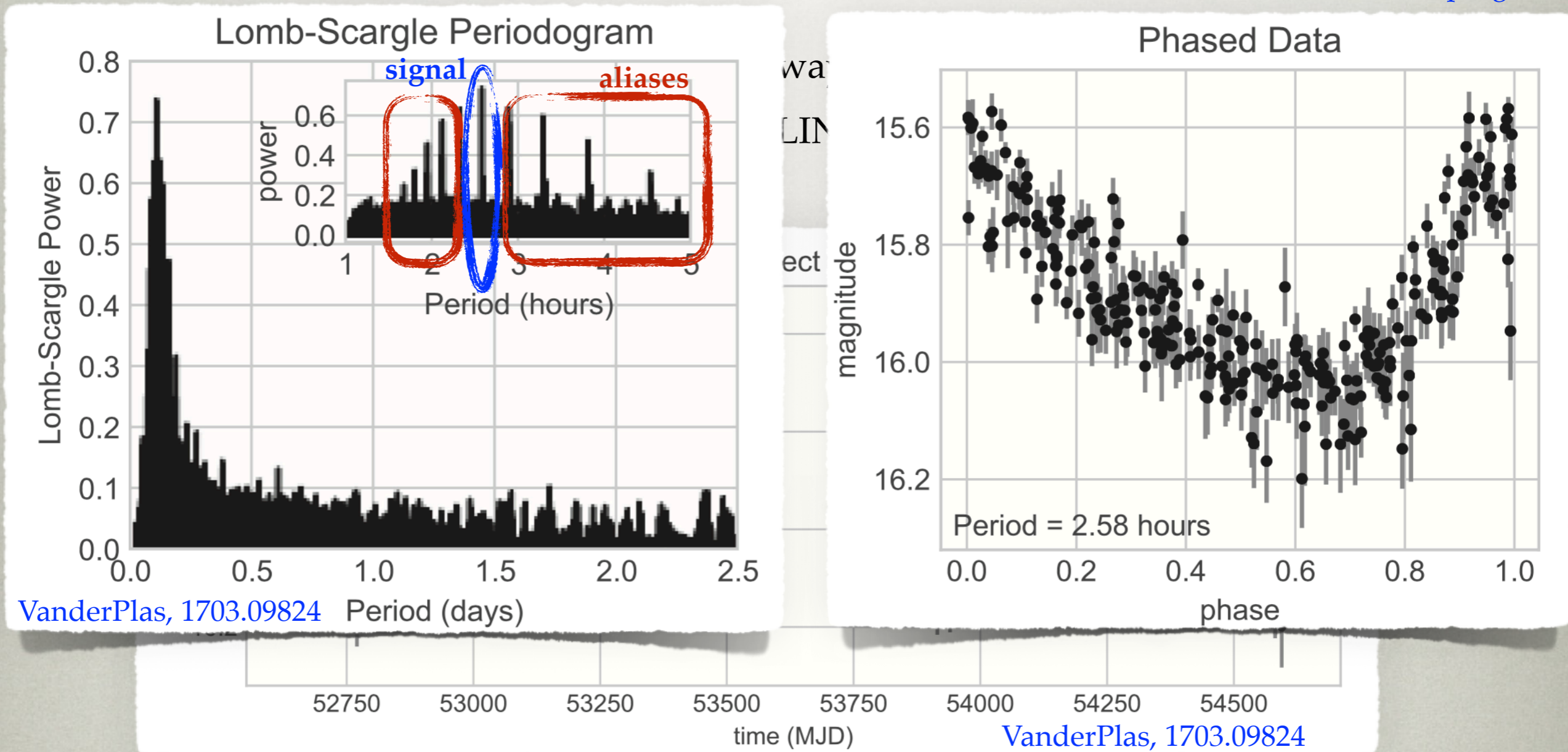
w/ Ilten et al, work in progress

- Lomb-Scargle periodogram an efficient way of searching for periodic signals
  - example: observed light curve from LINEAR object ID 11375941

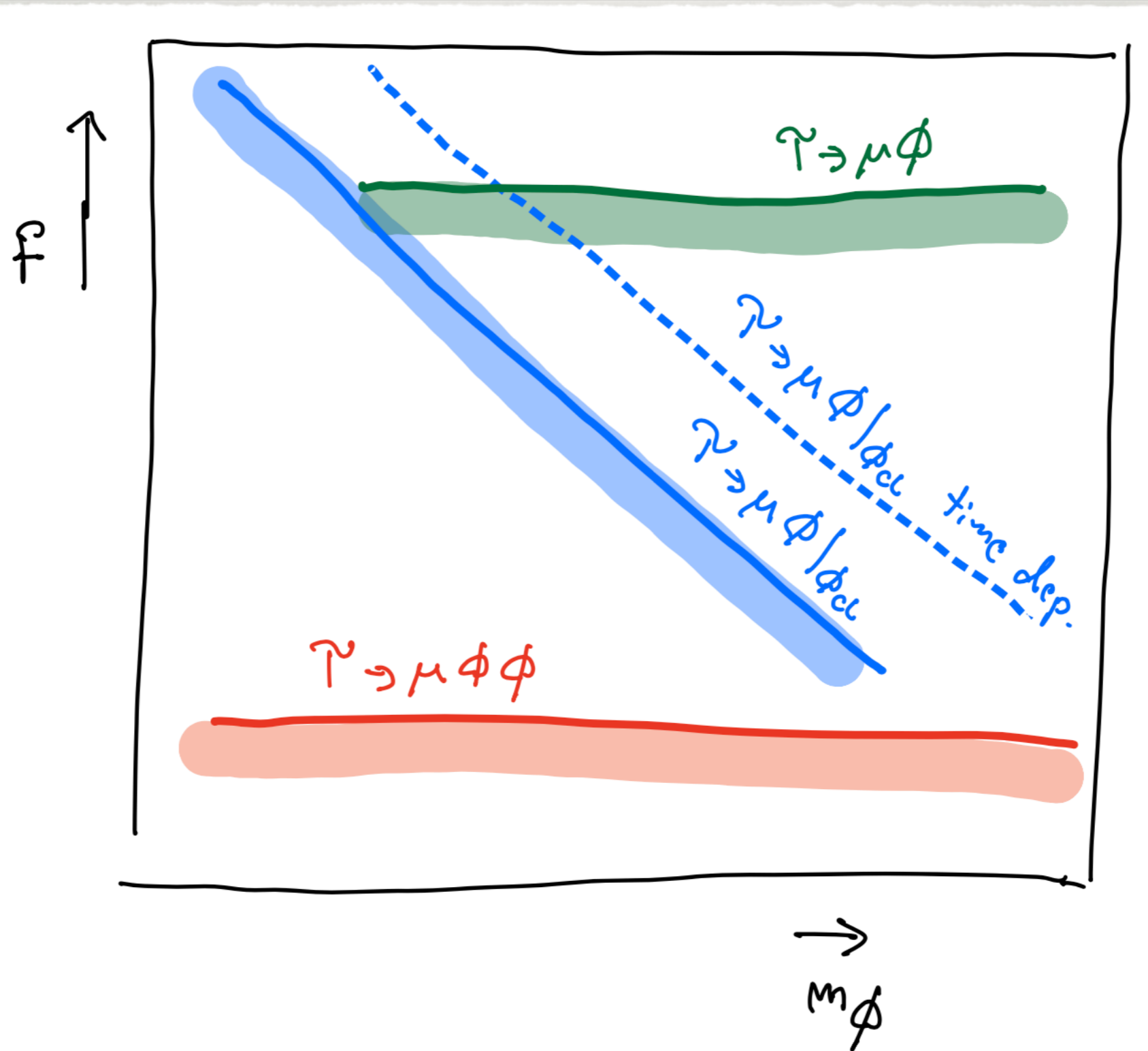


# HOW TO SEARCH FOR PERIODIC SIGNALS

w/ Ilten et al, work in progress



# TIME DEPENDENT $\tau \rightarrow \mu\phi$



# TIME DEPENDENT SIGNALS

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- any  $\psi_i \rightarrow \psi_j \phi$  FCNC process now time dependent
  - $\mu \rightarrow e \phi$
  - $s \rightarrow d \phi: K^+ \rightarrow \pi^+ \phi, K_{S,L} \rightarrow \pi^0 \phi, K^+ \rightarrow \pi^+ \pi^0 \phi, \Lambda \rightarrow n \phi, \dots$
  - $c \rightarrow u \phi: D \rightarrow \pi \phi, \rho \phi, D_s \rightarrow K^{(*)} \phi, \Lambda_c \rightarrow p \phi, \dots$
  - $b \rightarrow s \phi: B \rightarrow K^{(*)} \phi, \Lambda_b \rightarrow \Lambda \phi, \dots$
  - $b \rightarrow d \phi: B \rightarrow \pi \phi, B \rightarrow \phi \phi, \Lambda \rightarrow n \phi$
- the coherence of the signal is  $m_\phi / m_\phi v^2 \sim v^{-2} \sim 10^6$  oscillations
  - for month-scale oscillations only need time-stamps with precision of seconds
- in principle also contributions to FCNCs without missing energy
  - meson mixing, hadronic and leptonic meson decays

# CONCLUSIONS

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- we are immersed in DM background
- if DM light  $\Rightarrow$  coherently oscillating field
  - can be searched for through time dependent FCNC transitions
  - example:  $\tau \rightarrow \mu\phi$  (+many more)

# BACKUP SLIDES

# LOMB-SCARGLE PERIODOGRAM

- data with time-stamps,  $y(t_i)$ ,  $i = 1, \dots, N$
- Lomb-Scargle power for frequency  $f = \omega/2\pi$

$$P(f) = \frac{1}{2\sigma^2} \left( \frac{[\sum_{i=1}^N W_i (y(t_i) - \bar{y}) \cos \omega(t_i - \tau)]^2}{\sum_{i=1}^N W_i \cos^2 \omega(t_i - \tau)} + \frac{[\sum_{i=1}^N W_i (y(t_i) - \bar{y}) \sin \omega(t_i - \tau)]^2}{\sum_{i=1}^N W_i \sin^2 \omega(t_i - \tau)} \right)$$

- $\bar{y}$  is weighted average,  $\sigma$  weighted variance of data

- weights  $W_i$  are

$$W_i = \frac{1/\sigma_i^2}{\langle 1/\sigma_i^2 \rangle}$$

- phase factor  $\tau$

$$\tan(2\omega\tau) = \frac{\sum_{i=1}^N W_i \sin 2\omega t_i}{\sum_{i=1}^N W_i \cos 2\omega t_i}$$