Dark matter searches in Belle II

Belle II Summer School

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Gaps in the Standard Model

- Several open questions point to physics that is beyond the Standard Model
- Dark sectors often are proposed to address multiple open questions

- Dark Matter
- Baryon Asymmetry of the Universe
- Muon g-2 Anomaly
- Neutrino mass
- Strong CP problem
- Dark Energy
- Gravitation



Dark Matter

- Astronomical and cosmological observations show 84% of universe's matter content is Dark Matter
- What we know:
 - It's stable, interacts Gravitationally with the known particles and it is very abundant
- Cannot be explained by any of the particles in the Standard Model









Weakly Interacting Massive Particles (WIMPS)

- Stable particle with GeV-TeV mass
- Interaction cross section with Standard Model near weak scale
- Dark Matter candidate



Muon Anomalous Magnetic Moment

$$\vec{M} = g_{\mu} \frac{e}{2m_{\mu}} \vec{S}$$

$$a_{\mu} \equiv \frac{g_{\mu} - 2}{2}$$



T. Aoyama, N. Asmussen, M. Benayoun et al. Physics Reports 887 (2020) 1–166 B. Abi et al. (Muon g-2 Collaboration) Phys. Rev. Lett. 126, 141801 (2021)

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https://news.fnal.gov/2021/04/first-results-from-fermilabs-muon-g-2-experiment-strengthen-evidence-of-new-physics/









Collection of particles with no direct coupling to Standard Model



Collection of particles with no direct coupling to Standard Model



Dark Photons

Dark Photon

- Massive vector boson arising from broken U(1) gauge symmetry in dark sector
- Couples to Standard Model via kinetic mixing with photon
 - $m_{A'}$ Dark Photon Mass
 - $\boldsymbol{\epsilon}\,$ Coupling strength to SM
 - $\Gamma_{A' \rightarrow \chi \chi}$ Branching fractions to decay to DM
- Detector signature classified as invisible or visible

If
$$m_{A'} > 2m_{\chi}$$

~100% of A^{\prime} decays are to dark matter

$$A' \to \chi \bar{\chi}$$

"Invisible" Detector Signature

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$$\varepsilon e A'_{\mu} J^{\mu}_{\rm EM} + \mathcal{L}_{A'\chi\chi}$$

$$\gamma A'$$

If
$$m_{A'} < 2m_{\chi}$$

$$A' \rightarrow \chi \bar{\chi}$$
 is forbidden.

Must decay to Standard Model!

$$A' \rightarrow e^+ e^-, \mu^+ \mu^-, \dots$$

"Visible" Detector Signature

M. Graham, C. Hearty, and M. Williams, Annu. Rev. Nucl. Sci. 71 (2021)

Invisible Dark Photons

- Initial State Radiation photon recoils against Dark Photon
- "Single Photon Search"
- Specialized Single Photon Trigger" is essential (was not present at Belle)



Single Photon Triggers at Belle II:

- At least one photon with E_{CMS} > 2 GeV
- One E_{CMS} > 1 GeV photon in barrel + no other energetic photons
- One $E_{CMS} > 0.5$ GeV photon in central barrel + no other energetic photons

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Invisible Dark Photons

• Belle II analysis (in progress) will explore parameter space that is consistent with universes dark matter abundance



Invisible Dark Photons





Visible Dark Photons

If $A' \to \chi \bar{\chi}$ is forbidden. Must decay to Standard Model!







ATOMKI ANOMOLY: A 17 MeV Dark Photon?





-excitation can also occur through via dark photon. ould give peak in e^+e^- angular distribution

> A. J. Krasznahorkay, et al. Phys. Rev. Lett. 116, 042501 A. J. Krasznahorkay, et al. Phys. Rev. C 104, 044003

ATOMKI ANOMOLY: A 17 MeV Dark Photon?



De-excitation can also occur through via dark photon. Would give peak in e^+e^- angular distribution

x100

X17 Search at Belle II (In Progress)

Main background sources :

•
$$e^+e^- \rightarrow e^+e^-\gamma$$

 $e^+e^- \rightarrow \gamma\gamma \ (\gamma \rightarrow e^+e^-)$
 γ



Bellell Simulation $e^+e^- \rightarrow \gamma A'[A]$

e⁺

 e^+



 e^{-}

Visible Dark Photons

- Belle II search bypasses nuclear physics uncertainties
- Current Belle II dataset can search remaining parameter space for protophobic X17 ($2\epsilon_u + \epsilon_d \lesssim 0.1\epsilon_u$)





E. Kou et al. Prog Theor Exp Phys (2019) Feng et al, Phys. Rev. Lett. 117, 071803 D. Alves et al. Eur. Phys. J. C (2023) 83:230

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Muonic Dark Force: $L_{\mu} - L_{\tau}$ Model

Dark Z' Boson and $L_{\mu} - L_{\tau}$ Model

- Massive dark vector boson (similar to Dark Photon)
- $L_{\mu} L_{\tau}$ model: Z' couples only to second and third generation leptons muon g-2, dark matter
- Invisible and visible signatures at Belle II depending on how $Z^\prime\,\text{decays}$



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 $BF(Z' \rightarrow \nu \bar{\nu}) \sim 33 - 100\%$ BF(Z' $\rightarrow \chi \bar{\chi}$) ~ 100% if kinematically allowed

Detected muons used to compute recoil mass that peaks for Z' signal

B. Shuve and I. Yavin, PRD 89, 113004 (2014)
W. Altmannshofer, S. Gori, M. Pospelov, and I. Yavin, PRL 113, 091801 (2014)
W. Altmannshofer, S. Gori, S. Profumo, and F. S. Queiroz, JHEP 12 (2016) 106
P. Harris, P. Schuster, and J. Zupan (2022), arXiv:2207.08990
N. Tran, and A. Whitbeck, PRD 107, 116026 (2023)

Search for Invisible Z'

• Backgrounds arise from:

 $e^+e^- \rightarrow \mu^+\mu^-(\gamma)$ where photon is not reconstructed $e^+e^- \rightarrow \tau^+\tau^-(\gamma)$ neutrinos escape detector $e^+e^- \rightarrow e^+e^-\mu^+\mu^-$ with e^+e^- not in acceptance

- No significant excess observed in 79.7 fb⁻¹
- Excluded part of Z' parameter space, which could explain muon g-2 tension



Belle II Collaboration Phys. Rev. Lett. 130, 231801 (2023)

Search for Invisible Z'

• Backgrounds arise from:



Belle II Collaboration Phys. Rev. Lett. 130, 231801 (2023)

Search for Visible Z'

 e^+

 e^{-}

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- Search in channel $e^+e^- \rightarrow \mu^+\mu^- X$, $X \rightarrow \mu^+\mu^-$
- Select events with four muons with total centre-ofmass energy consistent with \sqrt{s}
- Main background $e^+e^- \to \mu^+\mu^-\mu^+\mu^-$ has distinct kinematics from signal
- Neural network for background suppression



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Search for Visible Z^\prime

- No significant excess observed in 178 fb⁻¹
- Limits also set on Z' interpretation
- Set first limits on muonphilic scalar, which constrain explanation for muon g-2 anomaly



Dark Scalar/Dark Higgs Boson





Axion-Like Particles

Axion-Like Particles (ALPs)

- Axions originally proposed to solve strong Charge-Parity problem
- Axion-like particles (*a*) are a generalization of the axion
- Searches focus on ALP interactions with gauge bosons



Pseudo-scalar with mass m_a

$$\mathcal{L} \supset -\frac{g_{a\gamma\gamma}}{4} aF_{\mu\nu}\tilde{F}^{\mu\nu} - \frac{g_{a\gamma Z}}{4} aF_{\mu\nu}\tilde{Z}^{\mu\nu} - \frac{g_{aZZ}}{4} aZ_{\mu\nu}\tilde{Z}^{\mu\nu} - \frac{g_{aWW}}{4} aW_{\mu\nu}\tilde{W}^{\mu\nu}$$

 $e^+e^- \rightarrow e^+e^-a, a \rightarrow \gamma\gamma$

• ALPs could be directly produced at Belle II through variety of channels

 $B \to Ka, a \to \gamma \gamma$



M. Dolan, T. Ferber, C. Hearty, F. Kahlhoefer & K. Schmidt-Hoberg, JHEP 12, 094 (2017) E. Izaguirre, T. Lin, and B. Shuve Phys. Rev. Lett. 118, 111802 (2017)



 $e^+e^- \rightarrow \gamma a, a \rightarrow \gamma \gamma$



ALPs at Belle II

• Analysis focuses on ALP-Photon coupling





F. Abudinén *et al.* (Belle II Collaboration) Phys. Rev. Lett. 125, 161806 (2020) M. Nuccio. Search for Axion-Like Particles produced in e^+e^- collisions and photon energy resolution studies at Belle II. [PhD Dissertation] University of Hamburg (2021)

Search for $\tau\tau$ resonance

Search for $\tau\tau$ resonance

- $\tau\tau$ resonance in $e^+e^- \rightarrow \mu\mu\tau\tau$ arise in many dark sector models:
 - Spin-1 particle coupling only to the heavier lepton families
 - Higgs-like spin-0 particle that couples preferentially to charged leptons (leptophilic scalar)
 - Axion-like particles



B. Shuve and I. Yavin, PRD 89, 113004 (2014)
W. Altmannshofer et al, JHEP 12, 106 (2016)
B. Batell et al, PRD 95, 075003 (2017)
M. Bauer et al, JHEP. 2022, 1 (2022)



- Event signature is four tracks with missing energy
- Muons used to compute $M_{\rm recoil}(\mu\mu)$, which peaks for signal
- Background suppression via neural network
- $e^+e^+ \rightarrow e^+e^-X_{\rm had}$ and $e^+e^+ \rightarrow 4\ell(\gamma)$ backgrounds not included in simulation

Search for $\tau\tau$ resonance

- No significant excess observed in 62.8 fb⁻¹
- Limits on $e^+e^- \rightarrow X(\rightarrow \tau^+\tau^-)\mu^+\mu^-$ cross section translated to limits on leptophilic scalar, Z', and ALP mediator interpretations

I. Adachi et al. (Belle II Collaboration) Phys. Rev. Lett. 131, 121802 (2023)







Getting Involved with Dark Sector Physics at Belle II

- Many opportunities to get involved with ongoing and new Dark Sector Physics analyses at Belle II!
- Only small selection of analyses was covered in this talk
- Mailing-list:

physics-dark-low-multiplicity@belle2.org

• Weekly Meetings (alternates timezones):

https://indico.belle2.org/event/12405/