New Physics in EWP B decays with missing E

David Marzocca



Outline

Flavour in the **Standard Model** and beyond.

Part I Effective Field Theory approach to New Physics (NP).

The New Physics Flavour Problem and the need for a flavour structure.

Rare decays as probes of heavy New Physics: focus on golden-channel decays.

Part II What is the preferred flavour alignment of NP?

Hints for a consistent picture emerging from data.

Beyond the leading SMEFT.

Part III L-violating operators in B→Kw.

Light New Physics in $\mathbf{B} \rightarrow \mathbf{K} \mathbf{X}$: q^2 spectrum shape and ALPs.

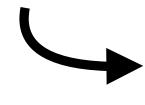
Part I

The success of symmetries and power counting

The Flavour of the Standard Model

Most of the richness and complexity of the Standard Model is in the Yukawa sector:

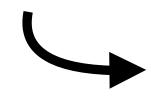
$$\mathcal{L}_{SH}^{vik} = -y_e^{ij} \, \overline{L}_i^i \, e_j^i \, H - y_a^{ij} \, \overline{Q}_i^i \, d_j^i \, H - y_a^{ij} \, \overline{Q}_i^i \, u_j^i \, \widetilde{H} + h.c.$$



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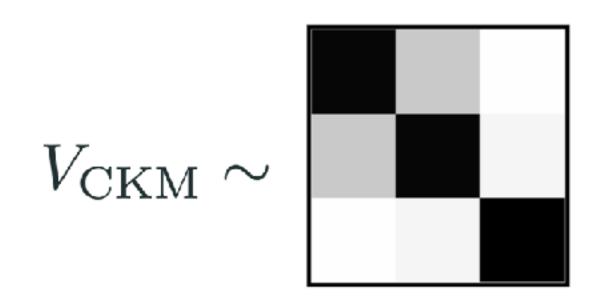
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- hierarchical fermion masses

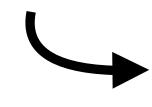
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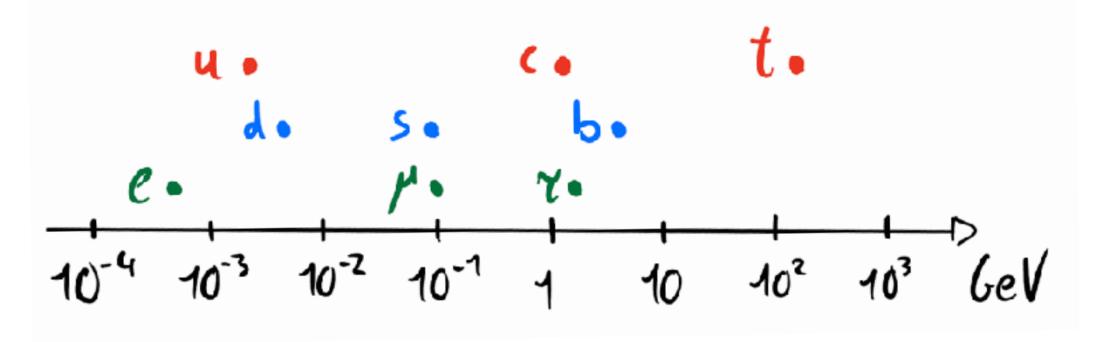


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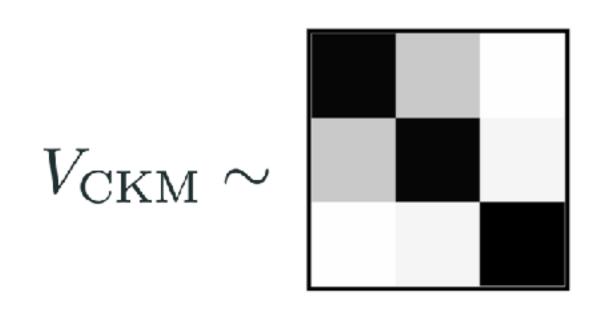
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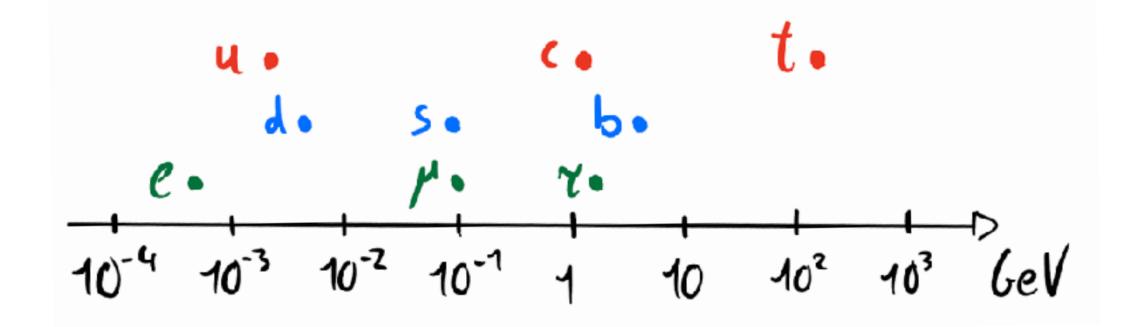


However, the SM gives no explanation for these hierarchies. Is there a more fundamental underlying theory which does?

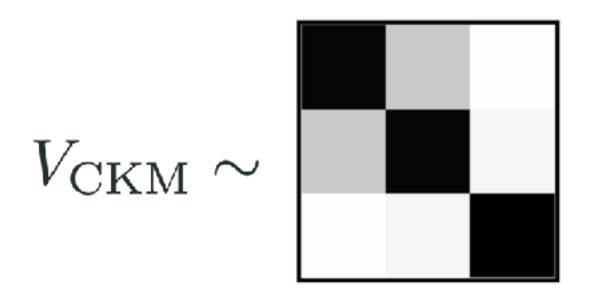
SM Flavour Puzzle

- hierarchical fermion masses

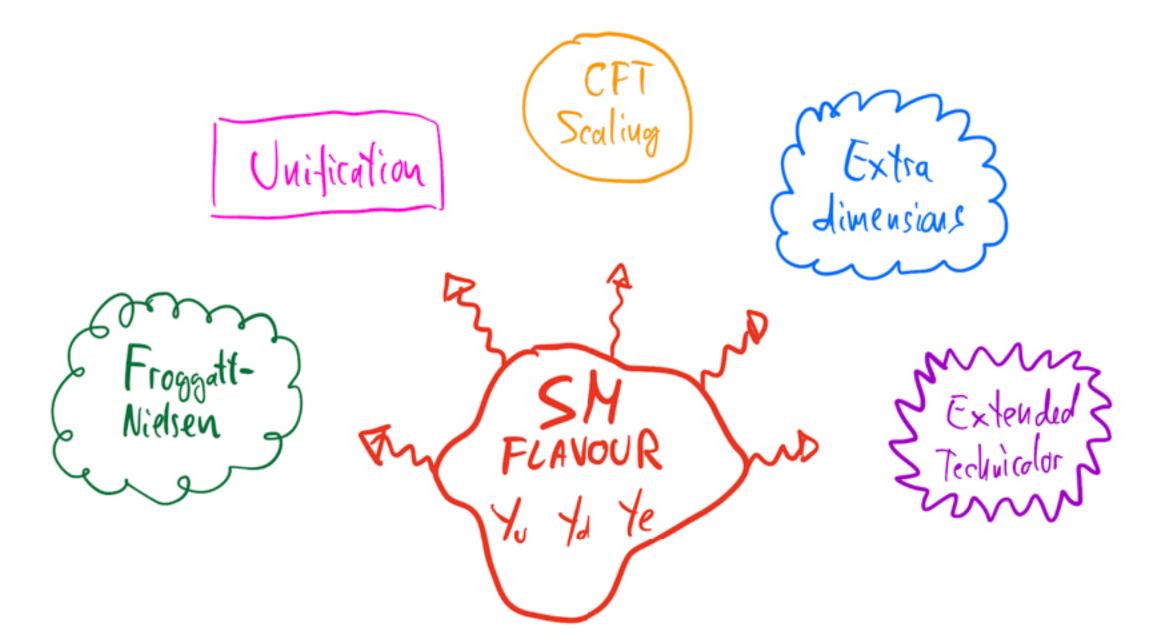
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This puzzle in general doesn't point to a specific New Physics scale for its solutions. They could be anywhere from near the TeV till up to GUT/Planck.

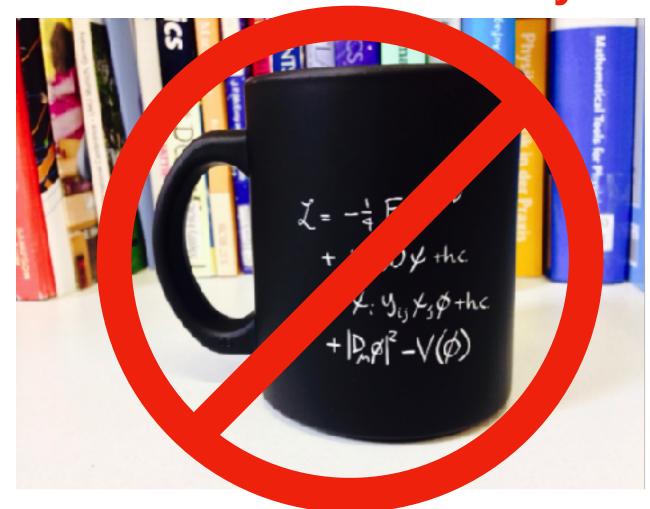


Necessarily Flavourful New Physics:

- non universal
- flavour changing

We know that the Standard Model must be extended at some high energy scale Λ .

not the whole story



- What is the nature of **dark matter**?
- What is the origin of neutrino masses?
- Why does **QCD** conserve **CP**?
- What is the origin of fermion masses and mixings?
- Why the specific assignment of charges in the SM?
- Why is the electroweak scale smaller than the Planck scale? How is it stable?
- What is the origin of the baryon asymmetry of the Universe?
- What induces the accellerated expansion of the Universe?
- What was the mechanism underlying inflation?
- How does gravity behave at the quantum level?

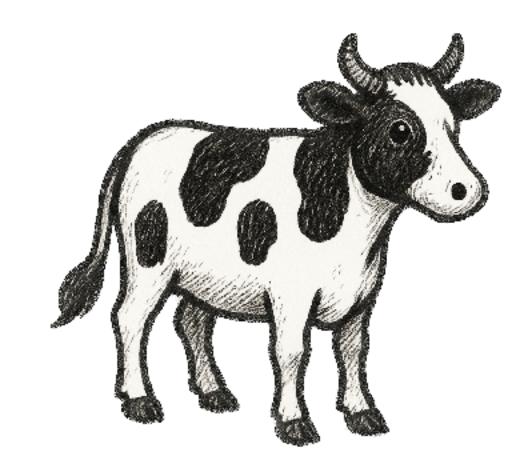
However, we don't know what Λ is, or what New Physics looks like.

So, how can we parametrise New Physics effects in our experiments?

Say New Physics has a characteristic scale $\Lambda\gg E_{exp}$ (or equivalently $\lambda\ll\ell_{exp}$), then we can only look at its effects on processes involving SM particles.

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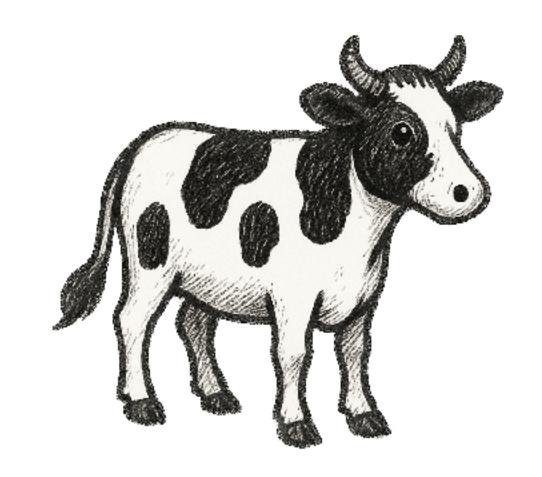
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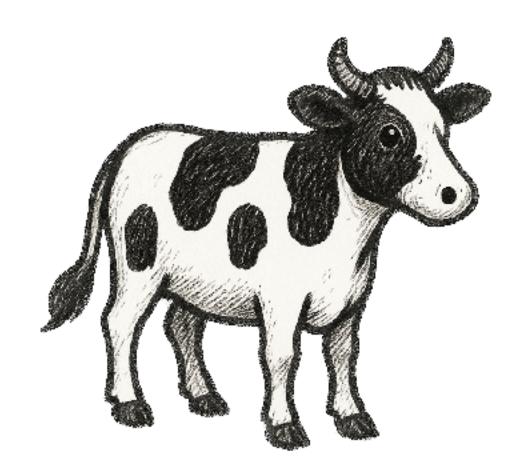
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The same model seen from a distance (from low energy)

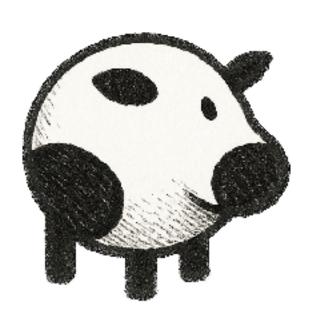
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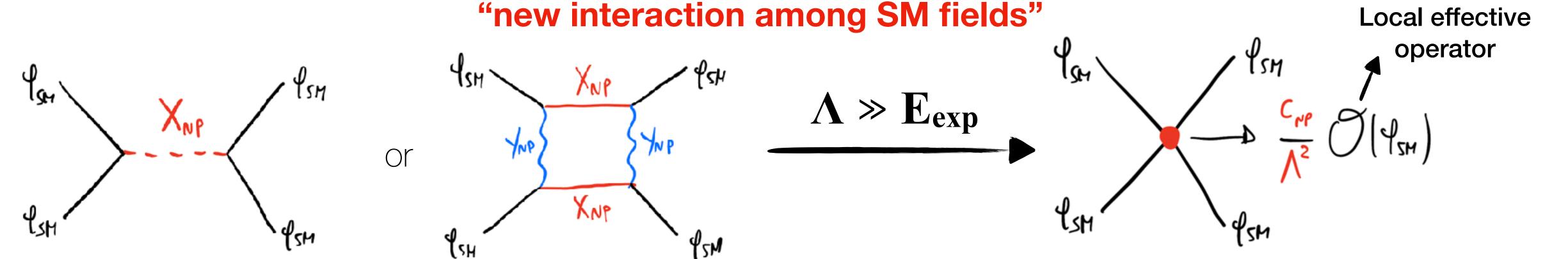
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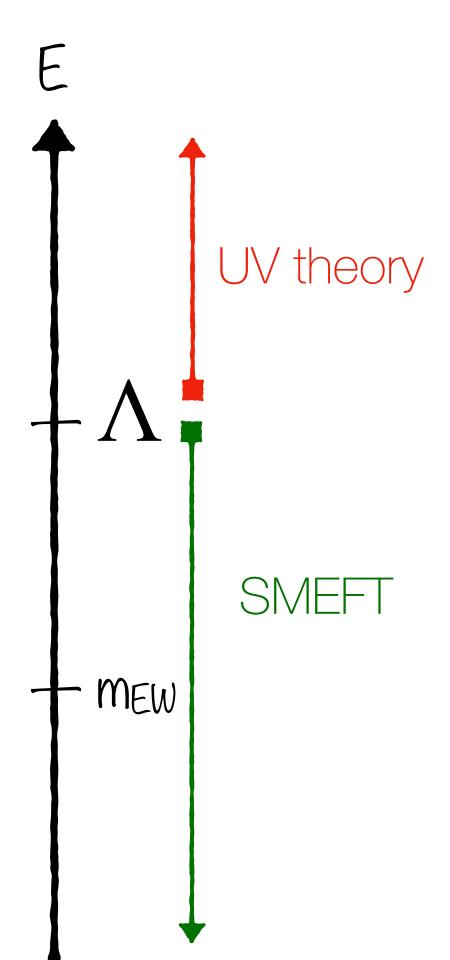
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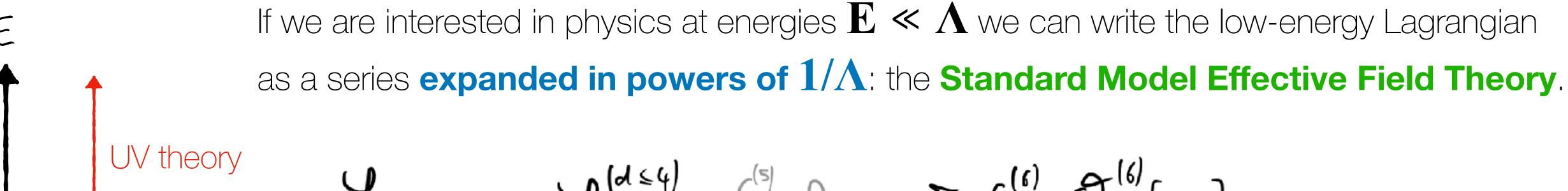
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UV theory

$$\mathcal{L}_{SMEFT} = \mathcal{L}_{SM}^{(A \le 4)} + \mathcal{L}_{i}^{(5)} \mathcal{O}_{i}^{(6)} \left[\mathcal{L}_{SM} \right] + \mathcal{O}(\Lambda^{-4})$$
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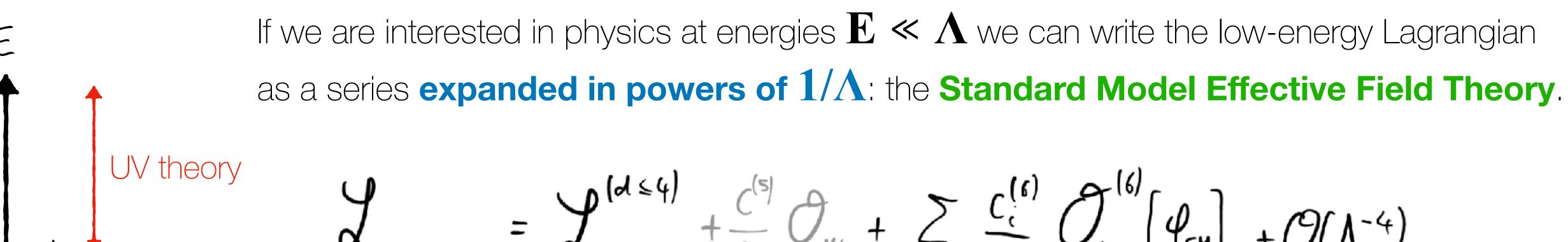


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Symmetries and power counting:

There can be different scales Λ associated to the violation of different SM properties: quark flavour, lepton flavour, L and B violation, etc..

SM: Accidental Features

The structure of the Standard Model implies several **accidental features**, i.e. properties that arise automatically, not imposed by hand.

Symmetries & conservation laws: conservation of B, L_e , L_{μ} , L_{τ}

Custodial symmetry: An approximate global SU(2)_C symmetry in the Higgs sector. Protects the ratio $m_W/(\cos\theta_W m_Z) \approx 1$.

Approximate <u>U(3)</u>⁵ Flavour Symmetry: Broken only by Yukawa interactions

Absence of FCNC at tree-level: Z boson, photon and gluon couple in a flavour-conserving way + Higgs Yukawa couplings are small.

Small CP-violation effects, even though the CP-phase is large: small quark masses and mixing angles.

Lepton-Flavour Universality: SM gauge couplings are generation-independent + Yukawa couplings are small and hierarchical (e.g. m_{e,µ} « m_b)

Massless neutrinos: a SM neutrino mass term is forbidden by gauge symmetries.

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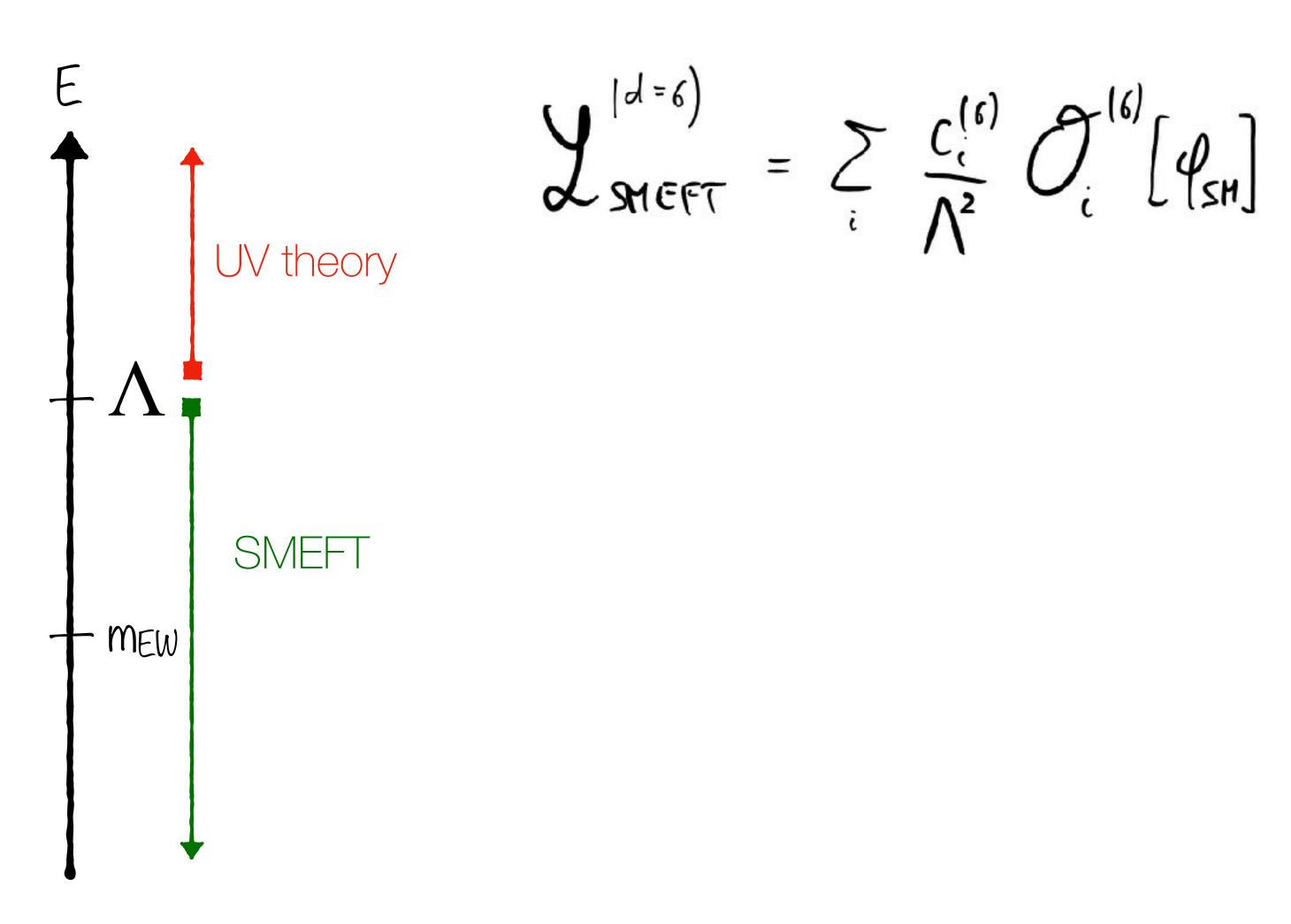
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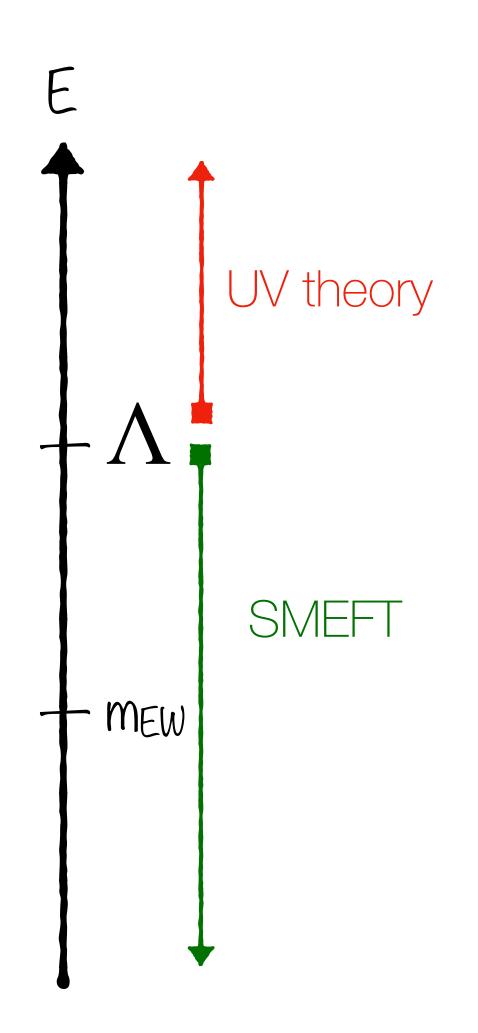
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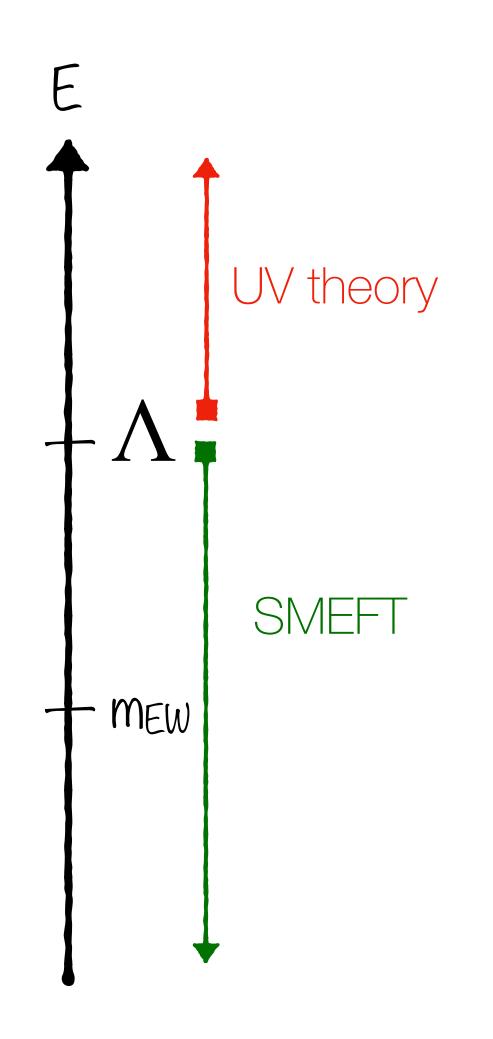


$$\int_{MEFT}^{|d=6|} = \sum_{i} \frac{C_{i}^{(6)}}{\Lambda^{2}} O_{i}^{(6)} [\varphi_{SH}]$$



$$\int_{\Re \mathbb{R} \in \mathbb{R}^{+}}^{|\mathcal{A}=6|} = \sum_{i} \frac{C_{i}^{(6)}}{\Lambda^{2}} O_{i}^{(6)} [\varphi] \qquad \text{in general violate all the accidental symmetries and properties of the SM}$$

We can expect large effects in rare or forbidden processes!

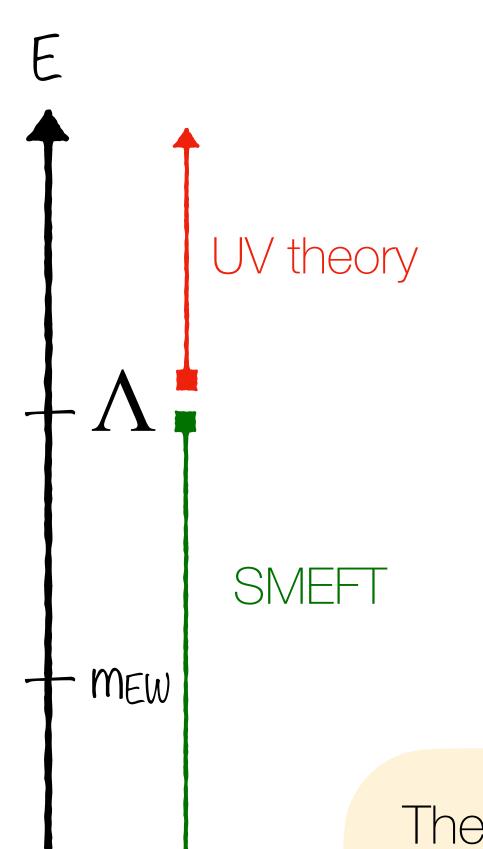


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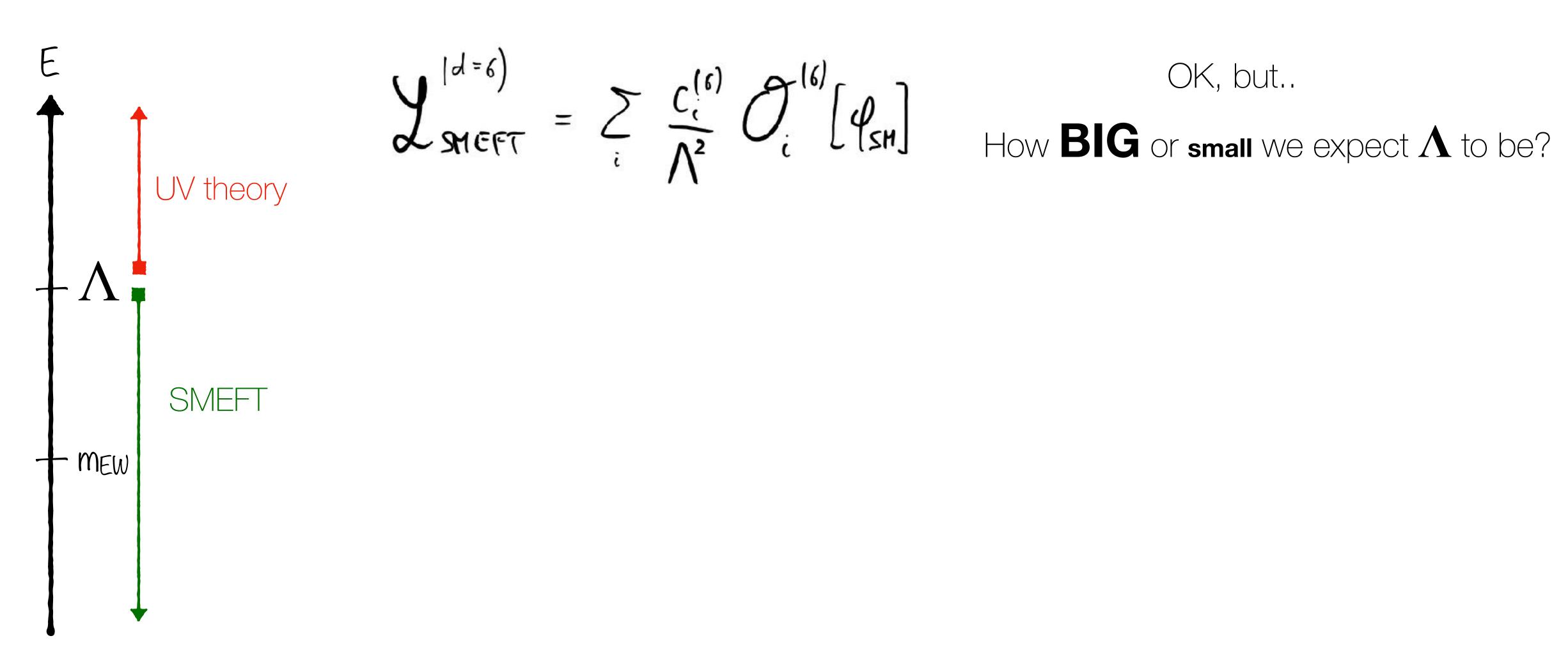
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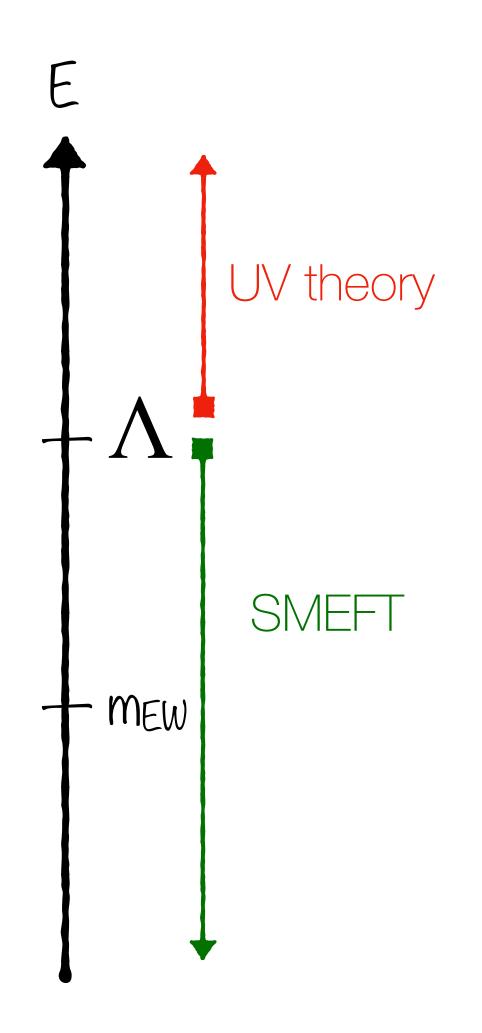
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The SMEFT is a consistent framework to systematically parametrise our knowledge of fundamental interactions between the known particles.

Every little improvement in any direction in the (big) EFT parameter space means that we learn something more of how particles behave at microscopic scales.

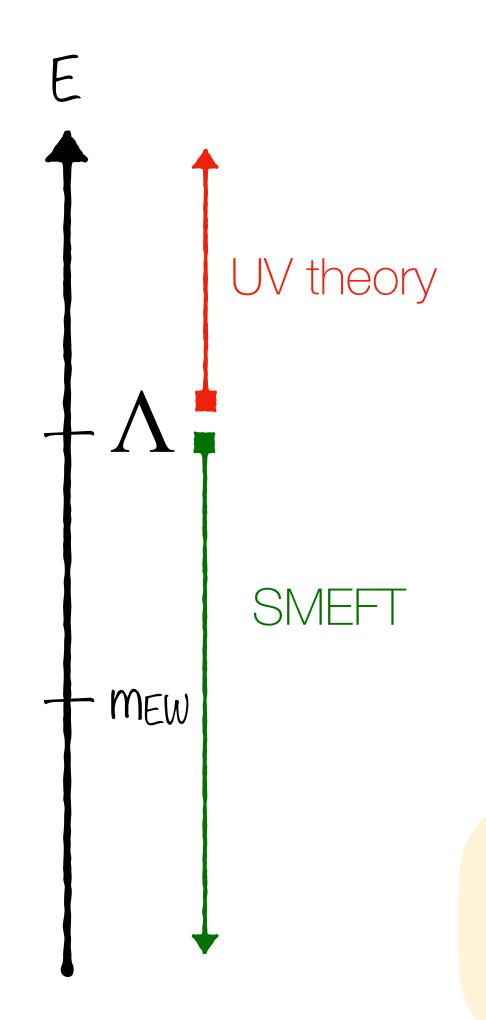




$$\mathcal{L}_{\text{SMEFT}}^{|d=6)} = \sum_{i} \frac{c_{i}^{(6)}}{\Lambda^{2}} \mathcal{O}_{i}^{(6)} [\varphi_{\text{SM}}] \qquad \text{How BIG or small we expect } \Lambda \text{ to be?}$$

Since the SM is renormalisable, we don't have a clear target (except $\Lambda \leq M_{Pl}$)

Our experiments typically test values of the NP scale not too far from the TeV.



$$\mathcal{L}_{\text{SMEFT}}^{|d=6)} = \sum_{i} \frac{c_{i}^{(f)}}{\Lambda^{2}} \mathcal{O}_{i}^{(6)} [\varphi_{\text{SM}}] \qquad \text{OK, but..}$$
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Motivated Reasons for a "low" 1:

Hierarchy problem of the EW scale, $\Lambda \sim \text{TeV}$

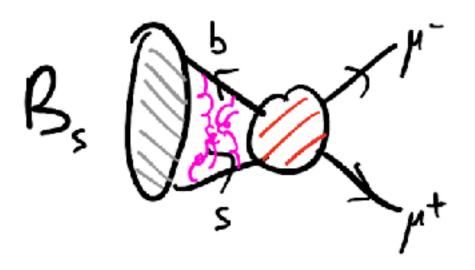
Experimental signatures of BSM physics (anomalies)

> (it depends on the measurement)

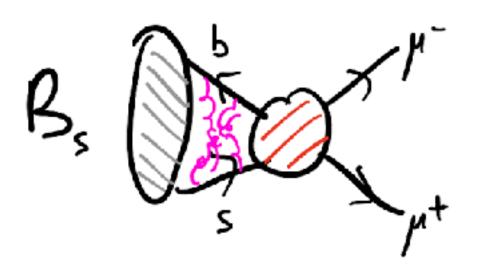
WIMP miracle

for Dark Matter

 $\Lambda \sim 0.1 - O(10) \text{ TeV}$



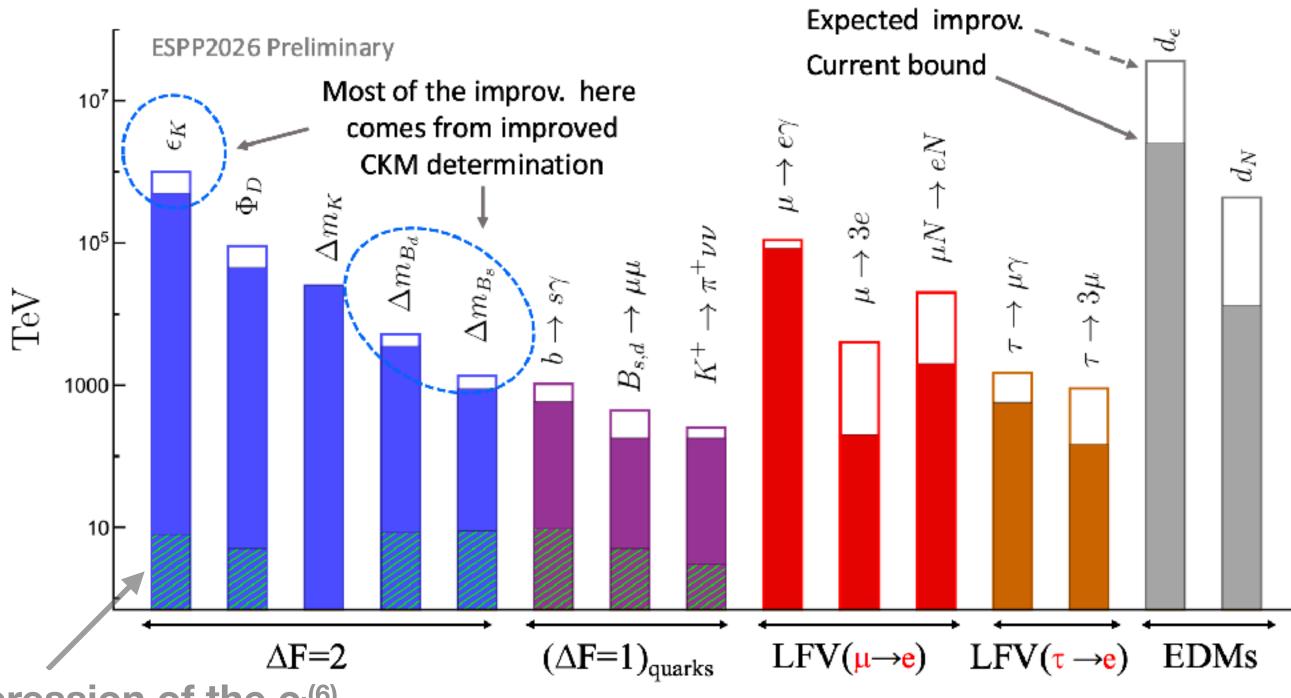
Measuring rare flavour transitions puts strong constraints on New Physics with generic flavour structure.

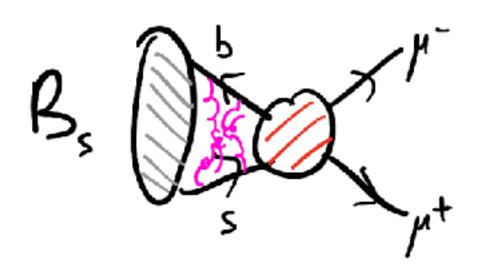


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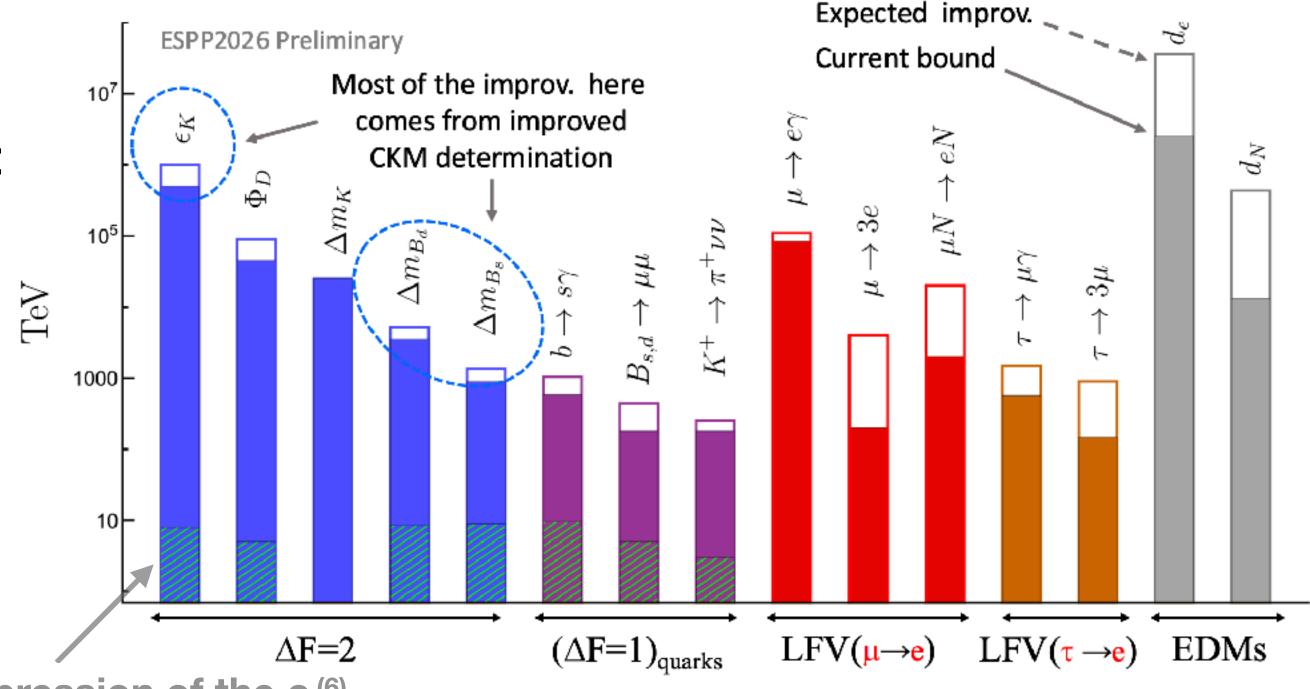
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If New Physics is present at the TeV scale, its flavour structure should be constrained by some "protecting" principle (symmetry or dynamics): the BSM Flavour Problem.

$$\mathcal{L}_{SMEFT}^{|d=6)} = \sum_{i} \frac{C_{i}^{(6)}}{\Lambda^{2}} \mathcal{O}_{i}^{(6)} [\varphi_{SM}]$$

Need: c⁽⁶⁾(Flav. Violating) ≪ 1 !!



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- Solutions to the Hierarchy Problem
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Cabibbo angle

$$\lambda \sim \sin \theta_c$$
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New physics likes the Top

(1)

Bounds from direct searches @ LHC are stronger for light fermions than for third generation ones.

E.g. squark: $M_{\tilde{q}_{(1,2)}} \gtrsim 2 \; TeV$ $M_{\tilde{t},\tilde{b}} \gtrsim 1.4 \; TeV$

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E.g. Scalar LQ:

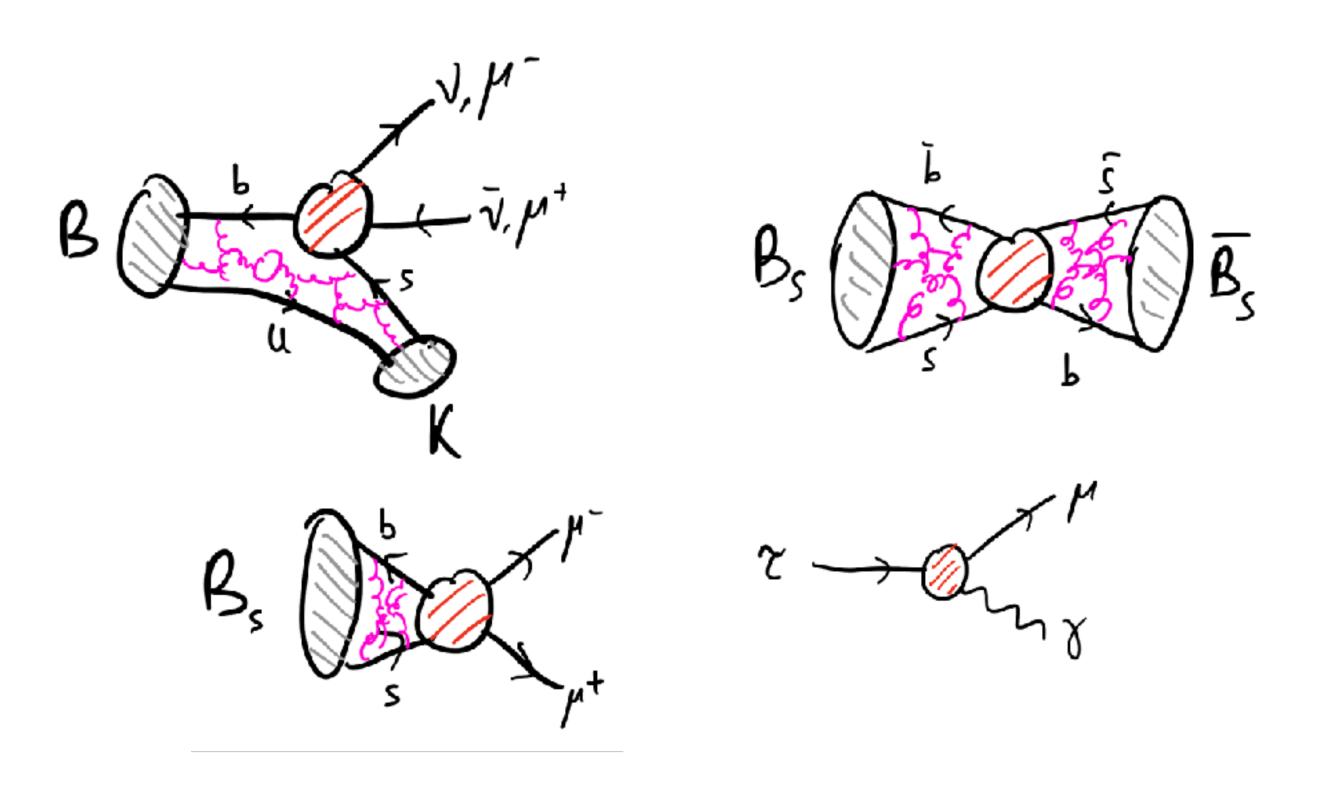
Non-universal couplings preferred

U(2)-like: $\xi_{1,2} \ll 1$

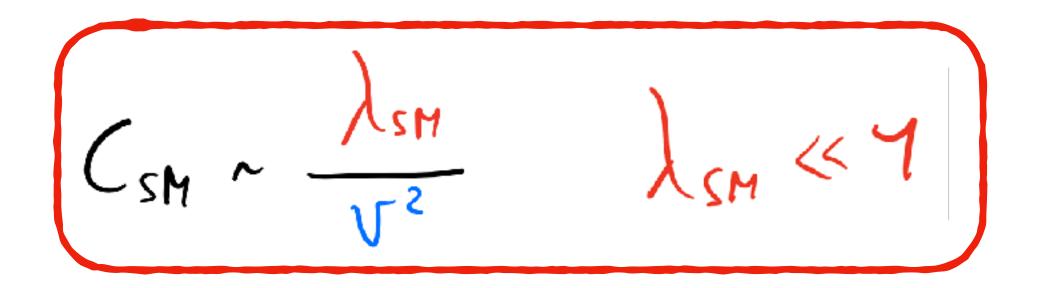
Part II

In rarity lies power

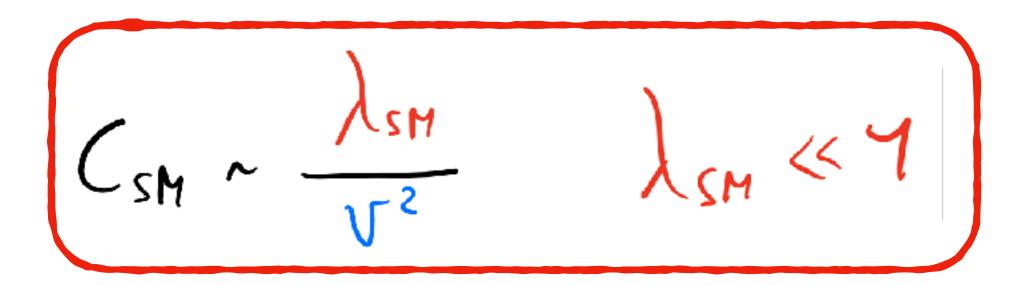
Probing New Physics with Rare or Forbidden Processes



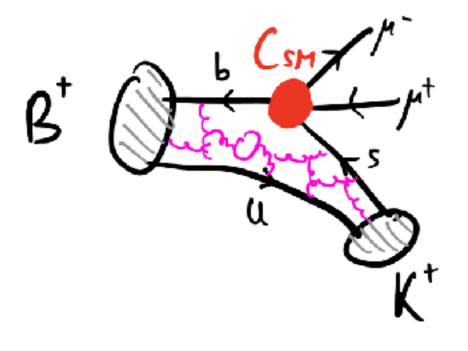
Consider a rare low-energy FCNC process in the SM Short-distance low-energy EFT coefficient



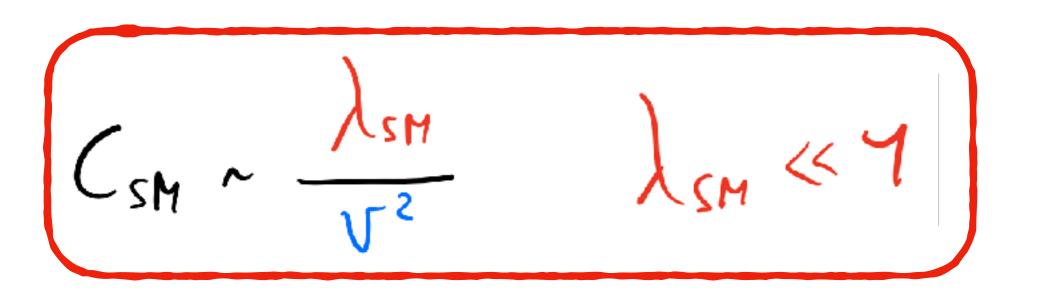
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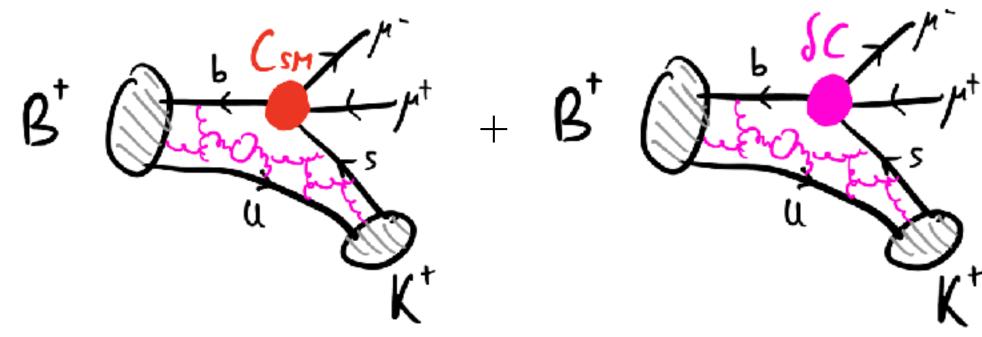


Example:
$$C_{sm}^{sb} \sim \frac{d}{4\pi} \frac{V_{ts}V_{tb}}{v^2}$$

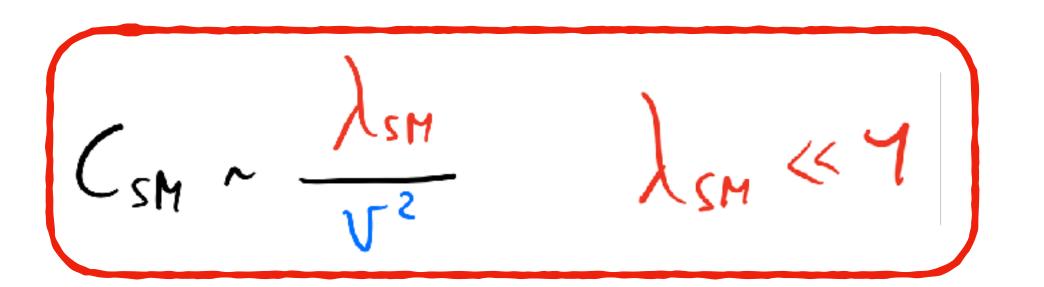


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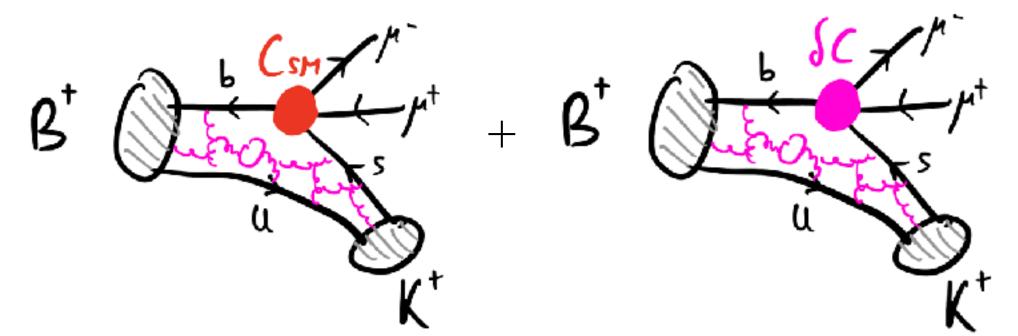




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Example:



Let us add a **SMEFT contribution**:

$$C_{\text{EFT}} \sim \frac{C}{\Lambda^2}$$

$$\frac{\int C}{C_{SM}} \sim \frac{C}{\lambda_{SM}} \frac{v^2}{\Lambda^2}$$

Relative deviation in the short-distance coefficient

> i.e. size of the deviation compared to the SM <

$$\frac{\int C}{C_{SM}} \sim \frac{V^2}{V^2}$$

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Say we measure the short-distance contribution C_{SM} with 10% precision.

$$C = C_{SH} \left(1 \pm 10\% \right)$$

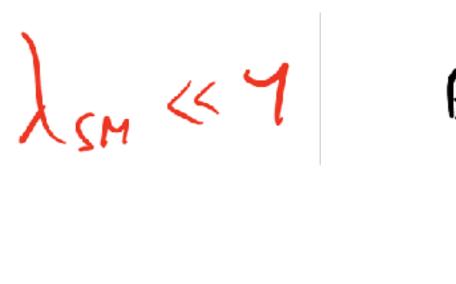
$$SC < C_{SL} \cdot 10\%$$

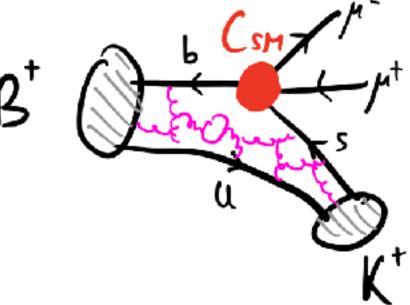
$$\frac{\int C}{C_{SM}} \sim \frac{C}{\lambda_{SM}} \frac{v^2}{\Lambda^2}$$

$$\frac{1}{\sqrt{s}} = \frac{1}{\sqrt{s}}$$

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Say we measure the short-distance contribution C_{SM} with 10% precision.

$$\frac{\int C}{C_{SH}} \sim \frac{C}{\lambda_{SM}} = \frac{\nabla^2}{\Lambda^2} < 10\%$$

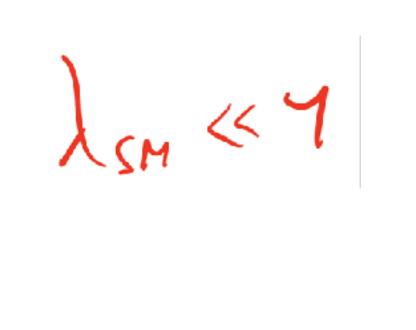
$$C = C_{SH} \left(1 \pm 10\% \right)$$

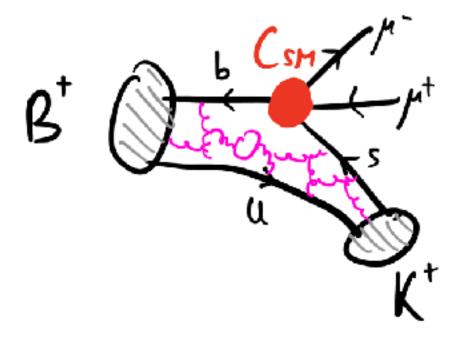
$$SC < C_{SH} \cdot 10\%$$

$$\frac{\int C}{C_{SM}} \sim \frac{C}{\lambda_{SM}} \frac{v^2}{\Lambda^2}$$

$$\frac{1}{\sqrt{s_{M}}} \sim \frac{\lambda_{sM}}{\sqrt{s_{M}}}$$

$$\frac{1}{\sqrt{s_{M}}} \sim \frac{1}{\sqrt{s_{M}}}$$





Say we measure the short-distance contribution C_{SM} with 10% precision.

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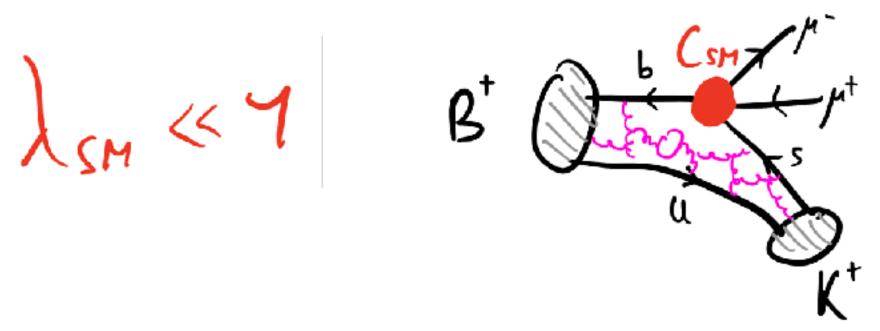
$$SC < C_{SH} \cdot 10\%$$

$$\frac{\int C}{C_{SM}} \sim \frac{C}{\lambda_{SM}} \frac{v^2}{\Lambda^2} < 10\% \qquad \frac{\text{for c=1}}{\sqrt{\lambda_{SM}}} \Lambda > \frac{v}{\sqrt{\lambda_{SM}}} \sqrt{10\%} \sim \frac{0.8}{\sqrt{\lambda_{SM}}} \text{TeV}$$

$$\frac{\int C}{C_{SM}} \sim \frac{C}{\lambda_{SM}} \frac{v^2}{\Lambda^2}$$

$$C_{SM} \sim \frac{\lambda_{SM}}{V^2}$$

$$C_{EFT} \sim \frac{C}{\Lambda^2}$$



Say we measure the short-distance contribution C_{SM} with 10% precision.

$$C = C_{SH} \left(1 \pm 10\% \right)$$

$$SC < C_{SH} \cdot 10\%$$

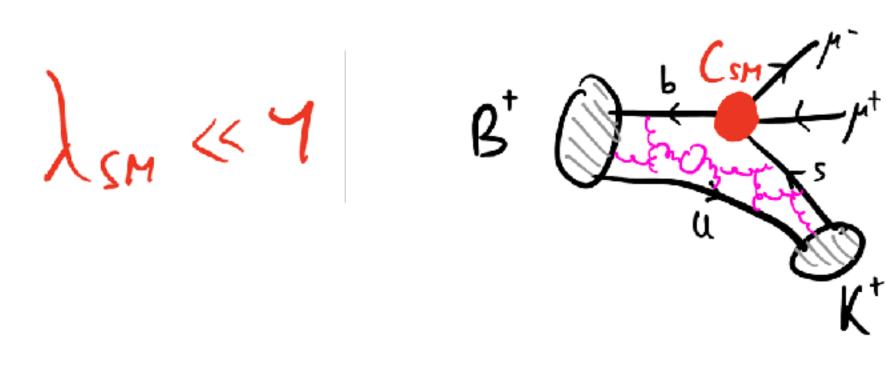
$$\frac{\int C}{C_{SM}} \sim \frac{C}{\lambda_{SM}} \frac{v^2}{\Lambda^2} < 10\% \qquad \frac{\text{for c=1}}{\sqrt{\lambda_{SM}}} \sim \frac{0.8}{\sqrt{\lambda_{SM}}} \text{ TeV}$$

Measuring this precisely puts strong constraints on the EFT combination c/Λ^2 , the better the smallest λ_{SM} is.

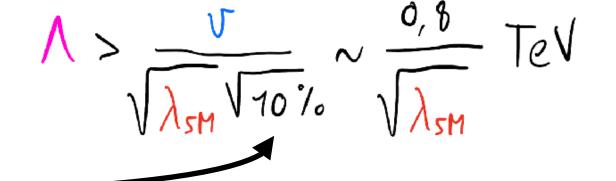
$$\frac{\int C}{C_{SM}} \sim \frac{C}{\lambda_{SM}} \frac{v^2}{\Lambda^2}$$

$$C_{SM} \sim \frac{\lambda_{SM}}{V^2}$$

$$C_{EFT} \sim \frac{C}{\Lambda^2}$$



For this goal it is also crucial to have the smallest possible uncertainty on the short-distance contributions:



Exp

- Very large statistics
- Small backgrounds and systematics
- Good control over the SM prediction:
- TH

- **SM inputs** (CKM matrix elements)
- QCD matrix elements (form factors)
- control over the possible long-distance contributions

Golden-channels of rare decays

$$b \rightarrow s \ v \ \overline{v}$$
 $B \rightarrow K(*) \ v \ \overline{v}$
BaBar, Belle, Belle II (JPARC)

$$s o d v \overline{v}$$
 $K^+ o \pi^+ v \overline{v}, \quad K_L o \pi^0 v \overline{v}$ NA62 (CERN) KOTO (JPARC)

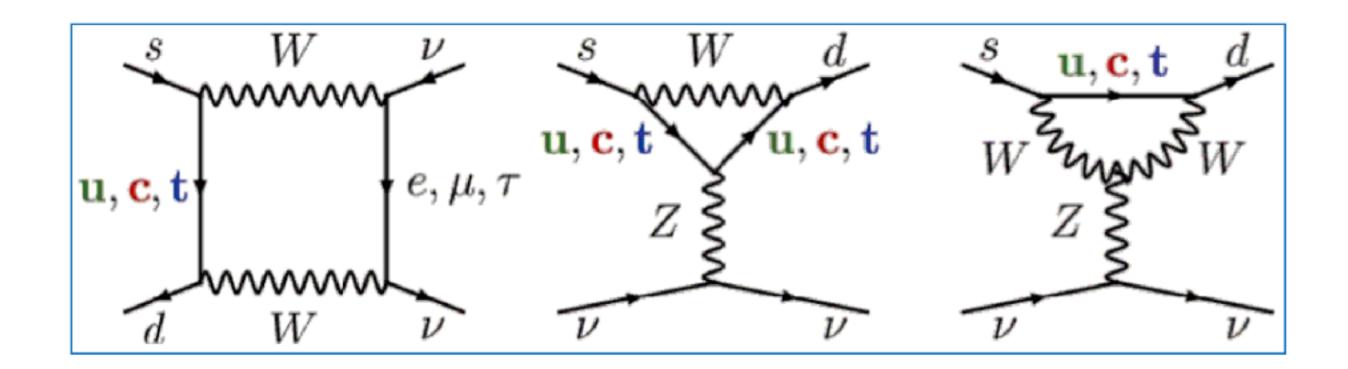
Golden-channels of rare decays

$$b o s \ v \ \overline{v}$$
 $s o d \ v \ \overline{v}$ $B o K^{(*)} \ v \ \overline{v}$ $K^+ o \pi^+ \ v \ \overline{v}, \quad K_L o \pi^0 \ v \ \overline{v}$ BaBar, Belle, Belle II (JPARC) NA62 (CERN) KOTO (JPARC)

Precise SM predictions possible due to absence of long-distance QCD effects:

[see previous lecture by Wolfgang] neutrinos do not couple to the electromagnetic current.

see 1409.4557, 1503.02693, 2109.11032, 2301.06990, ...



Main th. uncertainties due to:

- Hadronic form factors (Lattice QCD)
- CKM matrix elements

$$B^{+} \to K^{+} \nu \bar{\nu}$$
 $(5.06 \pm 0.14 \pm 0.28) \times 10^{-6}$ $B^{0} \to K_{S} \nu \bar{\nu}$ $(2.05 \pm 0.07 \pm 0.12) \times 10^{-6}$ $B^{+} \to K^{*+} \nu \bar{\nu}$ $(10.86 \pm 1.30 \pm 0.59) \times 10^{-6}$ $B^{0} \to K^{*0} \nu \bar{\nu}$ $(9.05 \pm 1.25 \pm 0.55) \times 10^{-6}$

Becirevic et al. 2301.06990

The SM rate is suppressed by loop and small CKM factors: high sensitivity to New Physics.

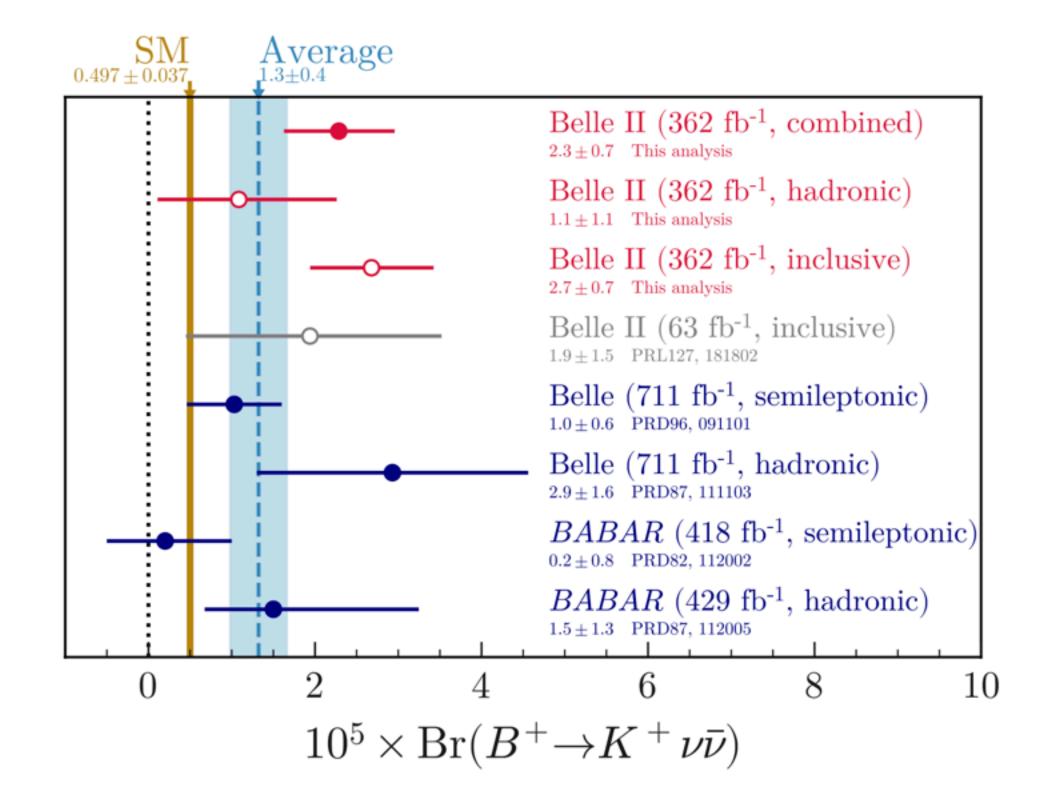
$B \longrightarrow K(*) \nu \overline{\nu}$

$$BR(B^+ \to K^+ \nu \overline{\nu})_{SM} = (0.444 \pm 0.030) \times 10^{-5}$$

Becirevic et al. 2301.06990

Belle-ll₂₀₂₃: BR($B^+ \to K^+ \nu \overline{\nu}$) = (2.3 ± 0.6) × 10-5

Combination: BR($B^+ \rightarrow K^+ \nu \overline{\nu}$) = $(1.3 \pm 0.4) \times 10^{-5}$



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Becirevic et al. 2301.06990

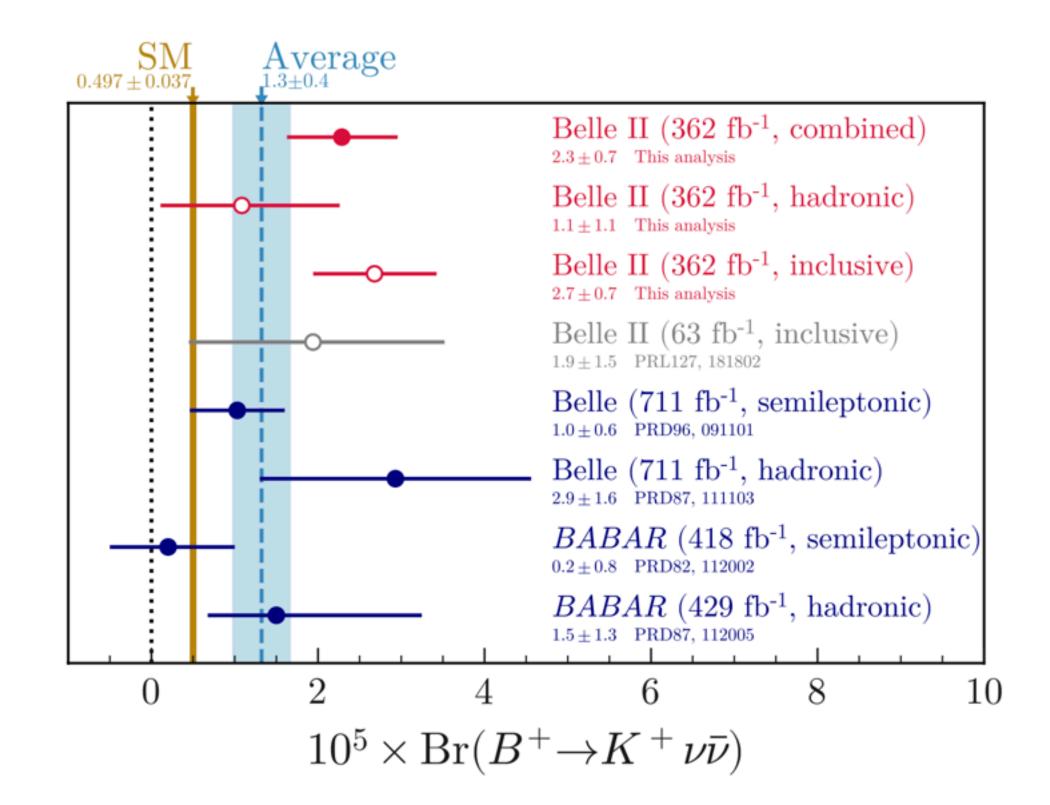
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Becirevic et al. 2301.06990

Belle₂₀₁₇: $BR(B \rightarrow K^* \nu \overline{\nu}) < 2.7 \times 10^{-5}$ @ 90%CL



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Becirevic et al. 2301.06990

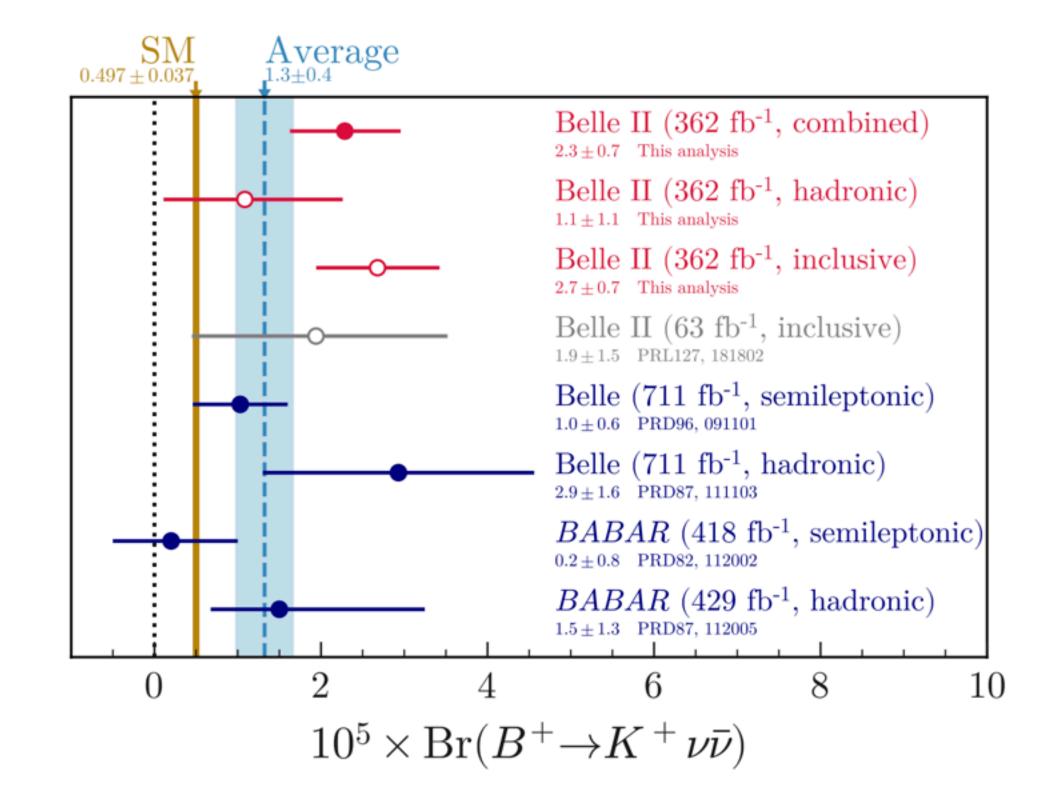
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$$R_{K}^{v} = \frac{BR(B \rightarrow Kvv)}{BR(B \rightarrow Kvv)^{SM}} = 2,93 \pm 0,90$$

$$R_{K^{\pm}}^{v} = \frac{BR(B \rightarrow K^{*}vv)}{BR(B \rightarrow K^{*}vv)^{SM}} = 1.0 \pm 1.4^{*}$$

^{*} Assuming SM to be the central value, also motivated by a small 2σ excess in the K*+ channel.

$B \to K(*) \nu \overline{\nu}$

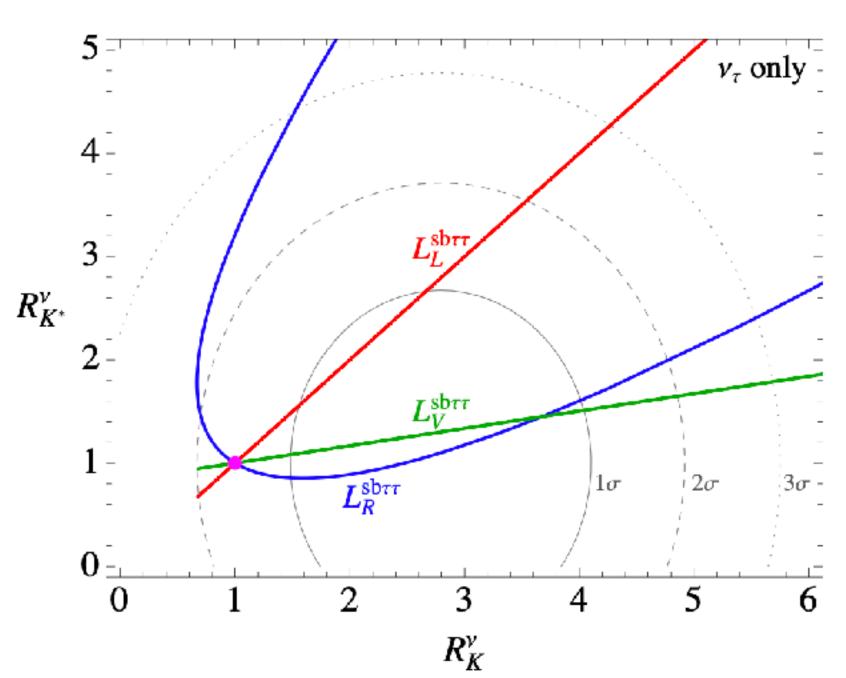
Assuming only NP in tau

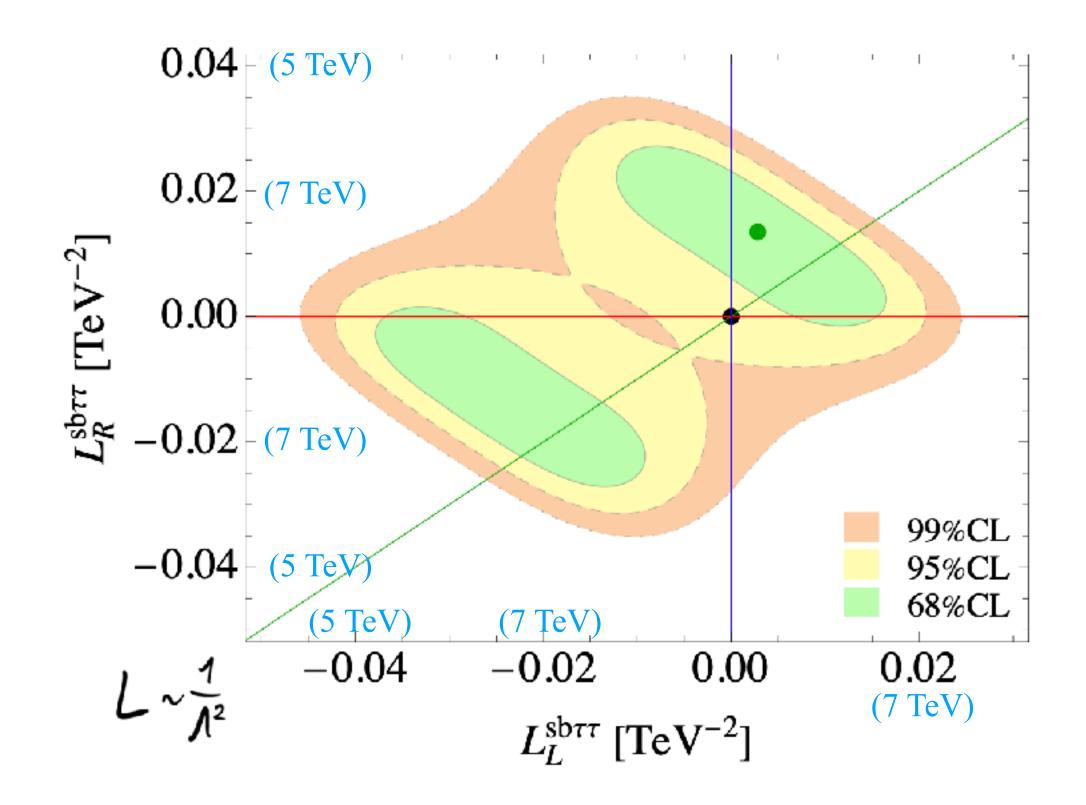
$$\mathcal{L}_{EFT} > \lfloor_{L,R}^{ijtr} \left(\bar{d}_{iL,R}^{i} \mathcal{V}_{\mu} d_{jL,R} \right) \left(\bar{\nu}_{e} \mathcal{V}^{\mu} \nu_{e} \right)$$

DM, M. Nardecchia, A. Stanzione, C. Toni [2404.06533]

The limits from R(K) and $B_s \rightarrow \mu\mu$ disfavour interpretations with electron or muon neutrinos

$$L_{V,A}^{sb\alpha\beta} \equiv L_R^{sb\alpha\beta} \pm L_L^{sb\alpha\beta}$$





 $\Lambda_{bsvv} \sim 7 \text{ TeV}$

Future Belle II results (in particular from the K* mode) will help to clarify the preferred chiral structure.

$$K^+ \longrightarrow \pi^+ \ \nu \ \overline{
u}, \quad K_L \longrightarrow \pi^0 \ \nu \ \overline{
u}$$

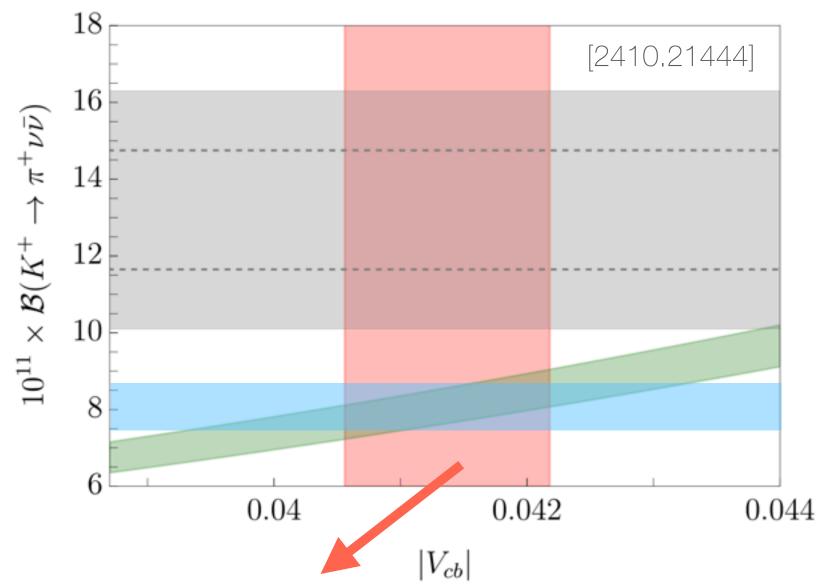
NA62 (CERN)

$$BR(K^+ \to \pi^+ \nu \overline{\nu})_{SM} = (8.09 \pm 0.63) \times 10^{-11}$$

Allwicher et al. [2410.21444] (see also Buras et al. 1503.02693, 2109.11032, etc..)

NA62₂₀₂₄:

$$\mathcal{B}(K^+ \to \pi^+ \nu \bar{\nu}) = (13.6 \, (^{+3.0}_{-2.7})_{\text{stat}} (^{+1.3}_{-1.2})_{\text{syst}}) \times 10^{-11}$$



$$|V_{cb}| = (41.37 \pm 0.81) \times 10^{-3}$$

Derived by combining exclusive and inclusive determinations. [2310,20324, 2406,10074]

KOTO (JPARC)

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Allwicher et al. [2410.21444]

KOTO₂₀₂₁:

$$BR(K_L \to \pi^0 \ \nu \ \overline{\nu}) < 4.9 \times 10^{-9}$$
 @ 90%CL

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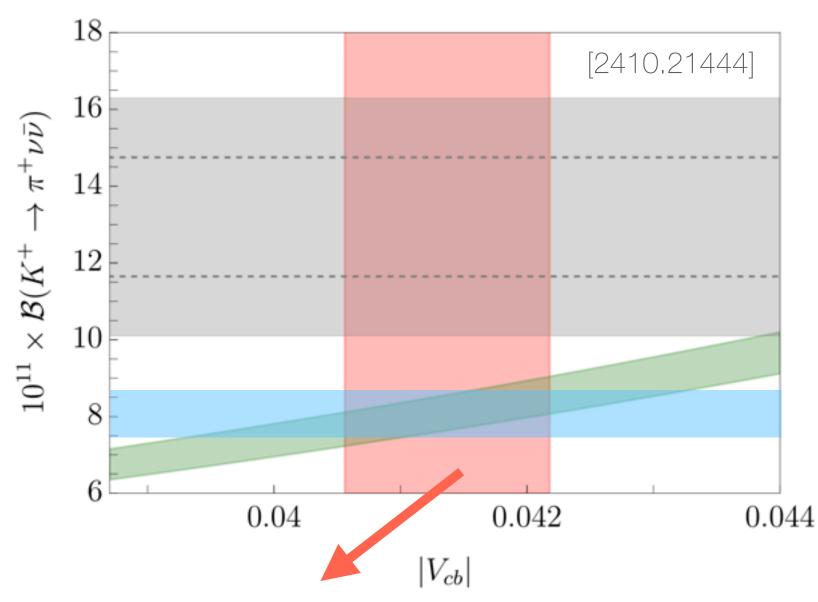
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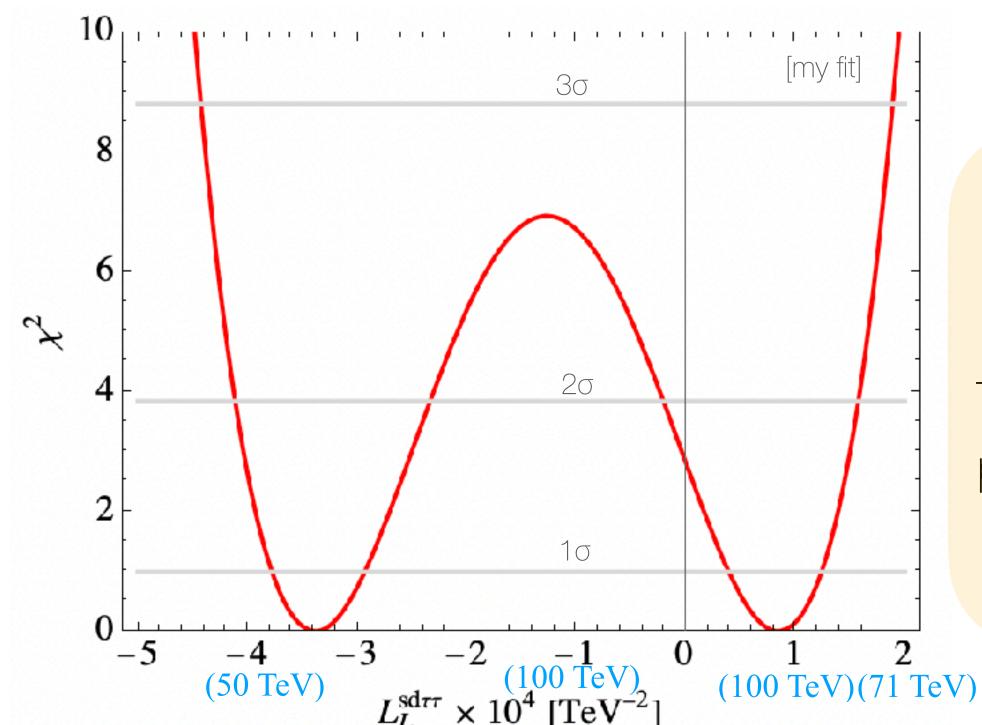
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Derived by combining exclusive and inclusive determinations. [2310.20324, 2406.10074]



 $L \sim \frac{1}{\Lambda^2}$

The slight ~1.7σ excess points to new physics scales

 $\Lambda_{\text{sdvv}} \sim 100 \text{TeV}$

Flavour alignment

How much should New Physics be aligned to the third generation?

We consider now a specific example:

- Overall New Physics scale set by the Belle-II excess in B→Kvv
- We assume a Rank-One flavour structure

Gherardi, DM, Nardecchia, Romanino 1903.10954 DM, Nardecchia, Stanzione, Toni 2404.06533

Consider the vector space spanned by the

3 generations of down quarks, $SU(3)_q$:

$$\hat{d} \sim \begin{pmatrix} 1 \\ 0 \\ 0 \end{pmatrix} \qquad \hat{S} \sim \begin{pmatrix} 0 \\ 1 \\ 0 \end{pmatrix} \qquad \hat{b} \sim \begin{pmatrix} 0 \\ 0 \\ 1 \end{pmatrix}$$

Gherardi, DM, Nardecchia, Romanino 1903.10954 DM, Nardecchia, Stanzione, Toni 2404.06533

Consider the vector space spanned by the

3 generations of down quarks, SU(3)q:

$$\hat{d} \sim \begin{pmatrix} 1 \\ 0 \\ 0 \end{pmatrix} \qquad \hat{\leq} \sim \begin{pmatrix} 0 \\ 1 \\ 0 \end{pmatrix} \qquad \hat{\int} \sim \begin{pmatrix} 0 \\ 0 \\ 1 \end{pmatrix}$$

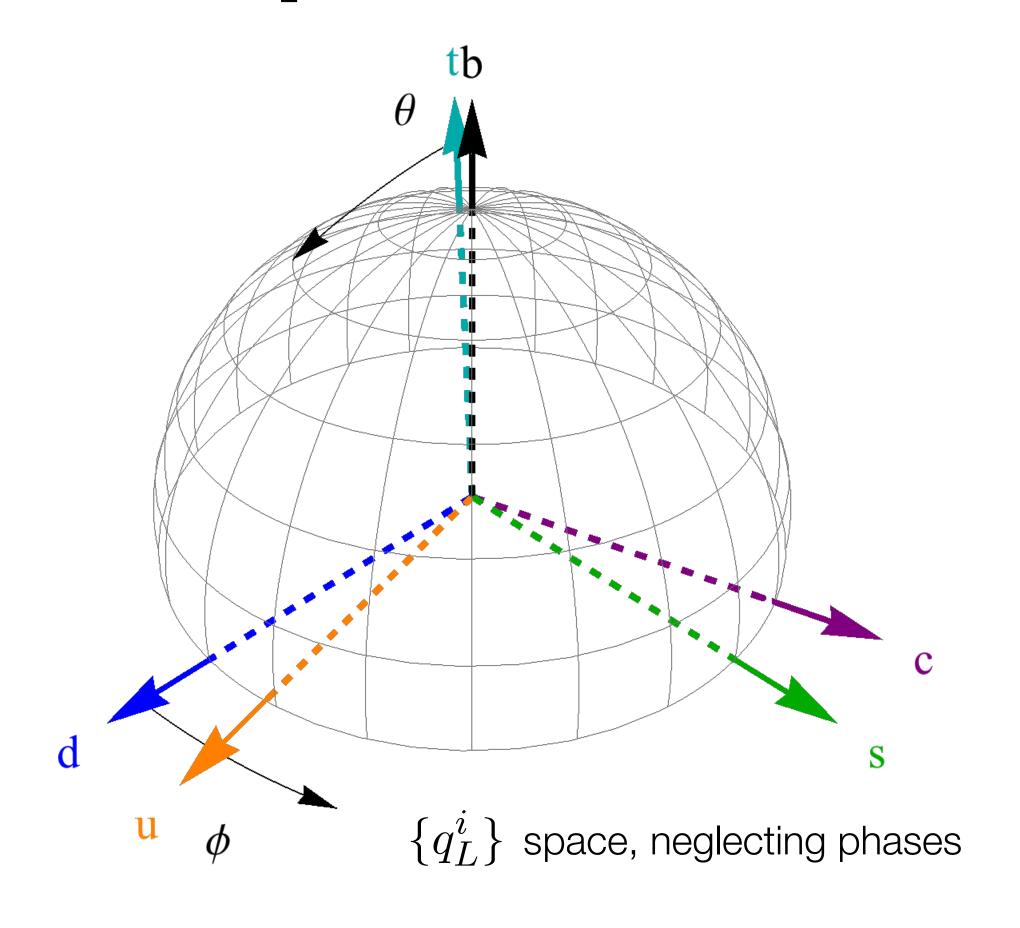
We can parametrise a generic directions as:

$$\hat{n} = \begin{pmatrix} \sin \theta \cos \phi e^{i\alpha_{bd}} \\ \sin \theta \sin \phi e^{i\alpha_{bs}} \\ \cos \theta \end{pmatrix}$$

neglecting phases, it is a unit-vector on a semi-sphere

$$\theta \in \left[0, \frac{\pi}{2}\right], \quad \phi \in \left[0, 2\pi\right), \quad \alpha_{bd} \in \left[-\frac{\pi}{2}, \frac{\pi}{2}\right], \quad \alpha_{bs} \in \left[-\frac{\pi}{2}, \frac{\pi}{2}\right]$$

The overall phase is unphysical: U(1)_B



Gherardi, DM, Nardecchia, Romanino 1903.10954 DM, Nardecchia, Stanzione, Toni 2404.06533

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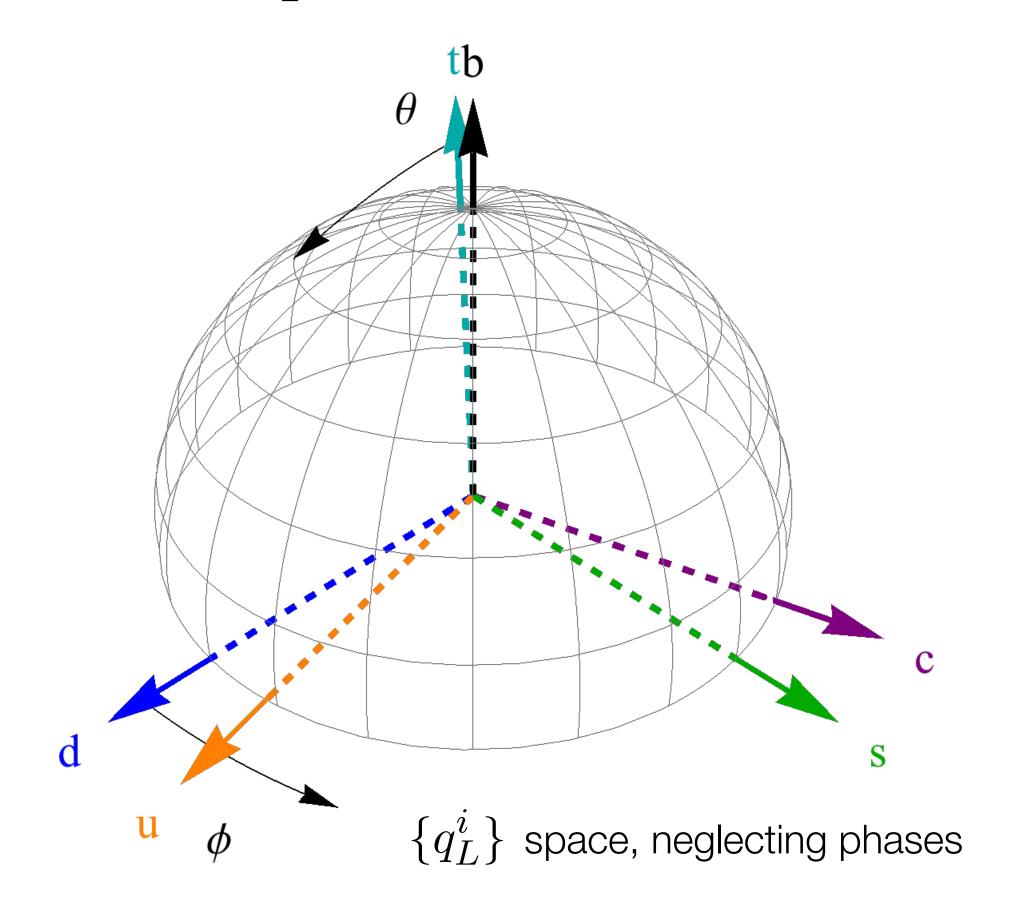
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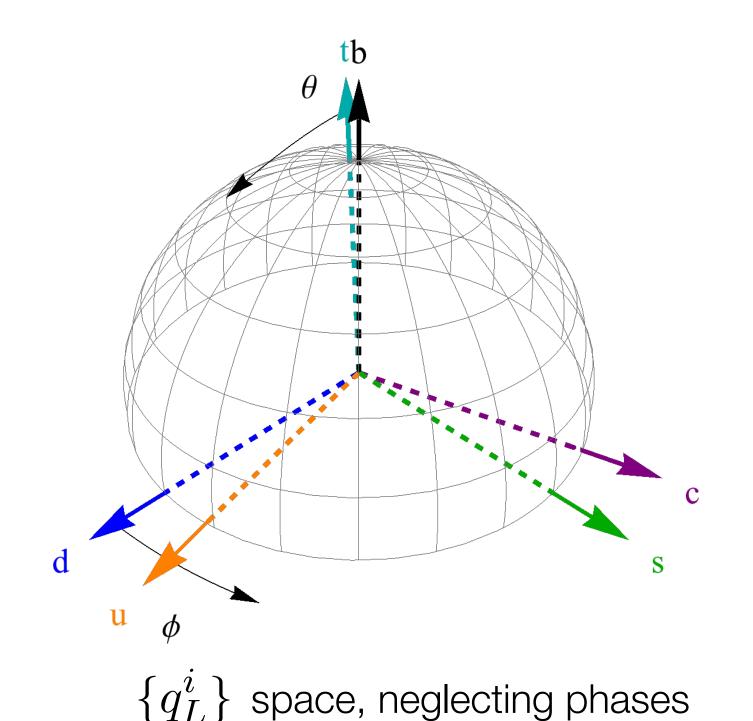
The overall phase is unphysical: U(1)_B



We show also up quarks using:

$$q_L^i = \left(\begin{array}{c} V_{ji}^* u_L^i \\ d_L^i \end{array}\right)$$

Gherardi, DM, Nardecchia, Romanino 1903.10954 DM, Nardecchia, Stanzione, Toni 2404.06533



$$\hat{n} = \begin{pmatrix} \sin \theta \cos \phi e^{i\alpha_{bd}} \\ \sin \theta \sin \phi e^{i\alpha_{bs}} \\ \cos \theta \end{pmatrix}$$

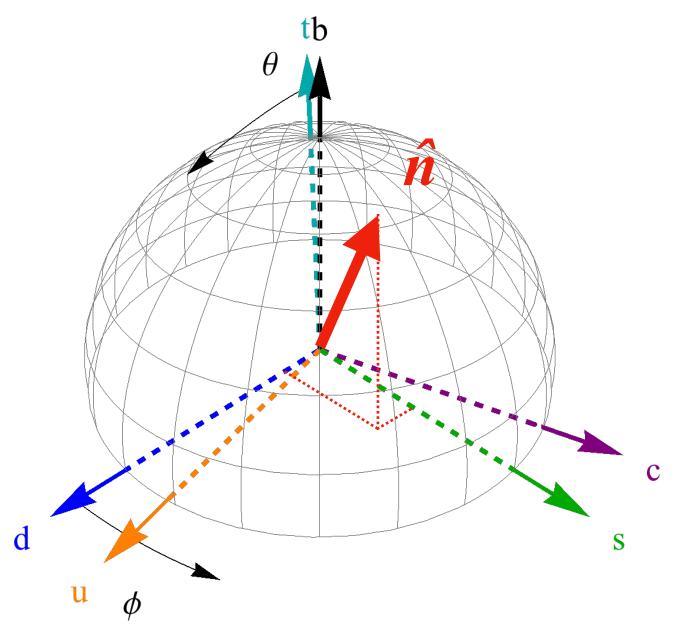
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$$q_L^i = \left(\begin{array}{c} V_{ji}^* u_L^i \\ d_L^i \end{array}
ight)$$

quark	\hat{n}	ϕ	θ	$lpha_{bd}$	$lpha_{bs}$
down	(1, 0, 0)	0	$\pi/2$	0	0
strange	(0,1,0)	$\pi/2$	$\pi/2$	0	0
bottom	(0, 0, 1)	0	0	0	0
up	$e^{i \arg(V_{ub})}(V_{ud}^*, V_{us}^*, V_{ub}^*)$	0.23	1.57	-1.17	-1.17
charm	$e^{i \arg(V_{cb})}(V_{cd}^*, V_{cs}^*, V_{cb}^*)$	1.80	1.53	-6.2×10^{-4}	-3.3×10^{-5}
top	$e^{i \arg(V_{tb})}(V_{td}^*, V_{ts}^*, V_{tb}^*)$	4.92	0.042	-0.018	0.39

The misalignment between down- and up-quarks is described by the CKM matrix.

Gherardi, DM, Nardecchia, Romanino 1903.10954 DM, Nardecchia, Stanzione, Toni 2404.06533



$$\{q_L^i\}$$
 space, neglecting phases

$$\mathcal{L}_{\mathrm{LEFT}}^{\mathrm{NP}} = \sum_{ij\alpha\beta} \left[L_L^{ij\alpha\beta} (\bar{d}_L^i \gamma_\mu d_L^j) (\bar{\nu}_L^\alpha \gamma^\mu \nu_L^\beta) + L_R^{ij\alpha\beta} (\bar{d}_R^i \gamma_\mu d_R^j) (\bar{\nu}_L^\alpha \gamma^\mu \nu_L^\beta) \right]$$

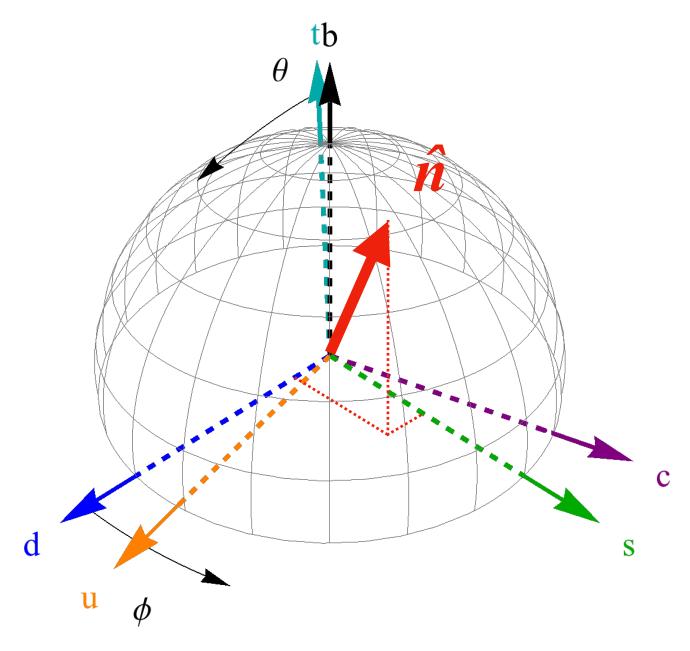
We assume that New Physics is aligned to a specific direction \hat{n} .

> the EFT coefficients are given by an overall scale times the projection of \hat{n} on the specific flavour direction

$$L^{ijvv} = C \hat{n}_i \hat{n}_j^*$$

$$\hat{n} = \begin{pmatrix} \sin \theta \cos \phi e^{i\alpha_{bd}} \\ \sin \theta \sin \phi e^{i\alpha_{bs}} \\ \cos \theta \end{pmatrix}$$

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 $\{q_L^i\}$ space, neglecting phases

$$\mathcal{L}_{\text{LEFT}}^{\text{NP}} = \sum_{ij\alpha\beta} \left[L_L^{ij\alpha\beta} (\bar{d}_L^i \gamma_\mu d_L^j) (\bar{\nu}_L^\alpha \gamma^\mu \nu_L^\beta) + L_R^{ij\alpha\beta} (\bar{d}_R^i \gamma_\mu d_R^j) (\bar{\nu}_L^\alpha \gamma^\mu \nu_L^\beta) \right]$$

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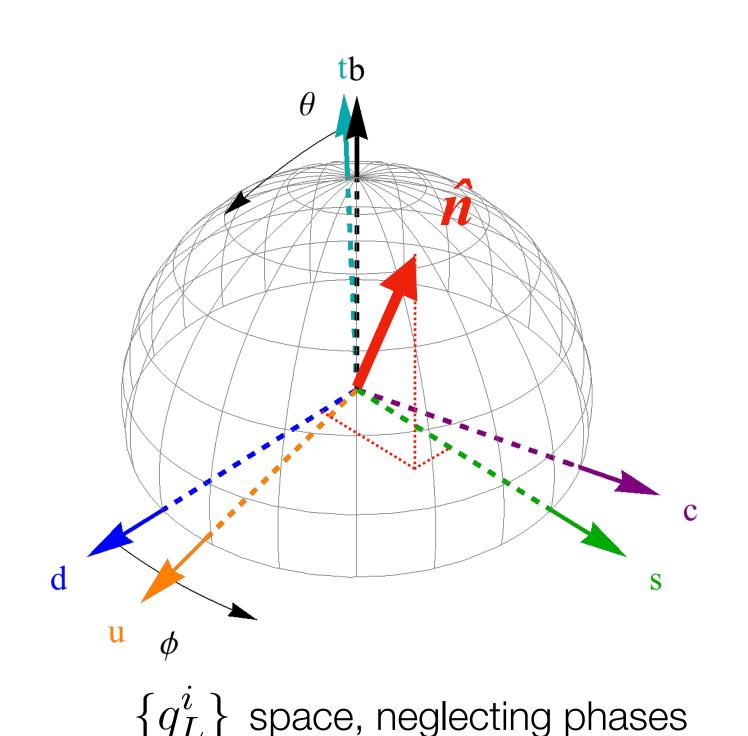
This structure is automatic if New Physics couples linearly to a single combination of quarks:

$$\mathcal{L} \supset \lambda_i \bar{q}^i \mathcal{O}_{\mathrm{NP}} + \mathrm{h.c.}$$

e.g.

- leptoquarks coupled mainly to 1 lepton family
- Vector coupled via the mixing of a single vector-like quark

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$$\mathcal{L}_{\mathrm{LEFT}}^{\mathrm{NP}} = \sum_{ij\alpha\beta} \left[L_L^{ij\alpha\beta} (\bar{d}_L^i \gamma_\mu d_L^j) (\bar{\nu}_L^\alpha \gamma^\mu \nu_L^\beta) + L_R^{ij\alpha\beta} (\bar{d}_R^i \gamma_\mu d_R^j) (\bar{\nu}_L^\alpha \gamma^\mu \nu_L^\beta) \right]$$

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At any value of (φ,θ) we can fix the overall scale

C by imposing the **best-fit** of $B \rightarrow K(*) vv$.

$$\frac{\text{sbvv}}{\text{cos}\,\theta} = \left(\frac{8\text{TeV}}{\text{cos}\,\theta}\right)^{-2}$$

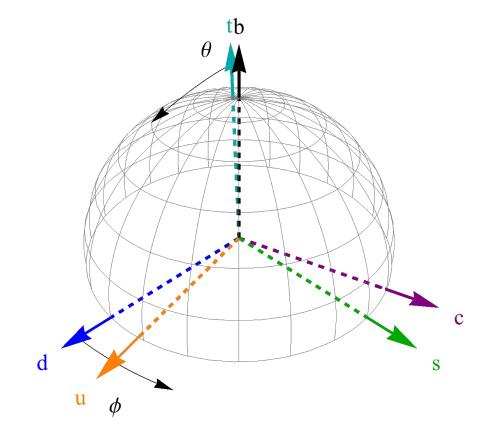
dbs = dbd =0

(fit in backup slides)

$$L^{ijvv} = C \hat{n}_i \hat{n}_j^*$$

$$L = C \hat{N}_i \hat{N}_i^*$$

$$= C \cos \theta \sin \theta \sin \phi = (8 \text{ TeV})^{-2}$$



Once C is fixed as function of (θ, ϕ) , all parameters are set and we can check the

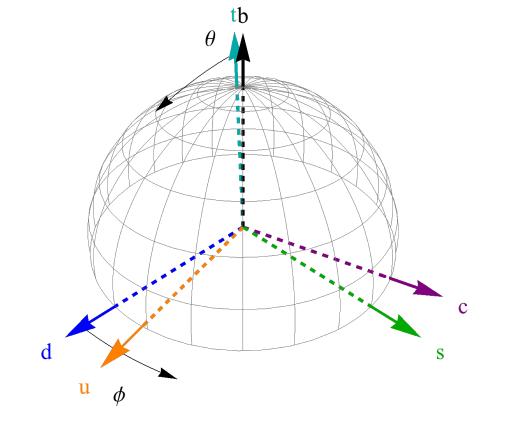
constraints from other observables

2404.06533

$$L^{ijvv} = C \hat{n}_i \hat{n}_j^*$$

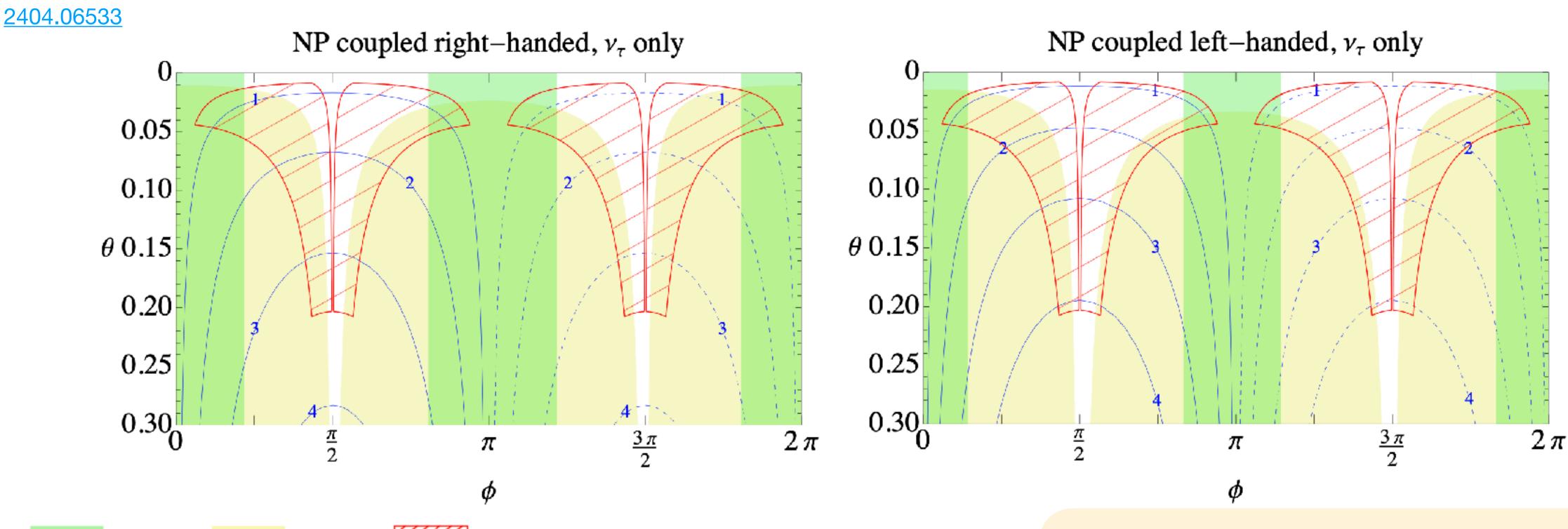
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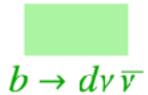
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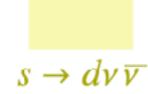


Once C is fixed as function of (θ, ϕ) , all parameters are set and we can check the

constraints from other observables









$$|\mathbf{C}|^{-1/2}$$
 [TeV]

The allowed region (white) is close to the third generation, with a misalignment of O(CKM).

Is there a larger picture emerging from data?

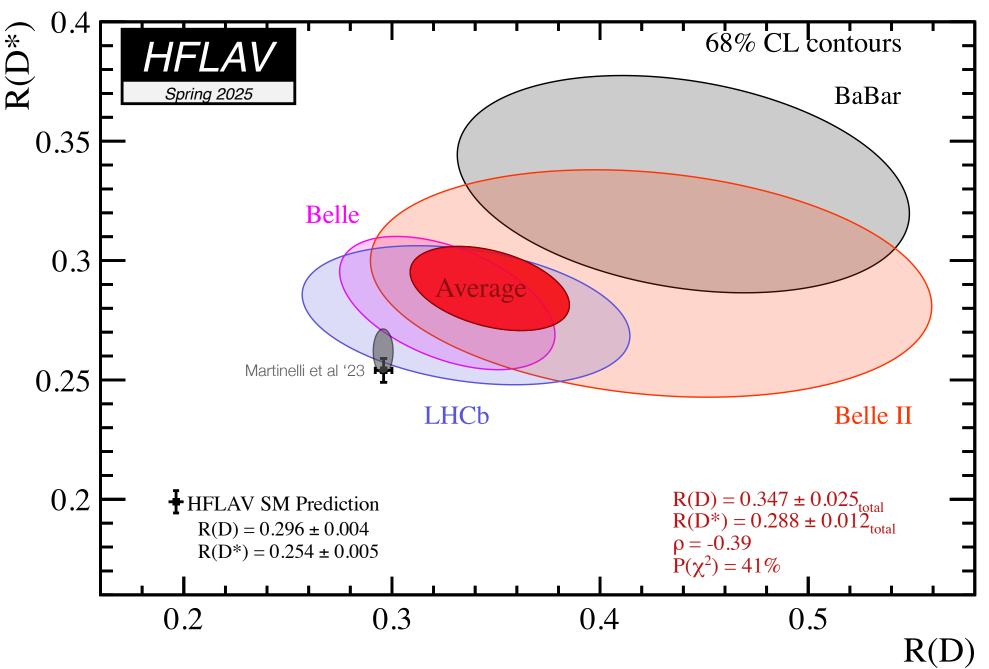
B-anomalies in charged current

$b \rightarrow c \, \tau \, \overline{\nu}_{\tau}$

Lepton Flavour Universality

$$R(D^{(*)}) \equiv \frac{\mathcal{B}(B^0 \to D^{(*)} + \tau \nu)}{\mathcal{B}(B^0 \to D^{(*)} + \ell \nu)}, \quad R(X) = \frac{\mathcal{B}(B \to X \tau \nu_{\tau})}{\mathcal{B}(B \to X \ell \nu_{\ell})}$$

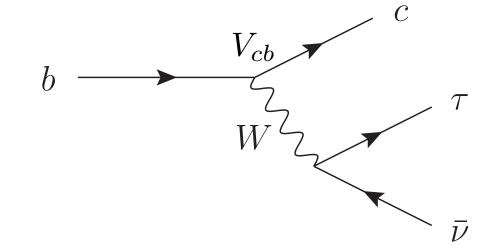
$$\ell = \mu, e$$



Most recent measurement by Belle-II

confirmed the tension: $3 - 4\sigma$.

Tree-level SM process with V_{cb} suppression.



SM prediction under control for R(D), less so for R(D*), related to Vcb incl/excl tension.

Martinelli et al. '23, '24

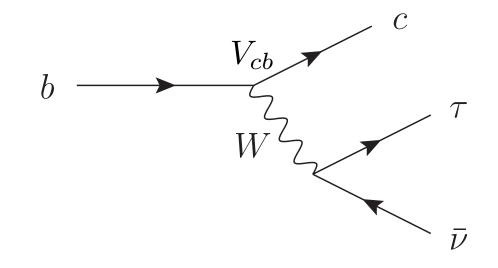
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$b \rightarrow c \tau \overline{\nu}_{\tau}$

Lepton Flavour Universality

$$R(D^{(*)}) \equiv \frac{\mathcal{B}(B^0 \to D^{(*)+}\tau\nu)}{\mathcal{B}(B^0 \to D^{(*)+}\ell\nu)}, \quad R(X) = \frac{\mathcal{B}(B \to X\tau\nu_{\tau})}{\mathcal{B}(B \to X\ell\nu_{\ell})}$$
$$\ell = \mu, e$$

Tree-level SM process with V_{cb} suppression.



SM prediction under control for R(D), less so for R(D*), related to Vcb incl/excl tension.

Martinelli et al. '23, '24

$$\mathcal{L}_{EFT} > C_{ijte}^{duve} \left(\bar{d}_{iL} \, \mathcal{V}_{\mu} \, u_{jL} \right) \left(\bar{\nu}_{e} \, \mathcal{V}^{r} \, \mathcal{V}_{L} \right)$$

Corresponds to a New Physics scale of

$$\Lambda_{bc\tau v} \sim 4 \text{ TeV}$$

0.4

##FLAV
Spring 2025

Belle

0.35

LHCb

Belle II

0.25

HFLAV SM Prediction
R(D) = 0.296 ± 0.004
R(D*) = 0.296 ± 0.004
R(D*) = 0.254 ± 0.005
P(x²) = 41%

0.2

0.36

0.4

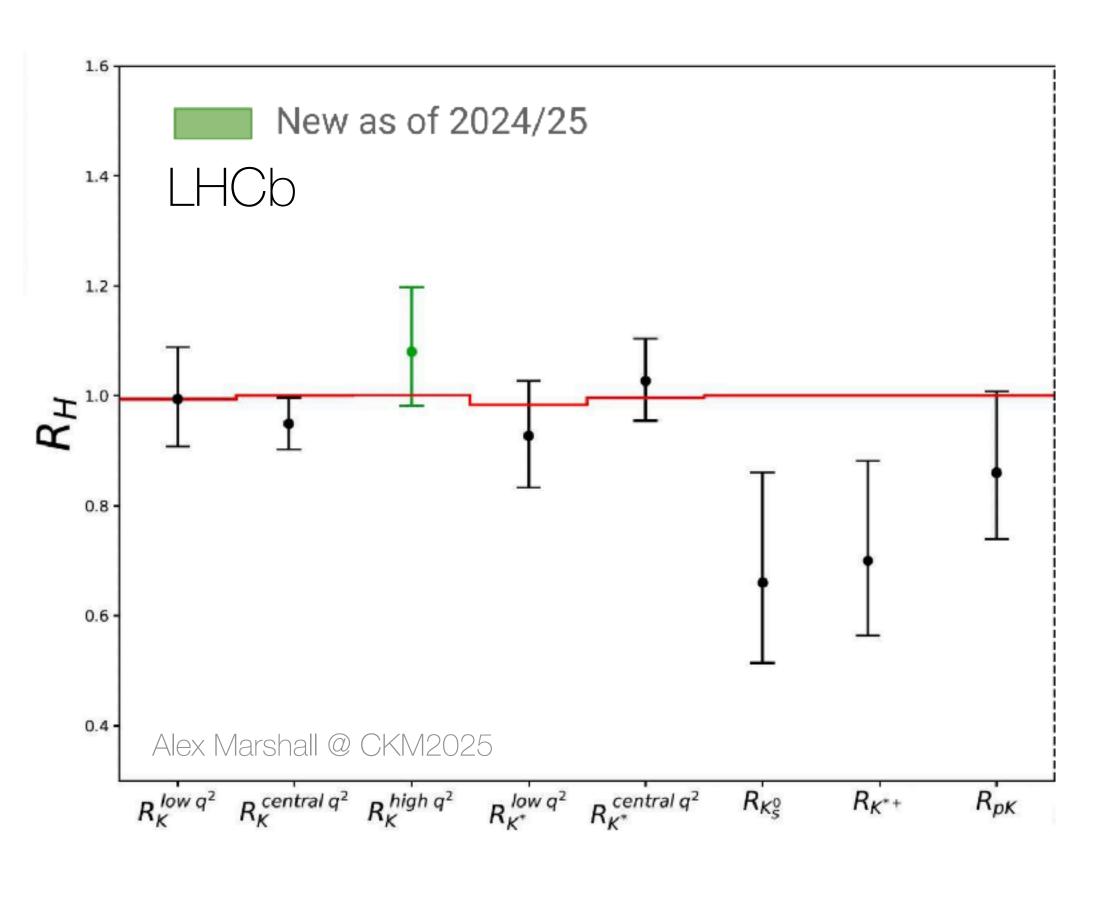
0.5

Most recent measurement by Belle-II confirmed the tension: $3 - 4\sigma$.

We eagerly wait for more data by Belle-II and LHCb.

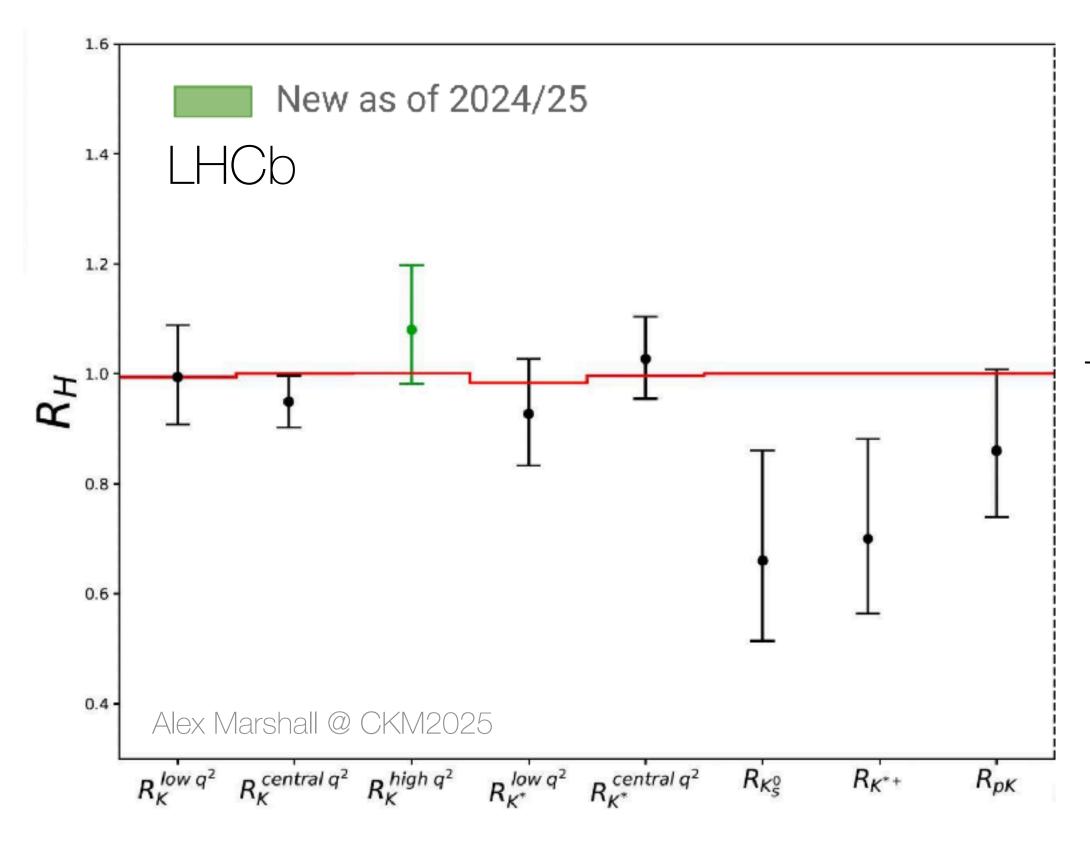
SM predictions will take advantage of larger and more precise datasets!

$$b \rightarrow s \mu^+ \mu^-/b \rightarrow s e^+ e^-: R(K^{(*)})$$



Clean SM prediction ($R_X = 1$), test of LFU between μ and e. μ vs. e LFU established at ~5% level.

$$b \rightarrow s \mu^+ \mu^-/b \rightarrow s e^+ e^-: R(K^{(*)})$$



Clean SM prediction ($R_X = 1$), test of LFU between μ and e.

μ vs. e LFU established at ~5% level.

To which NP scale Λ are these measurements sensitive to?

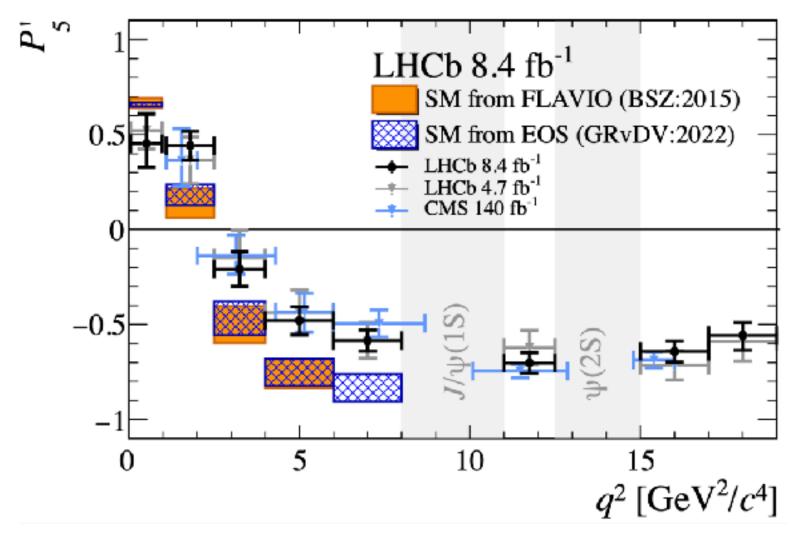
Take this *current* x *current* LFUV operator as example

$$\mathcal{L}_{CFT} > \frac{C}{N^2} \left(\overline{b}_{\lambda} \times_{\lambda} S_{\lambda} \right) \left(\overline{\mu}_{\lambda} \times^{\delta} \mu_{\lambda} \right)$$

if
$$c = 1$$
: $\Lambda_{bs\mu\mu} \gtrsim 56 \text{ TeV}$

Lower scales require same couplings to electrons and muons.

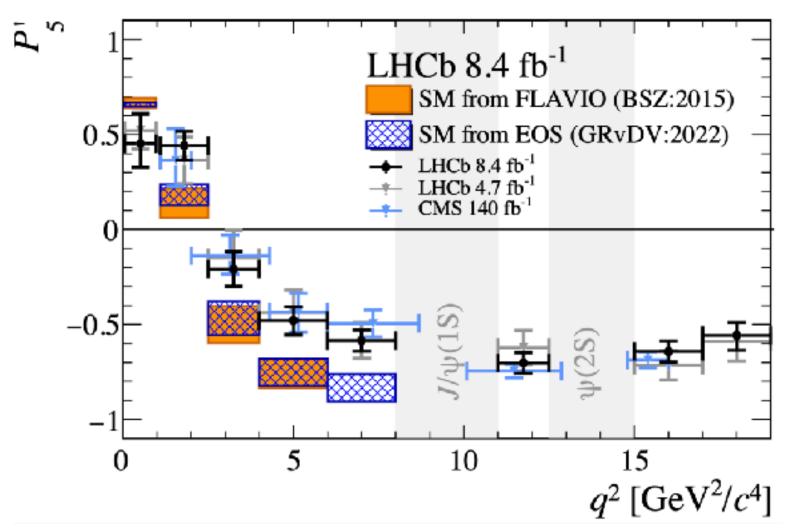
LHCb Run1 + Run2



$b \rightarrow s \mu^+ \mu^-$: P₅' and Br's

Very significant tension (>4 σ) between data and SM prediction in angular observables and Br's of $b \rightarrow s\mu^+\mu^-$ transitions.

LHCb Run1 + Run2



$$b \rightarrow s \mu^+ \mu^-$$
: P₅' and Br's

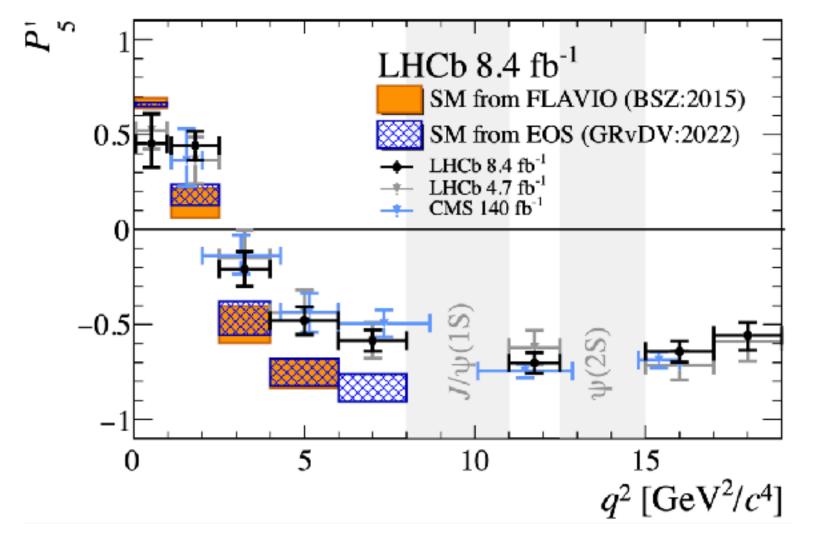
Very significant tension (>4 σ) between data and SM prediction in angular observables and Br's of $b \rightarrow s\mu^+\mu^-$ transitions.

If it is due to New Physics, it must respect LFU (to give R(K)=1)

$$\mathcal{L}_{CFT} > \frac{c}{\Lambda^2} \left(\bar{b}_{L} \chi_{A} \zeta_{L} \right) \left[\left(\bar{\mu}_{L} \chi^{A} \mu_{L} \right) + \left(\bar{e}_{L} \chi^{A} e_{L} \right) \right] \qquad \text{if } c = 1:$$

$$\Lambda_{bs\ell\ell} \sim 40 \text{ TeV}$$

LHCb Run1 + Run2



$b \rightarrow s \mu^+ \mu^-$: P₅' and Br's

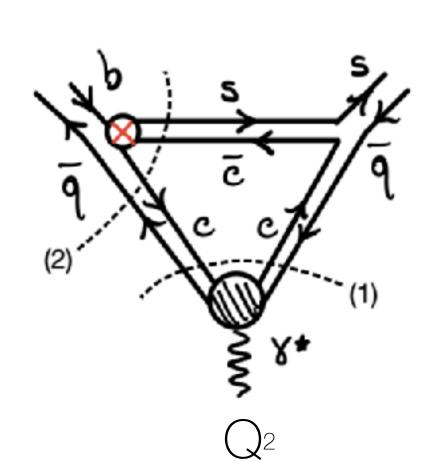
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$$\Lambda_{\text{bsll}} \sim 40 \text{ TeV}$$

However, non-perturbative long-distance QCD dynamics could reproduce the same effect.



Charm-rescattering: effects not accounted for in the SM predictions above.

Ciuchini et al. 2212.10516

Model estimates based on HHChPT estimate impact at 5% to 20% of short-distance. Isidori et al. 2405.1755, 2507.17824

Recent progress towards a lattice calculation! Rome group 2508.03655

More data will help in clarifying: allows for check of Q² dependence of the result and more detailed studies.

$$b \rightarrow s \mu^+ \mu^-$$
: P₅' and Br's

$$\mathcal{L}_{\text{CFT}} > \frac{c}{\Lambda^{2}} \left(\int_{\lambda} \mathcal{L}_{\lambda} \mathcal{L}_{\lambda} \right) \left[\left(\bar{\mu}_{L} \mathcal{L}^{A} \mu_{L} \right) + \left(\bar{e}_{L} \mathcal{L}^{A} e_{L} \right) \right] \quad \text{if } c = 1:$$

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$$b \rightarrow s \mu^+ \mu^-$$
: P5' and Br's

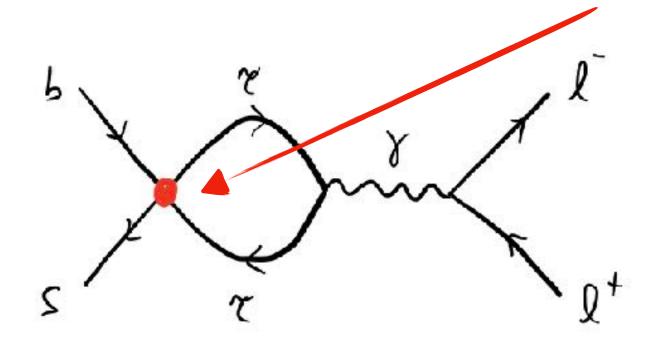
$$\mathcal{L}_{CFT} > \frac{c}{\Lambda^2} \left(\int_{\mathcal{L}} \mathcal{L}_{\lambda} \mathcal{L}_{\lambda} \right) \left[\left(\bar{\mu}_{L} \mathcal{L}^{\alpha} \mu_{L} \right) + \left(\bar{e}_{L} \mathcal{L}^{\alpha} e_{L} \right) \right] \quad \text{if } c = 1:$$

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An interesting New Physics contribution

Bobeth et al. 1109.1826, Capdevila et al. 1712.01919, Crivellin et al. 1807.02068, Alguerò et al. 1903.09578, Cornella et al. 2001.04470, Aebischer, Isidori, et al. 2210.13422,

$$(\overline{b}_L \gamma^{\mu} c_L)(\overline{v}_L \gamma^{\mu} \tau_L) \longleftrightarrow (\overline{b}_L \gamma^{\mu} s_L)(\overline{\tau}_L \gamma^{\mu} \tau_L)$$



$$C_9^{\text{U}} \approx 7.5 \left(1 - \sqrt{\frac{R_{D^{(*)}}}{R_{D^{(*)}\text{SM}}}} \right) \left(1 + \frac{\log(\Lambda^2/(1\text{TeV}^2))}{10.5} \right)$$

- → Related to R(D(*))
 → Induce C₉^U, R(K)=1

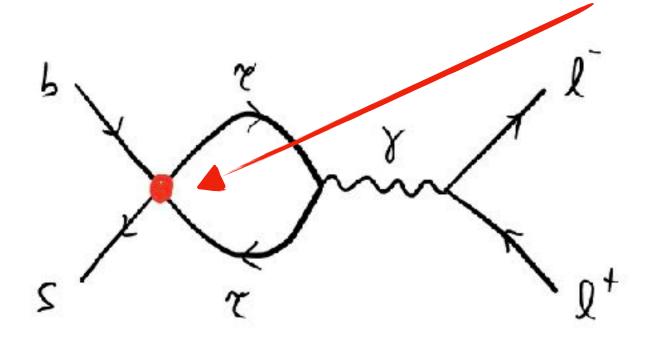
$$\Lambda_{bs\tau\tau} \sim O(4) \text{ TeV}$$

 $b \rightarrow s \mu^+ \mu^-$: P₅' and Br's

An interesting New Physics contribution

Bobeth et al. 1109.1826, Capdevila et al. 1712.01919, Crivellin et al. 1807.02068, Alguerò et al. 1903.09578, Cornella et al. 2001.04470, Aebischer, Isidori, et al. 2210.13422,

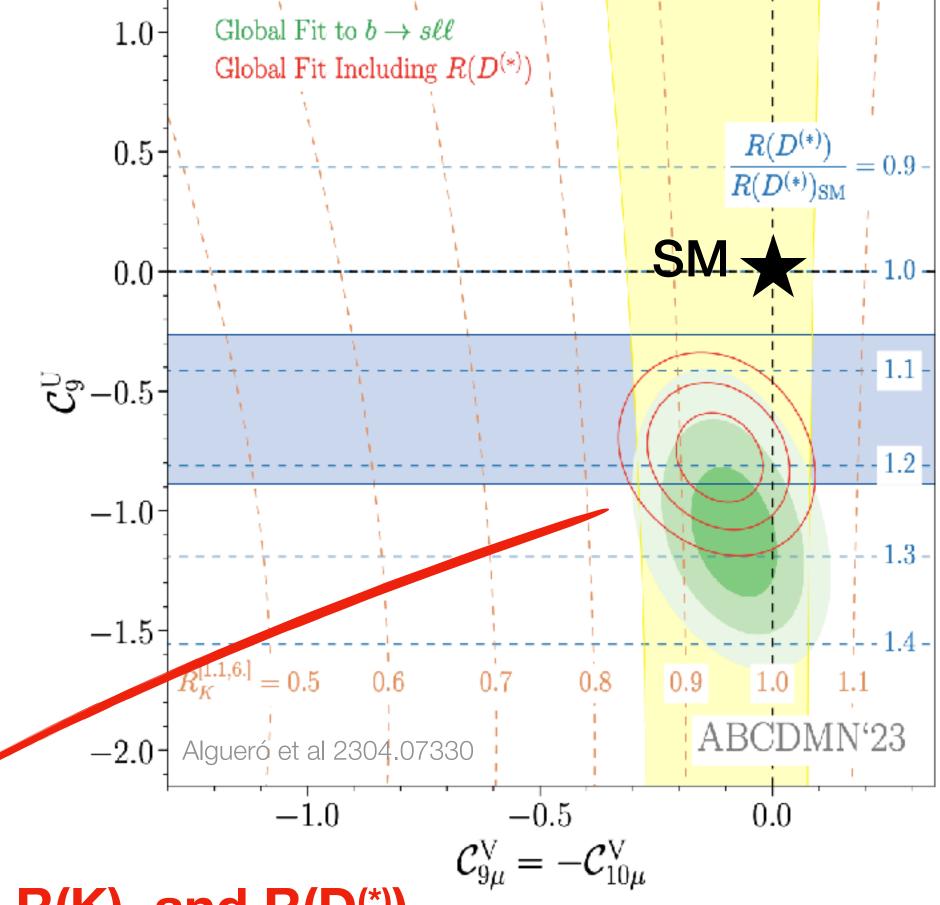
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$$\Lambda_{\rm bs\tau\tau} \sim {\rm O}(4)~{
m TeV}$$

$$C_9^{
m U} \approx 7.5 \left(1 - \sqrt{\frac{R_{D^{(*)}}}{R_{D^{(*)}{
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m TeV}^2))}{10.5}\right)$$

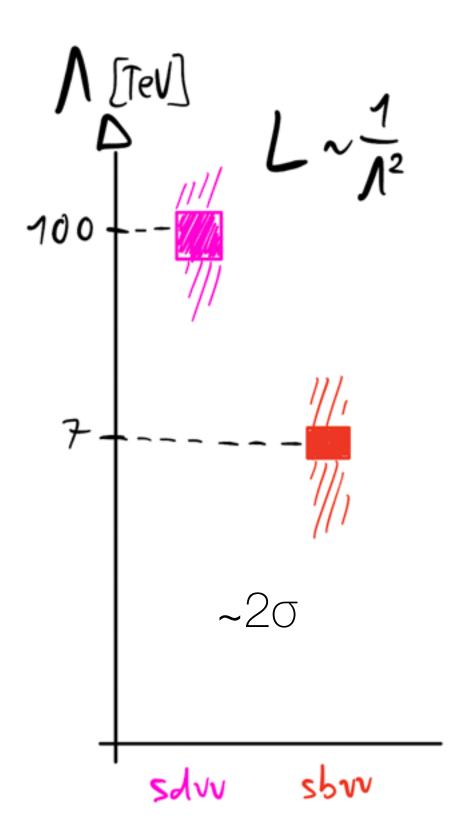
- → Related to R(D(*))
- → Induce C_9^U , R(K)=1



Compatible fit between b→sll, R(K), and R(D(*)).

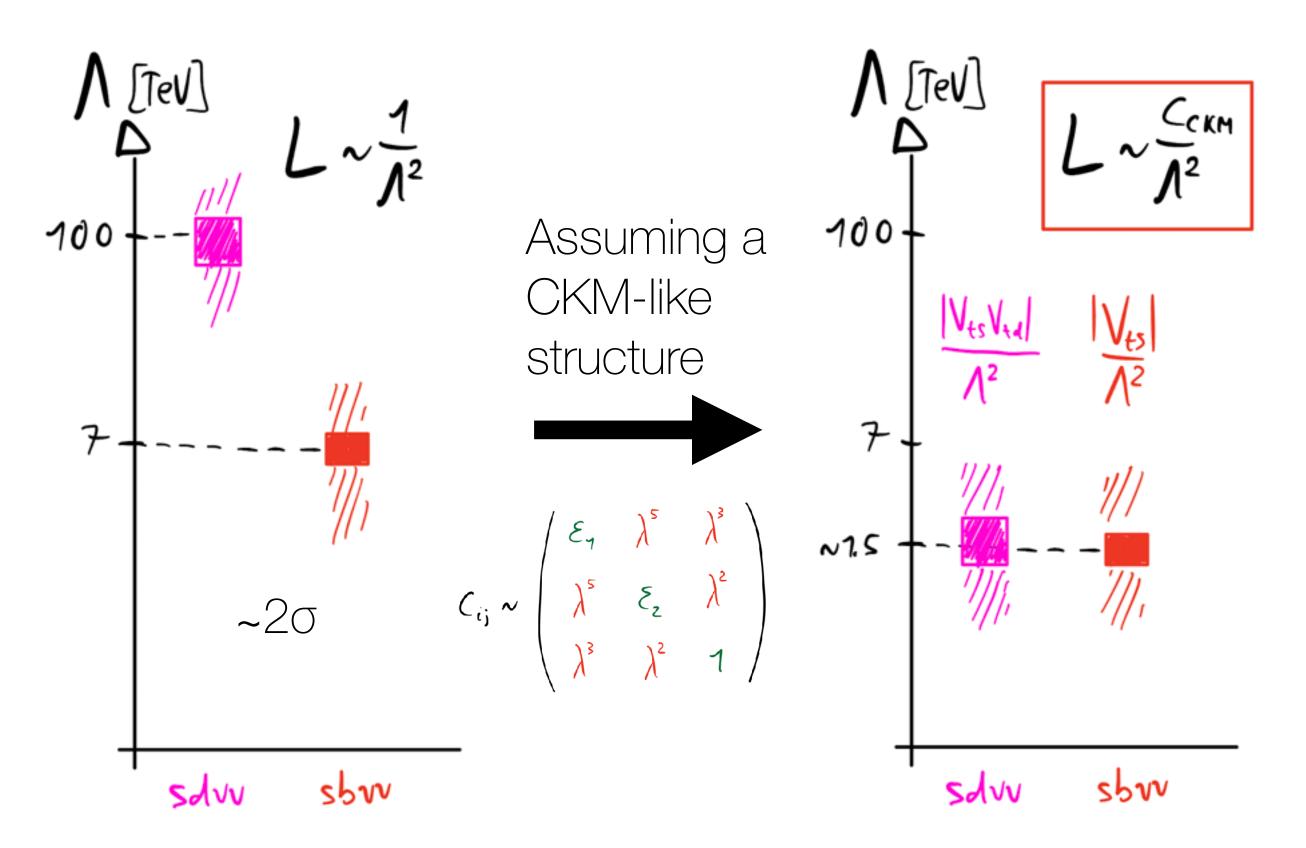
Neutral-current

$$\mathcal{L}_{EFT} > \left[\frac{ijtr}{L_{i,R}} \left(\bar{d}_{iL,R} \chi_{r} d_{jL,R} \right) \left(\bar{\nu}_{r} \chi^{r} \nu_{r} \right) \right]$$



Neutral-current

$$\mathcal{L}_{EFT} > \left[\frac{ijtr}{L_{i,R}} \left(\bar{d}_{iL,R} \gamma_{\mu} d_{jL,R} \right) \left(\bar{\nu}_{\tau} \gamma^{\mu} \nu_{\tau} \right) \right]$$



The physics scales become compatible!

Neutral-current Charged-current

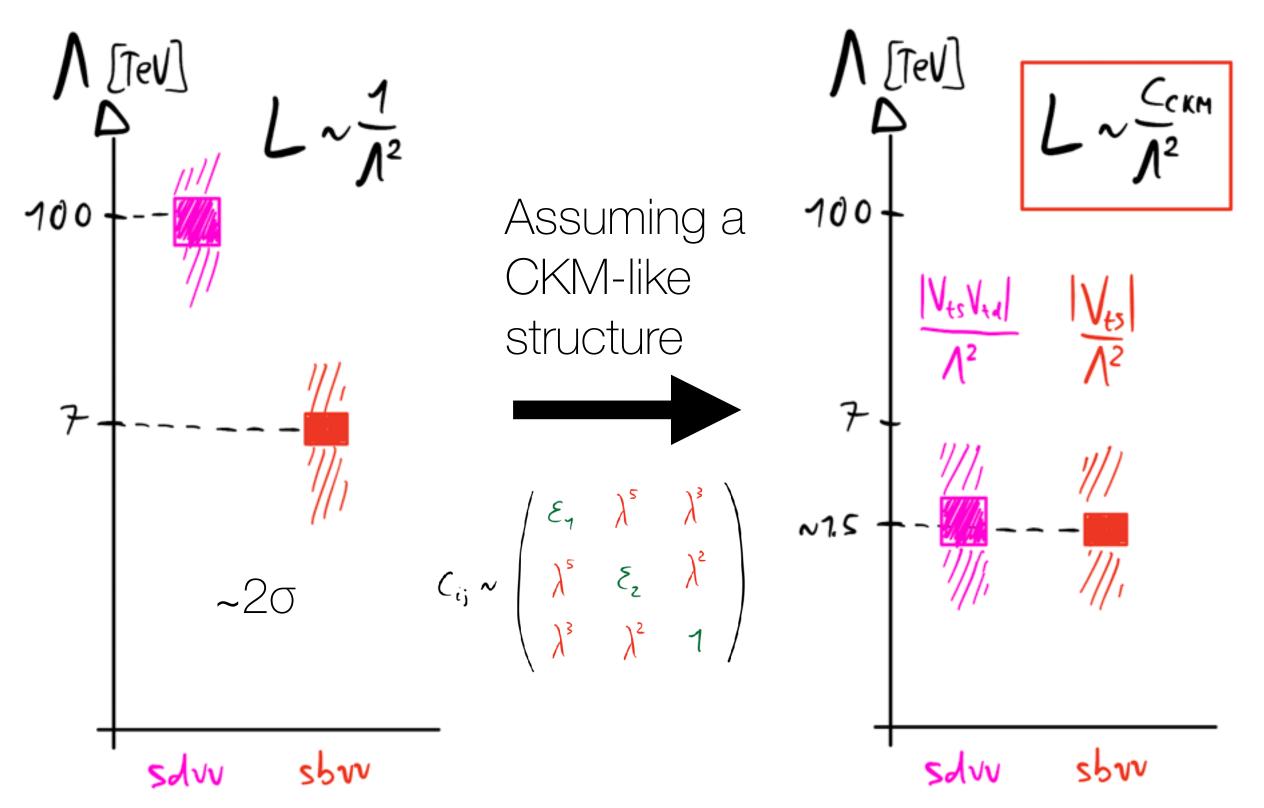
Neutral-current

$$\mathcal{L}_{EFT} > \bigcup_{L,R}^{ijtr} \left(\overline{d}_{iL,R} \mathcal{V}_{\mu} d_{jL,R} \right) \left(\overline{\mathcal{V}}_{\tau} \mathcal{V}^{r} \mathcal{V}_{\tau} \right) \xrightarrow{SU(2)_{L}} \mathcal{L}_{EFT} > \bigcup_{ijtr}^{cc} \left(\overline{d}_{iL} \mathcal{V}_{\mu} u_{jL} \right) \left(\overline{\mathcal{V}}_{\tau} \mathcal{V}^{r} \mathcal{V}_{L} \right)$$
The precise correlation is model-dependent

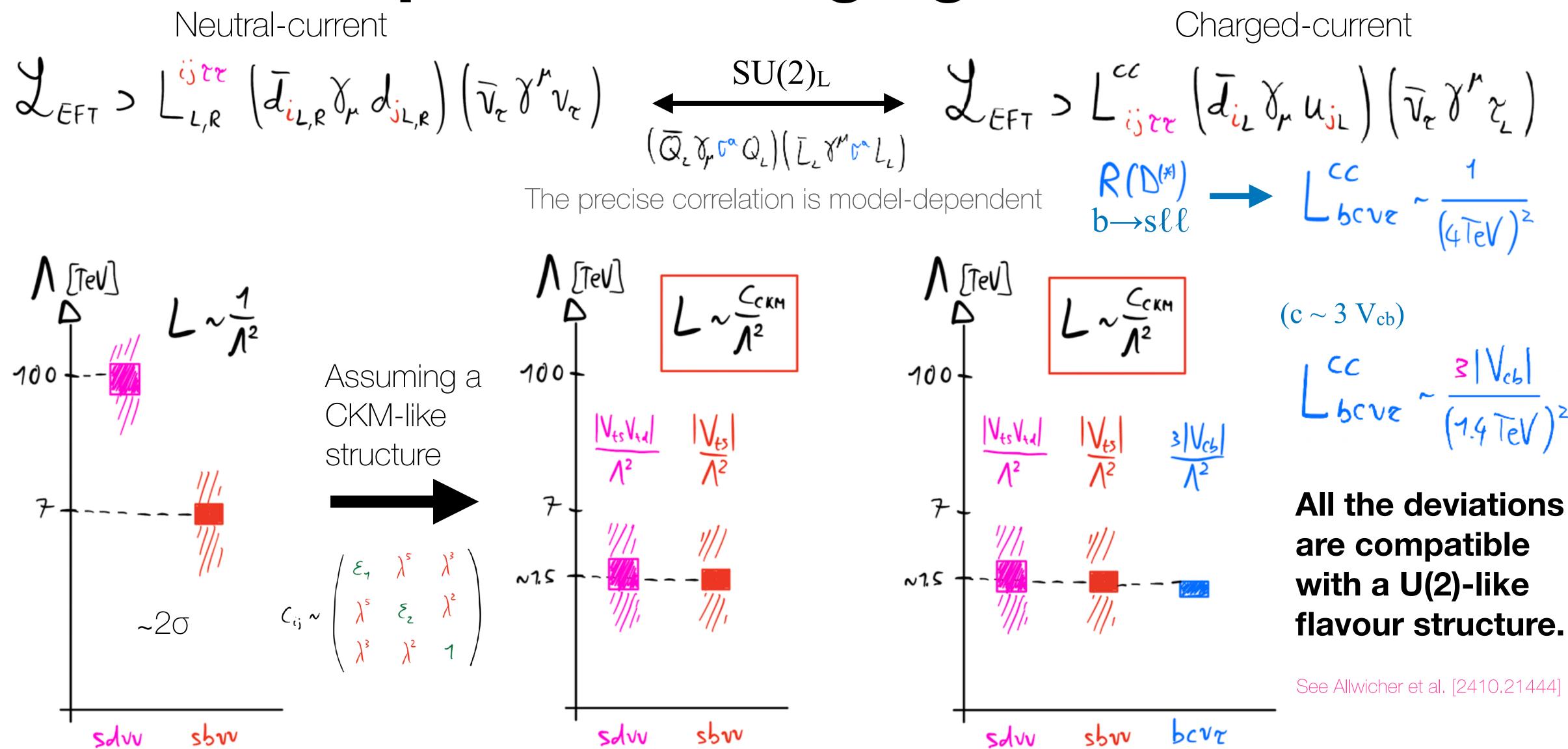
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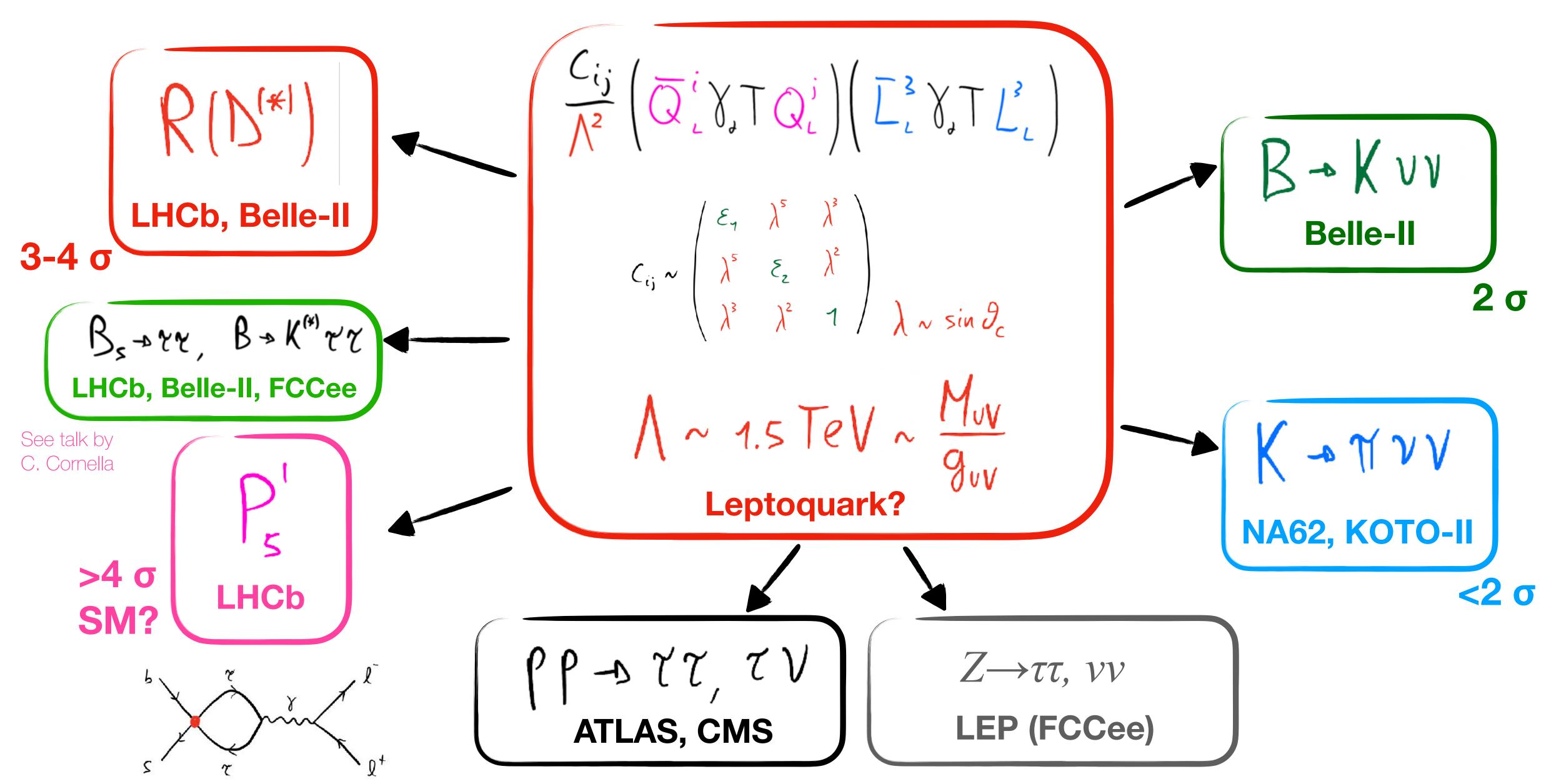
$$\mathcal{L}_{bcve} \sim \frac{1}{(4 \text{ TeV})^{2}}$$



The physics scales become compatible!



The physics scales become compatible!



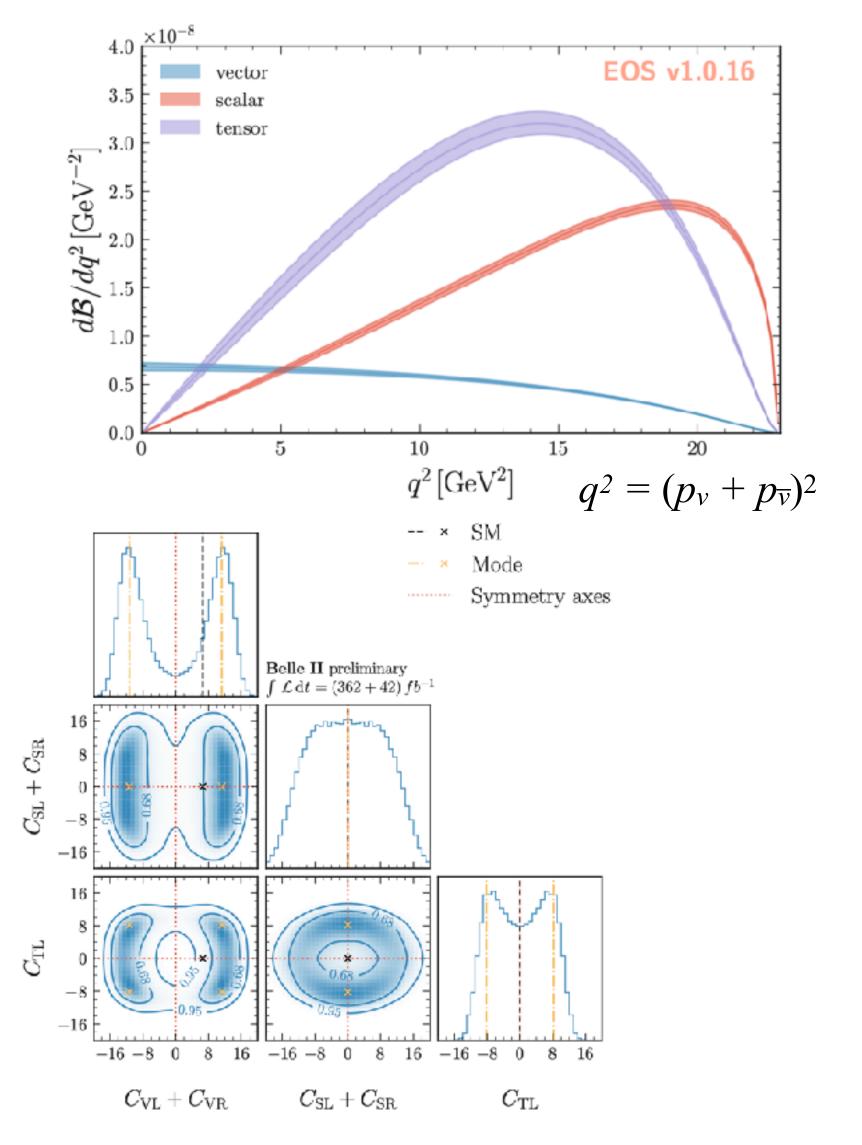
Part III

Exploring the exotic frontiers

(but not all theories are created equal)

Reinterpretation framework of $B \rightarrow K^+ \nu \nu$, generalising the EFT beyond the d=6 SMEFT:

Gartner at al. 2402.08417, Belle-II 2507.12393



$$\mathcal{L}^{\text{WET}} = -\frac{4G_{\text{F}}}{\sqrt{2}} \frac{\alpha}{2\pi} V_{ts}^* V_{tb} \sum_{i} C_{i}(\mu_{b}) O_{i} + \text{h.c.}$$

$$\mathcal{O}_{\text{VL}} = (\overline{\nu_{L}} \gamma_{\mu} \nu_{L}) (\overline{s_{L}} \gamma^{\mu} b_{L})$$

$$\mathcal{O}_{\text{VR}} = (\overline{\nu_{L}} \gamma_{\mu} \nu_{L}) (\overline{s_{R}} \gamma^{\mu} b_{R})$$

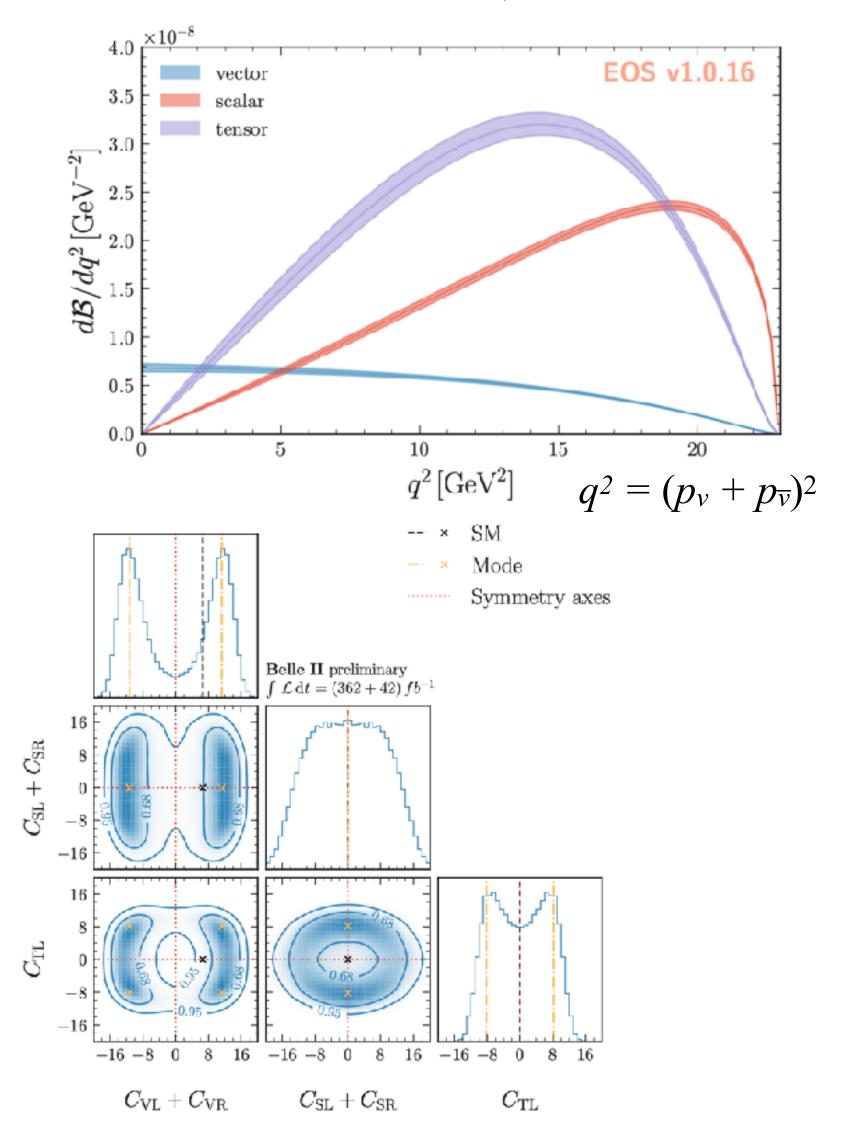
$$\mathcal{O}_{\text{SL}} = (\overline{\nu_{L}^{c}} \nu_{L}) (\overline{s_{R}} b_{L})$$

$$\mathcal{O}_{\text{SR}} = (\overline{\nu_{L}^{c}} \nu_{L}) (\overline{s_{L}} b_{R})$$

$$\mathcal{O}_{\text{TL}} = (\overline{\nu_{L}^{c}} \sigma_{\mu\nu} \nu_{L}) (\overline{s_{R}} \sigma^{\mu\nu} b_{L})$$

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$$-0. Oiq^{(1)}, Oia$$

$$\Delta L = 2$$

Lepton number-violating operators Generated at d=7 in SMEFT

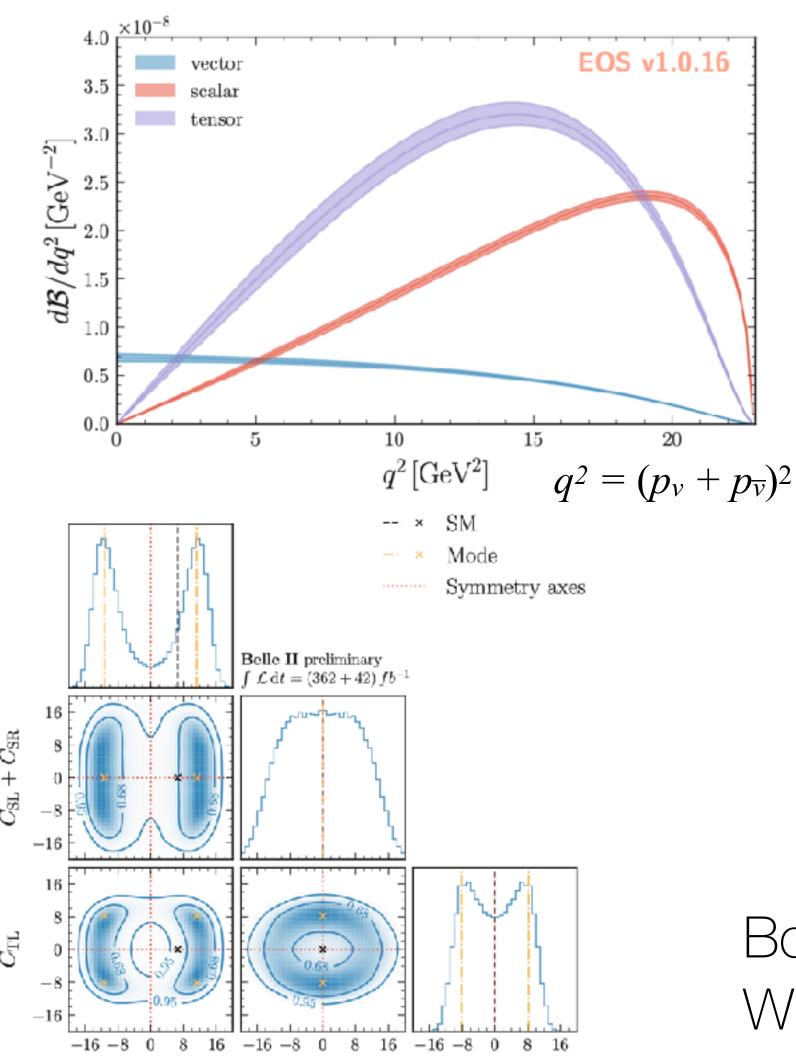
Fridell et al. 2306.08709

$$\frac{C^{(3)}}{\Lambda^3} (\overline{d}_R L_L) (\overline{Q}_L^c L_L) H$$

If at d=6 the EFT scale required was $\Lambda^{(6)} \sim 7 \text{ TeV}$, at d=7 it becomes $\Lambda^{(7)} \sim 2$ TeV.

Reinterpretation framework of $B \rightarrow K^+ \nu \nu$, generalising the EFT beyond the d=6 SMEFT:

Gartner at al. 2402.08417, Belle-II 2507.12393



 $C_{
m VL}+C_{
m VR}$

 $C_{\mathrm{SL}}+C_{\mathrm{SR}}$

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d=6:
$$O_{lq}^{(1,3)}$$
, O_{ld}

Lepton number-violating operators Generated at d=7 in SMEFT

Fridell et al. 2306.08709

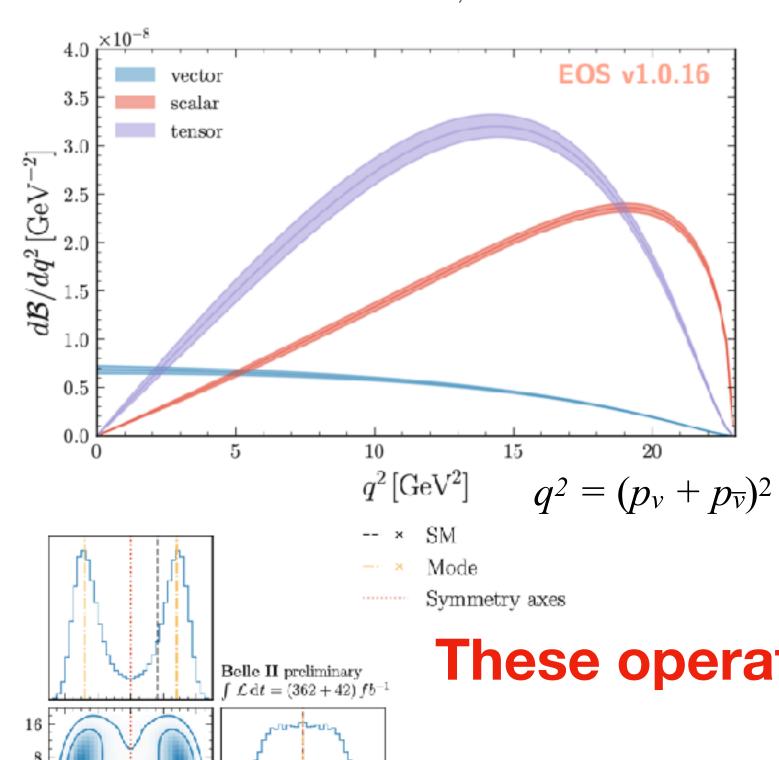
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Bounds from $0\nu\beta\beta$ decay ~ 100 TeV (for down quarks). 2306.08709 Why should flavour-conserving couplings be more suppressed than violating ones?

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$$d=6: O_{lq}(1,3), O_{ld}$$

Lepton number-violating operators Generated at d=7 in SMEFT

Fridell et al. 2306.08709

$$\frac{C^{(3)}}{\Lambda^3} (\overline{J}_R L_L) (\overline{Q}_L^c L_L) H$$

These operators are on all the same footing!

If at d=6 the EFT scale required was $\Lambda^{(6)} \sim 7 \text{ TeV}$, at d=7 it becomes $\Lambda^{(7)} \sim 2$ TeV.

Bounds from $0\nu\beta\beta$ decay ~ 100 TeV (for down quarks). 2306.08709 Why should flavour-conserving couplings be more suppressed than violating ones?

Neutrino masses in the SM EFT are generated at dim=5 by the Weinberg operator

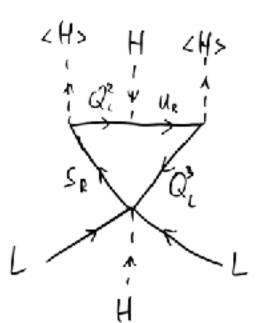
$$\Delta L = 2: \frac{c_W}{\Lambda_L} (L\tilde{H})(L\tilde{H})$$

 $\Lambda_L/c_W \sim 10^{14} \text{ GeV}$

This is the scale of breaking of lepton number L.

$$\Delta L = 2: \qquad \frac{\zeta^{(7)}}{\Lambda^{3}} (\widetilde{J}_{R} L_{L}) (\widetilde{Q}_{L}^{c} L_{L}) H \rightarrow \Lambda^{(7)}_{BKvv} \sim 2 \text{ TeV}$$

If this operator is present, with a much lower scale, there is **no symmetry argument that prevents the generation of the Weinberg operator with a too large coefficient** (too low scale). Indeed, it is induced radiatively.

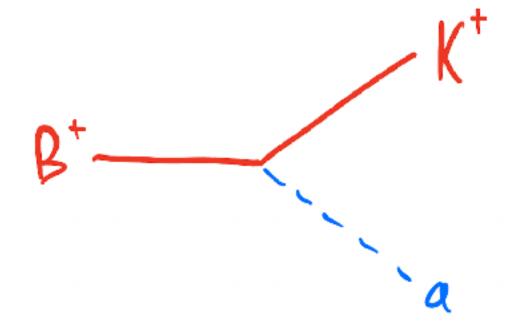


"Everything which is not forbidden is allowed"

Similar conclusions can be obtained if HNL are used instead of L-violating operators: that operator will generically induce at the radiative level too large neutrino masses.

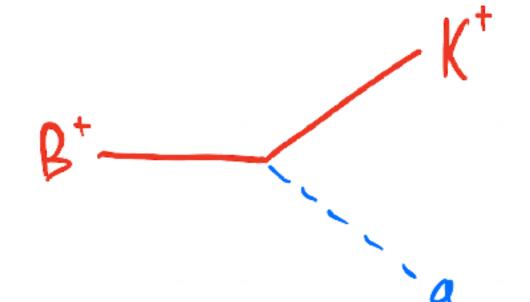
$B+ \rightarrow K+X$

A 2-body decay is only allowed into a neutral scalar (or pseudo-scalar): ALPs. In this case, the q^2 dependence is **peaked** at $q^2 = m_a^2$.



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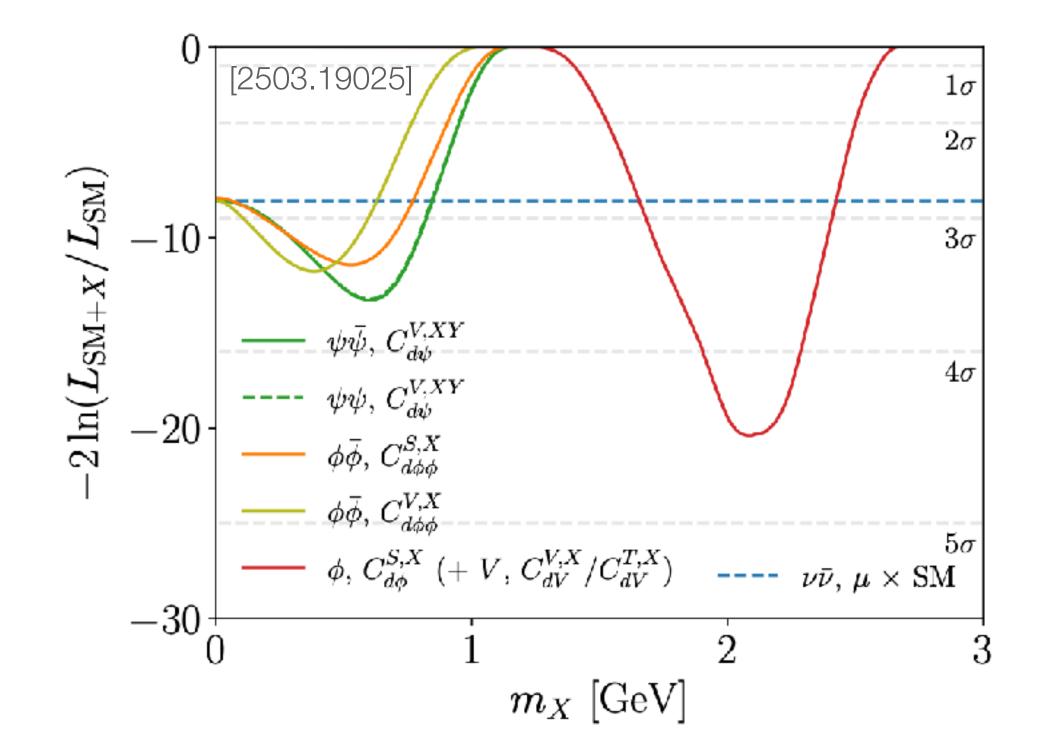


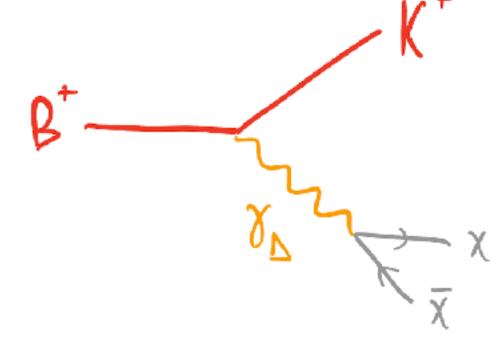
Vectors can contribute only as mediators, going then into 2 dark fermions (3-body decay).

(perhaps these could then be **DM candidates**?) e.g. Gabrielli et al. 2402.05901

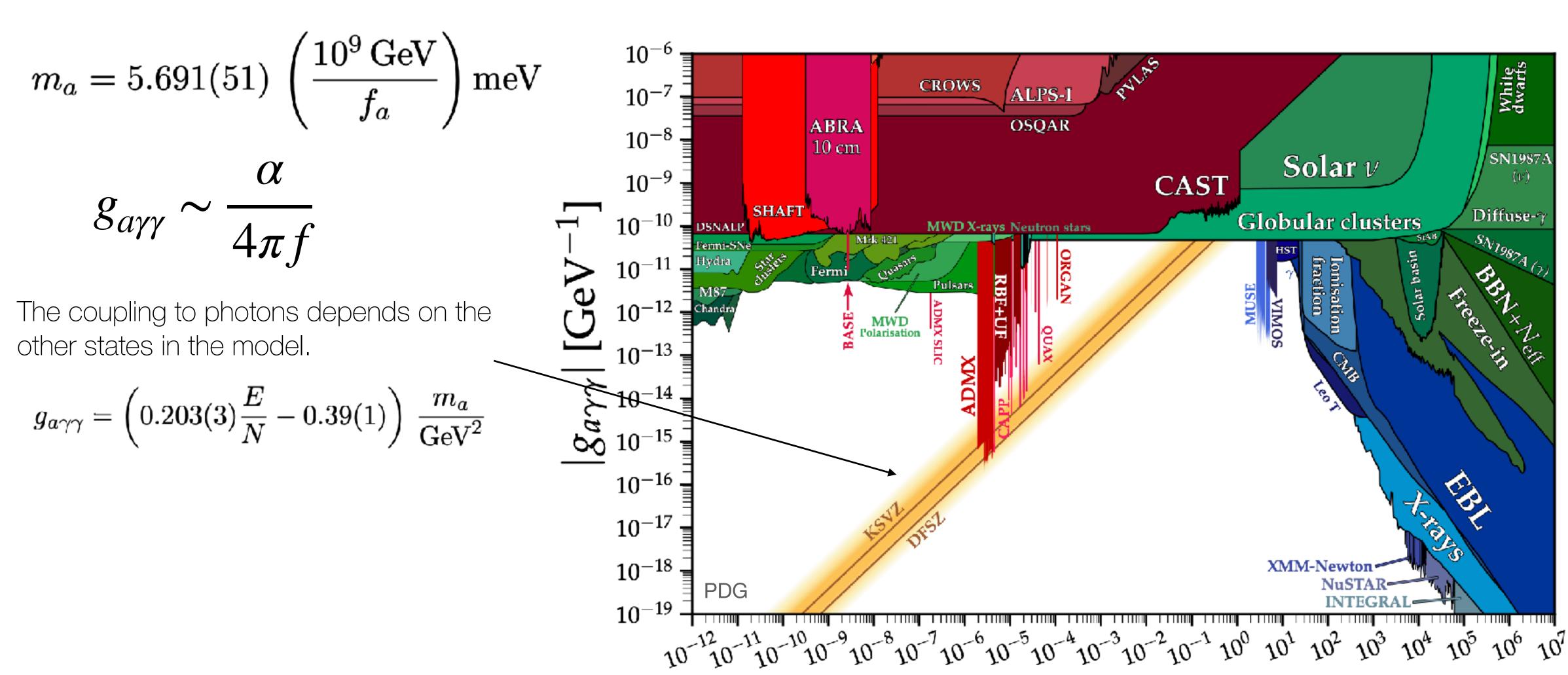
Data can differentiate between these scenarios

See e.g. Bolton et al. 2403.13887, 2503.19025





The QCD axion is the main prediction of the solution of the QCD θ -problem via a spontaneously broken $U(1)_{PQ}$ symmetry,



 m_a [eV]

The QCD axion is the main prediction of the solution of the QCD θ -problem via a spontaneously broken $U(1)_{PQ}$ symmetry,

$$m_a = 5.691(51) \left(\frac{10^9 \,\text{GeV}}{f_a}\right) \,\text{meV}$$

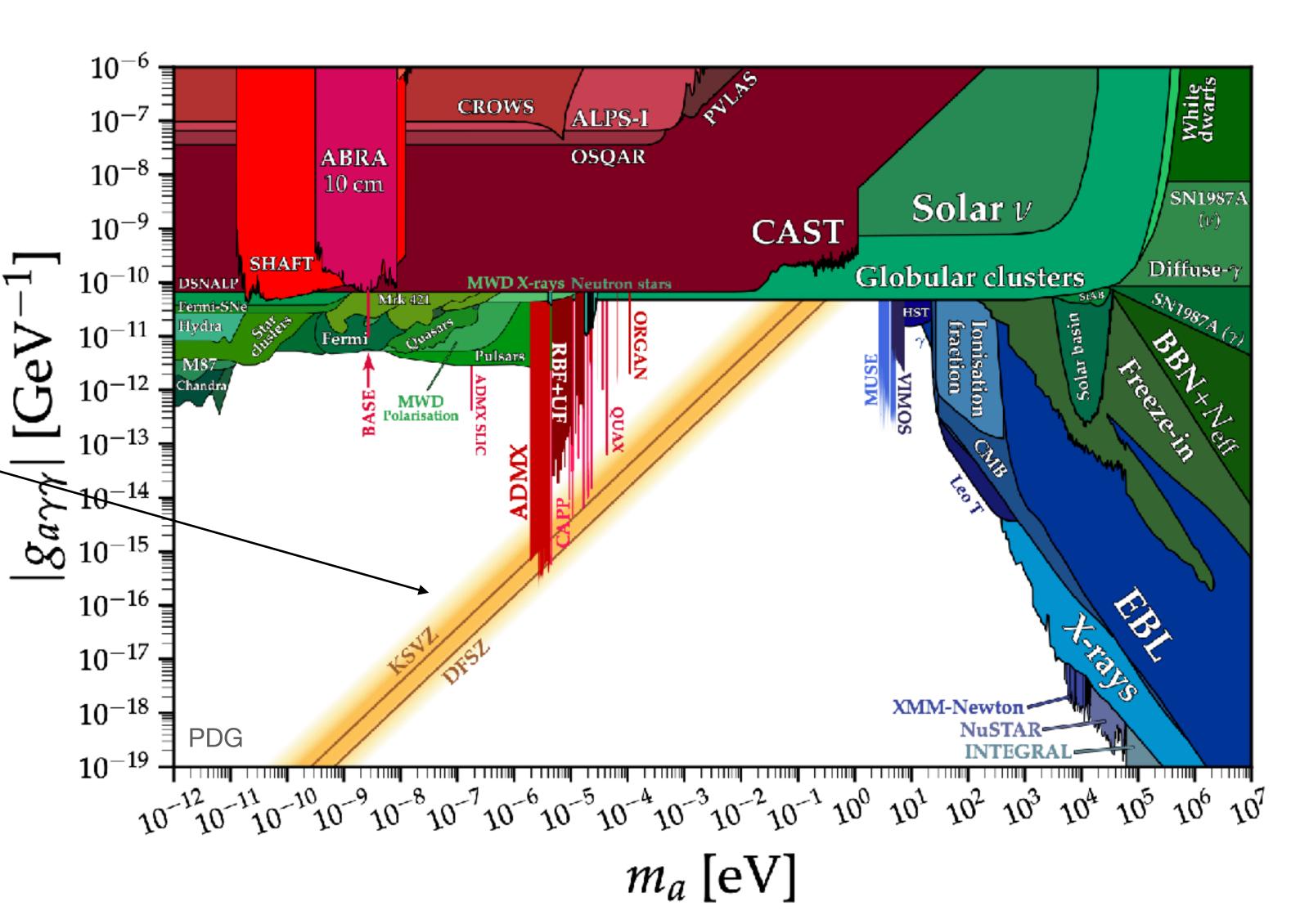
$$g_{a\gamma\gamma} \sim \frac{\alpha}{4\pi f}$$

The coupling to photons depends on the other states in the model.

$$g_{a\gamma\gamma} = \left(0.203(3)\frac{E}{N} - 0.39(1)\right) \frac{m_a}{\text{GeV}^2}$$

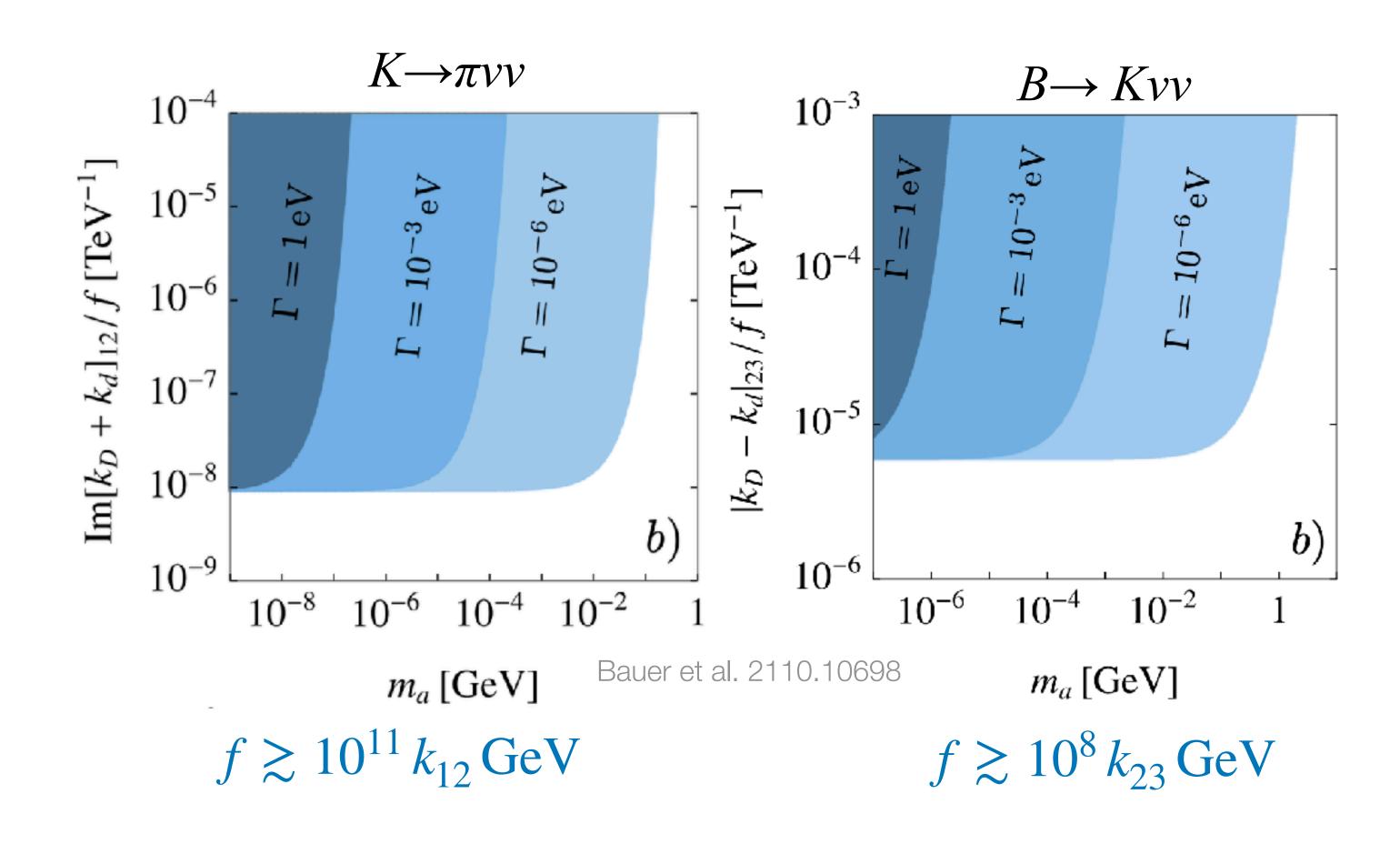
Other regions in m-g plane are not motivated by the θ -problem, they represent possible generic extensions of the SM.

The whole log-log plane can be populated by BSM theories.



The QCD axion couplings to fermions can be flavour-violating, if the PQ charges are family-dependent. This is well motivated in models where the PQ symmetry is connected with the explanation of SM Yukawa's.

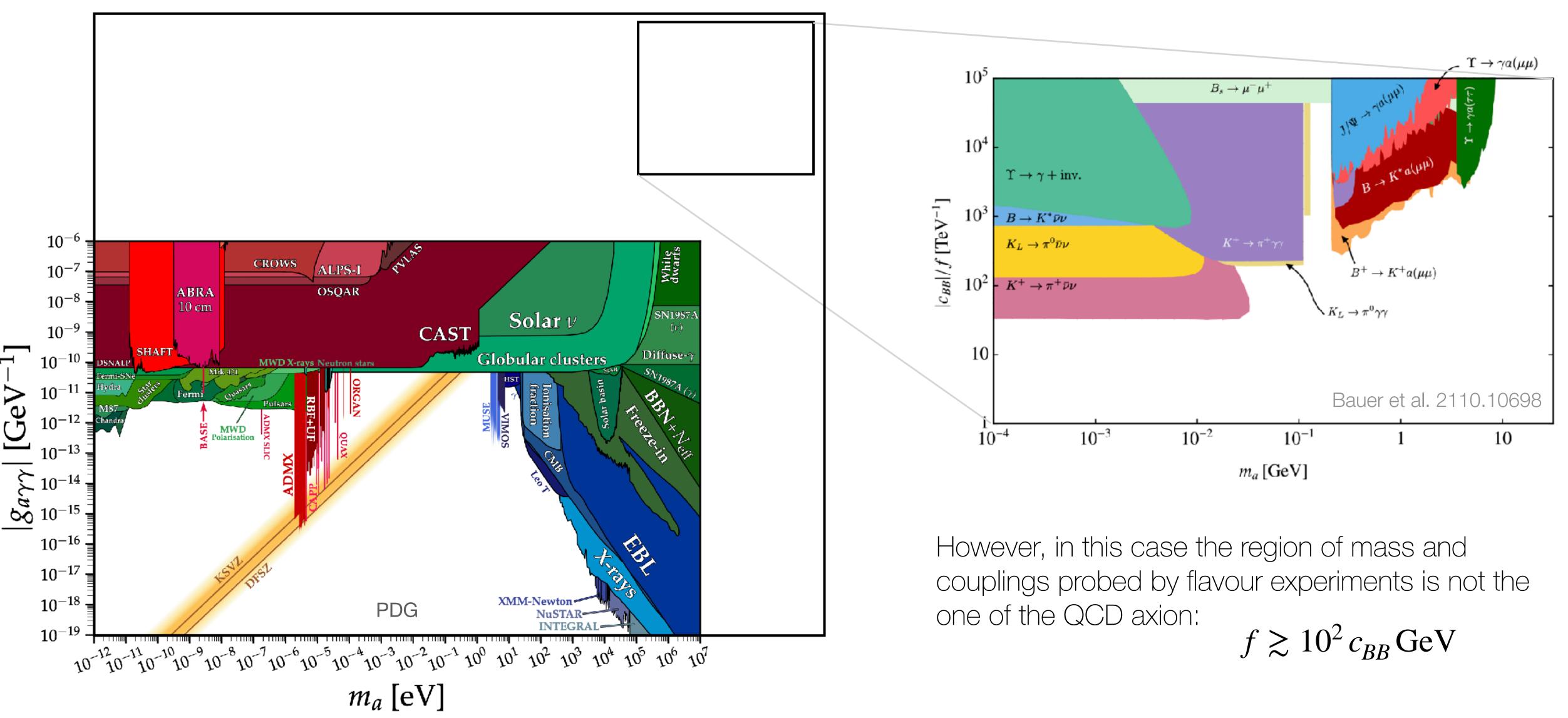
e.g. 1806.00660, 1905.01084, 1911.02591, Axion review 2003.01100



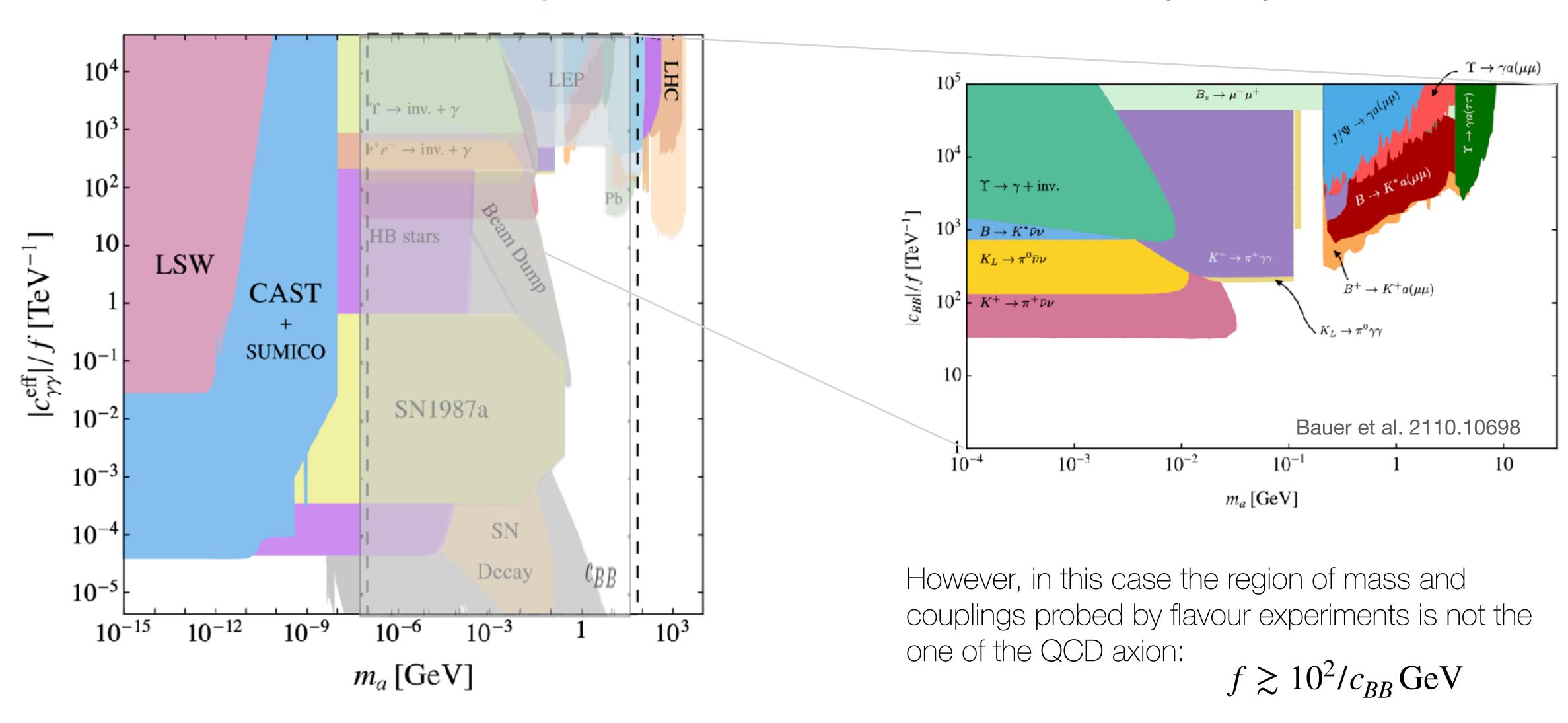
$$\mathcal{L}_{\text{fermion}} = \frac{\partial^{\mu} a}{f} \bar{\psi} \, \mathbf{k}_{\psi} \gamma_{\mu} \, \psi$$

For large flavour-mixings, $k_{ij}\sim O(1)$, these bounds are very strong, also surpassing astrophysical ones!

Also flavour-universal couplings to EW bosons induce FCNC (via SM-like penguin diagrams)



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Backup

U(2)⁵ flavour symmetry

In first approximation only the 3rd generation couples to the Higgs

In this case the SM enjoys a $U(2)^5$ global symmetry

$$G_F = U(2)_q \times U(2)_\ell \times U(2)_u \times U(2)_d \times U(2)_e$$
 Barbieri et al. [1105.2296, 1203.4218, 1211.5085]

The minimal breaking of this symmetry to reproduce the SM Yukawas is:

$$Y_{u(d)} = y_{t(b)} \begin{pmatrix} \Delta_{u(d)} & x_{t(b)} \mathbf{V}_q \\ 0 & 1 \end{pmatrix}, \qquad Y_e = y_{\tau} \begin{pmatrix} \Delta_e & x_{\tau} \mathbf{V}_{\ell} \\ 0 & 1 \end{pmatrix} \begin{array}{c} X_{t,b,\tau} \text{ are } \mathcal{O}(1), \ \mathbf{V}_{\ell} \ll 1 \end{array}$$

This is a very good approximate symmetry: the largest breaking has size $\epsilon \approx y_t |V_{ts}| \approx 0.04$

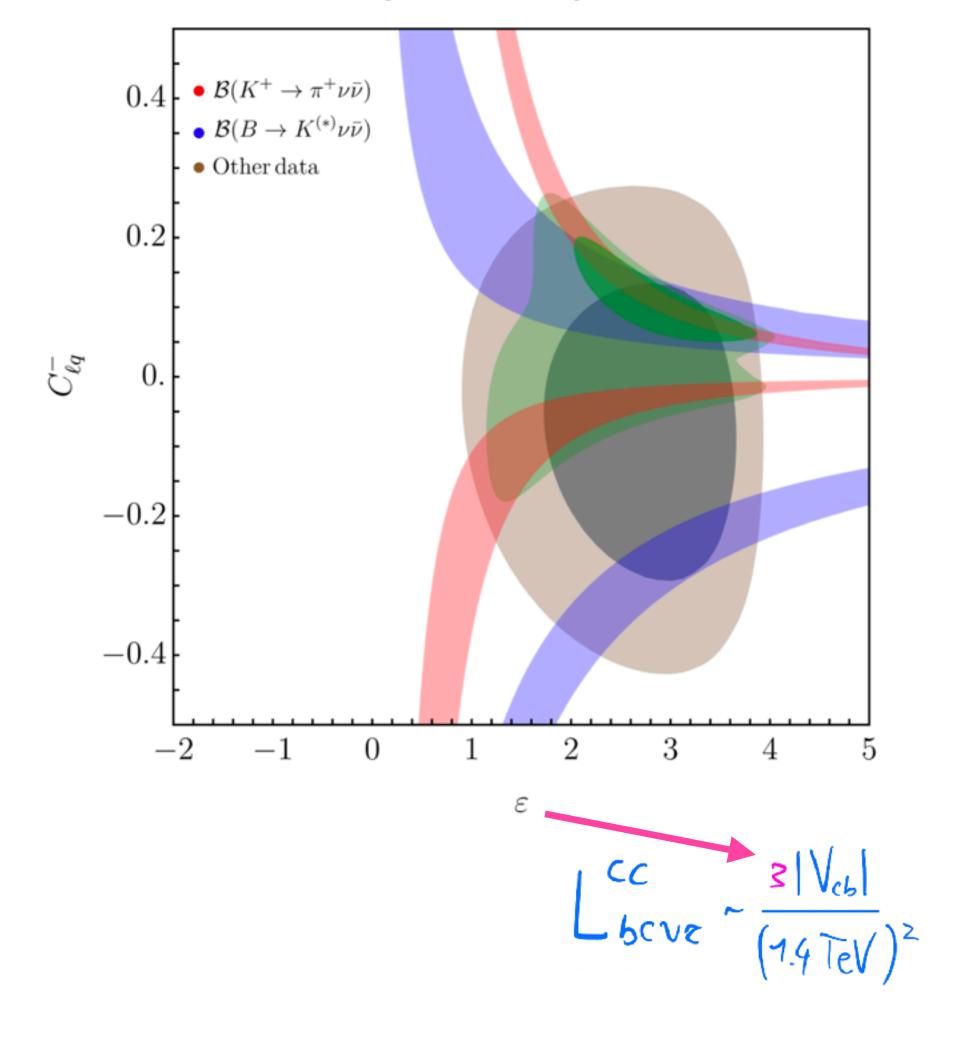
Diagonalizing quark masses, the V_q doublet spurion is fixed to be ${
m V}_q=\kappa_q(V_{td}^*,V_{ts}^*)^T$ See also Fuentes-Martin, Isidori, Pagès, Yamamoto [1909.02519]

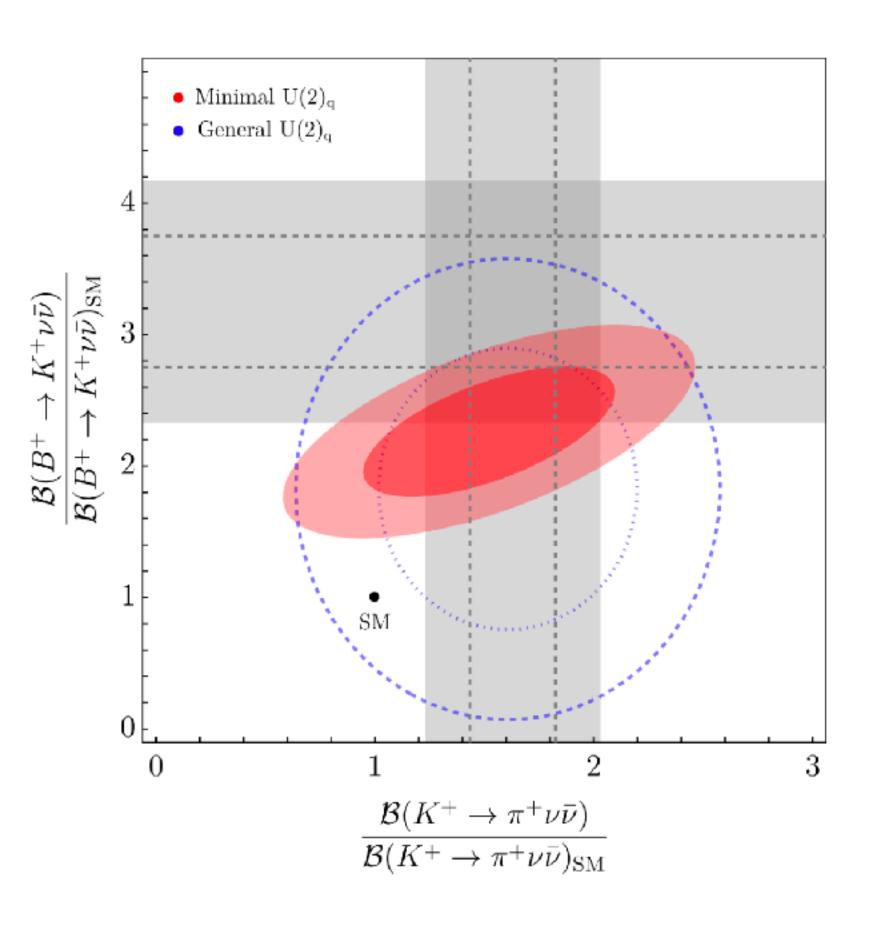
U(2)⁵ flavour symmetry and data

$$Q^\pm_{\ell q} = (\bar q_L^3 \gamma^\mu q_L^3)(\bar \ell_L^3 \gamma_\mu \ell_L^3) \pm (\bar q_L^3 \gamma^\mu \sigma^a q_L^3)(\bar \ell_L^3 \gamma_\mu \sigma^a \ell_L^3)$$

$$\tilde{V} = -\varepsilon V_{ts} \begin{pmatrix} \kappa V_{td}/V_{ts} \\ 1 \end{pmatrix}$$
 Minimal U(2)q: $\kappa = 1$.

Allwicher et al. [2410.21444]





Rare Semileptonic and Leptonic decays

Let us look at the flavour structure: other rare decays into muons

$$\mathcal{L}_{CFT} > \frac{C_{ij}}{\Lambda^2} \left(\overline{q}_i^i \, \chi_{\alpha} \, q_i^j \right) \left(\overline{\mu}_{\alpha} \, \chi^{\alpha} \mu_{\alpha} \right)$$

2σ bound on		LHCb '23	2210.07221	PDG 2024	hep-ph/0311084	LHCb '20	2011.09478
	Λ	R(K)	$B_s \rightarrow \mu\mu$	$B_d \rightarrow \mu\mu$	$K_L \rightarrow \mu\mu$	$K_S \rightarrow \mu\mu$	$D^0 \rightarrow \mu\mu$
Anarchic flavour	c = 1	56 TeV	33 TeV	18 TeV	74 TeV	c = i $10.7 TeV$	6.9 TeV
CKM-like		$c_{CKM} = V_{ts} $	$c_{CKM} = V_{ts} $	$c_{CKM} = V_{td} $	$c_{CKM} = V_{td}V_{ts} $	$c_{CKM} = i V_{td}V_{ts} $	$c_{CKM} = V_{cb}V_{ub} $
(MFV, U(2),)	$c = c_{CKM}$	11 TeV	6.6 TeV	1.6 TeV	1.4 TeV	0.2 TeV	0.086 TeV

$$C_{ij} \sim \begin{pmatrix} \mathcal{E}_{\gamma} & \lambda^{5} & \lambda^{3} \\ \lambda^{5} & \mathcal{E}_{z} & \lambda^{2} \\ \lambda^{3} & \lambda^{2} & 1 \end{pmatrix}$$

In new physics scenarios with **CKM-like flavour structure**, the **strongest constraints in the quark-muon couplings come from bsµµ observables**.