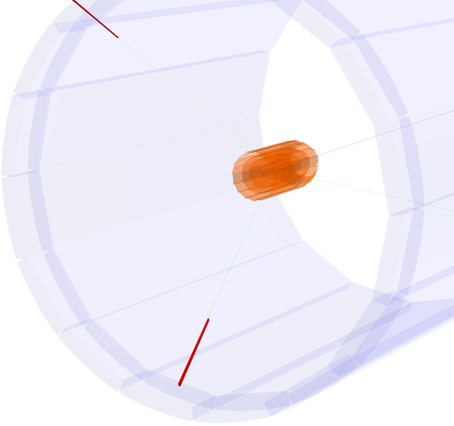


Search for Axion-Like Particles produced in e^+e^- collisions at the Belle II experiment.

10th January 2021 Epiphany Conference 2021

Michael De Nuccio (michael.de.nuccio@desy.de)

Belle II

HELMHOLTZ RESEARCH FOR GRAND CHALLENGES

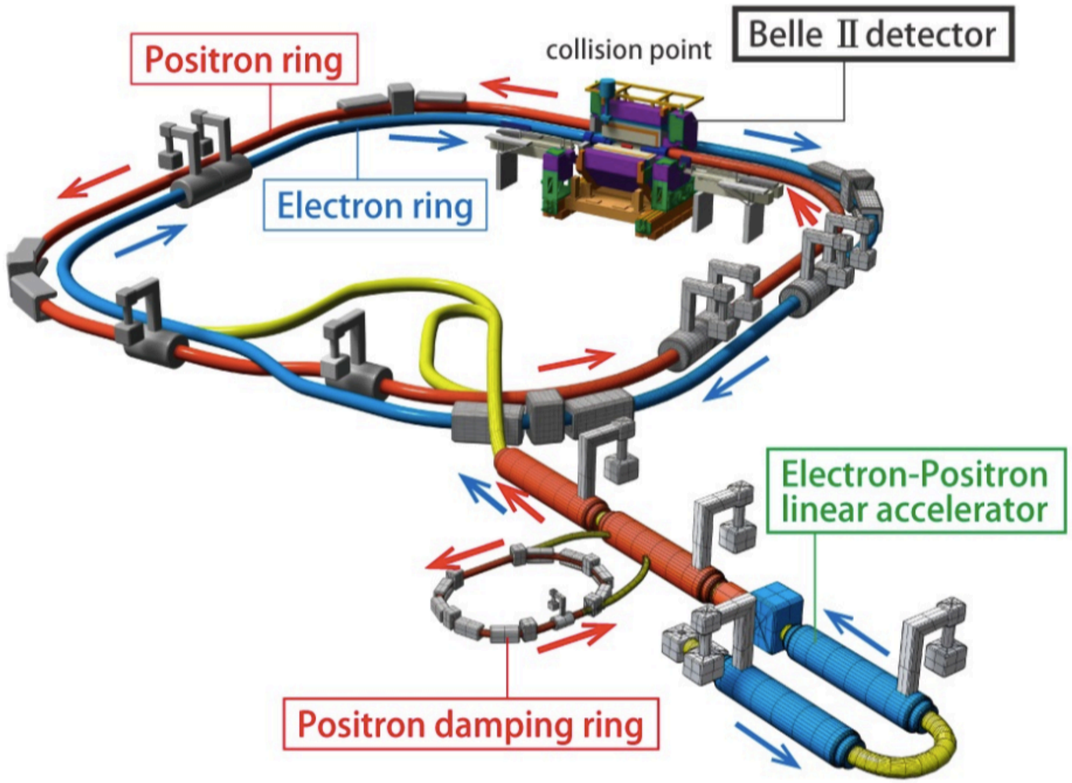
Universität Hamburg
DER FORSCHUNG | DER LEHRE | DER BILDUNG

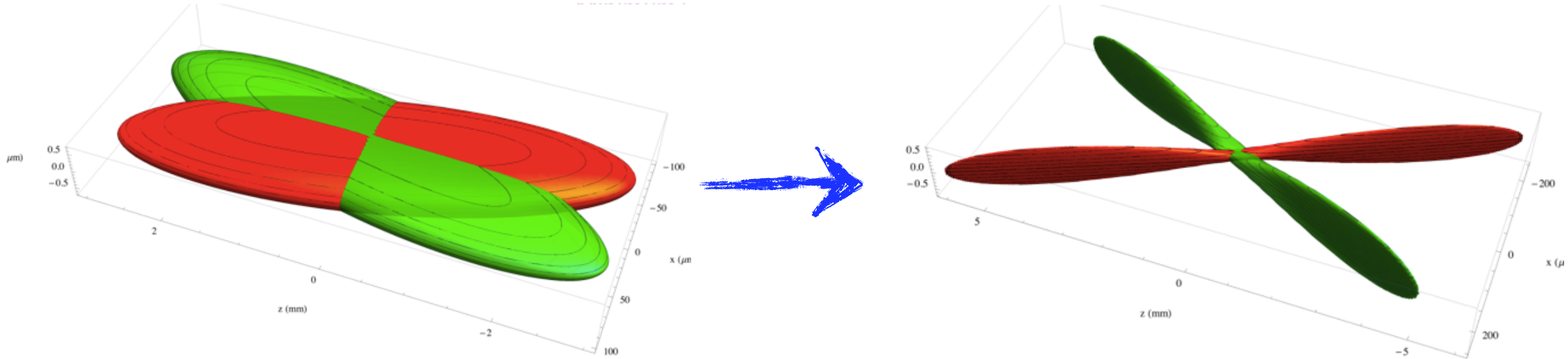
DESY.

SuperKEKB and Belle II.

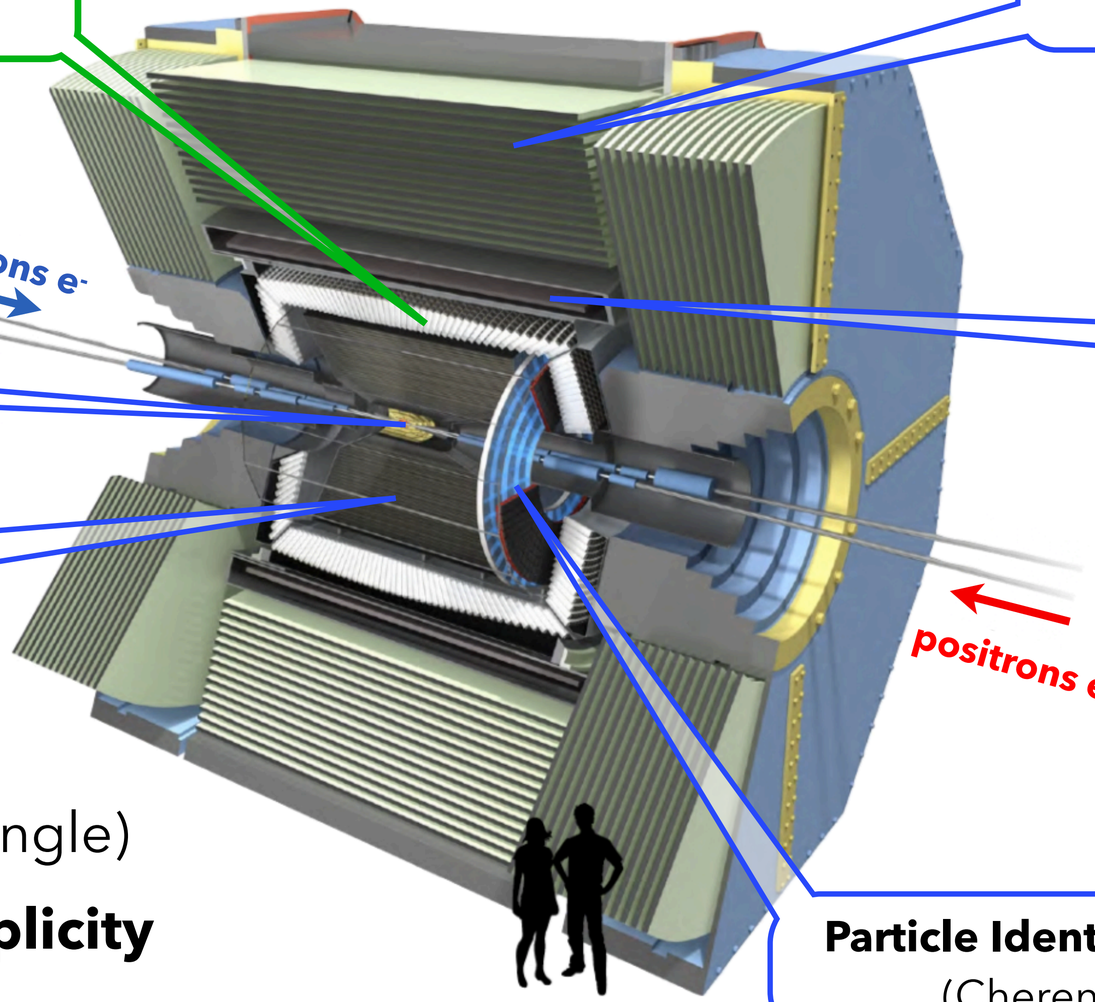
SuperKEKB

- **Asymmetric e^+e^- collider**
  @ $\Upsilon(4S)$ energy = 10.58 GeV
- Second-generation **B factory**
  (optimized to produce a lot of B mesons)
- **50 times** increase in luminosity with respect to predecessor KEKB: ~50 ab^{-1}





Belle II



Electromagnetic Calorimeter
(CsI(Tl) crystals)

K_L and μ detector
(resistive plates and scintillators)

Vertex Detectors
(silicon pixels & strips)

Magnet
(1.5 T)

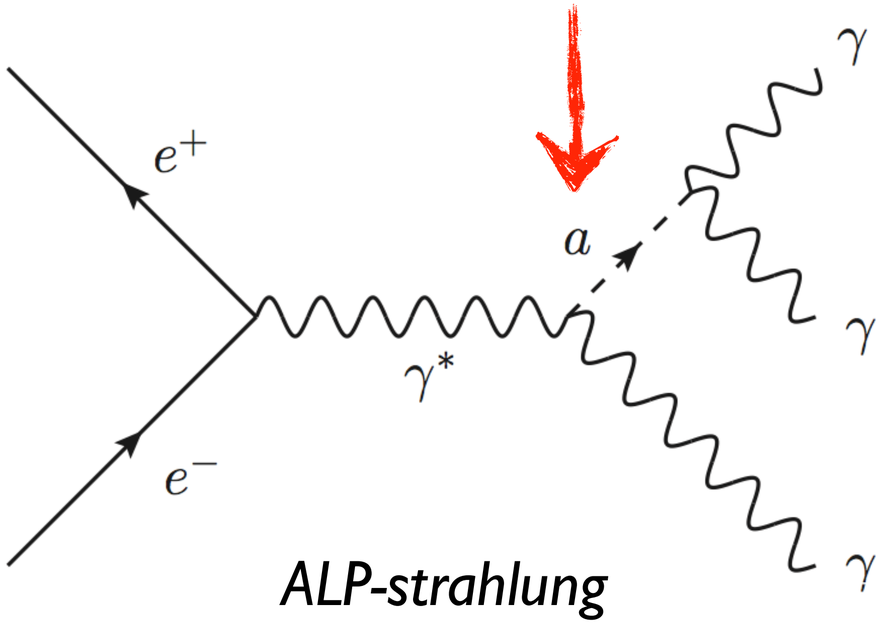
Tracking Detector
(drift chamber)

Particle Identification Detectors
(Cherenkov radiation)

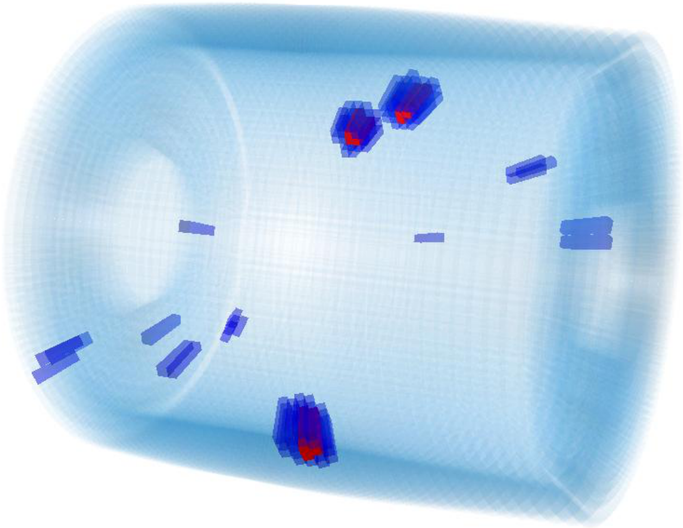
- **Hermetic detector** (90% of solid angle)
- **Dedicated triggers for low multiplicity**
- **Clean environment** (e^+e^- collider)

Axion-Like Particles.

Physical process



ALP-strahlung

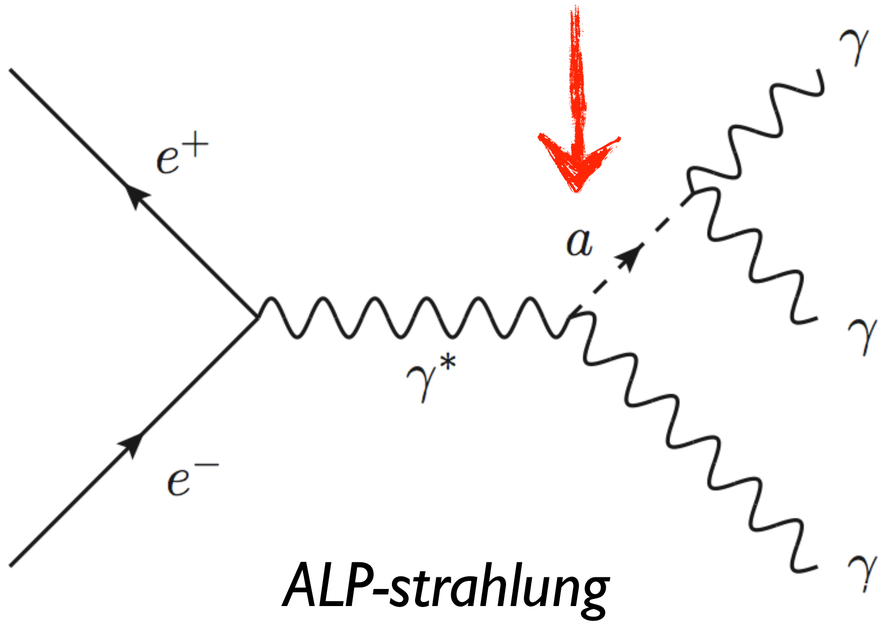


Belle II simulation, signal only

- Axion-like particles (***a***, **ALP**s):
  ~axions, but no mass-coupling constraint
  Light, neutral, pseudoscalar
- Possible portals to **Dark Sector**

Dolan, Ferber, Hearty, Kahlhoefer, Schmidt-Hoberg, JHEP 1712 (2017) 094

Physical process



ALP-strahlung

Signal: $e^+e^- \to \gamma a$, $a \to \gamma\gamma$

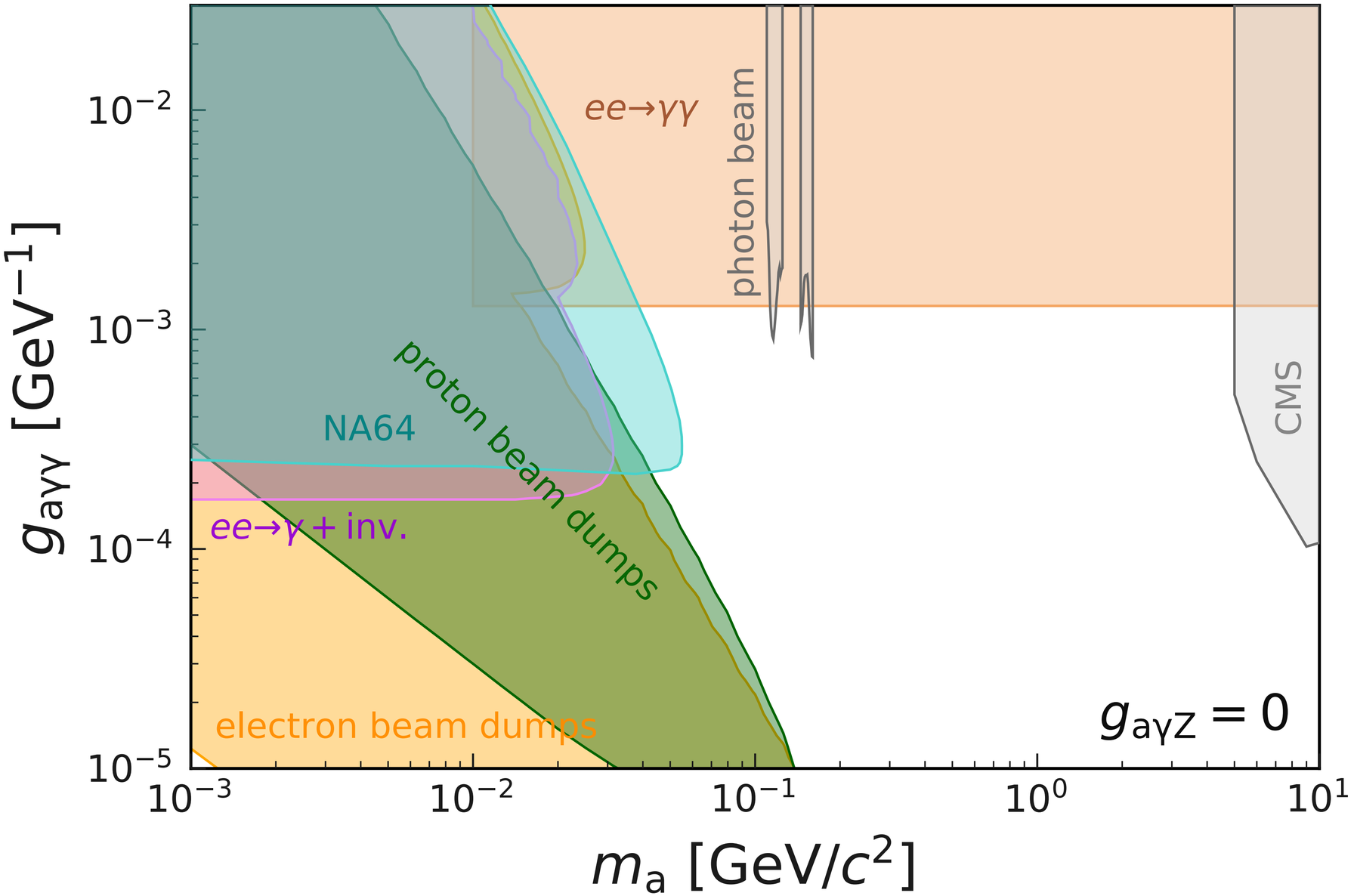
- 3-γ final state
- No tracks
- No missing energy

Main **backgrounds**:

- $e^+e^- \to \gamma\gamma(\gamma)$
- $e^+e^- \to e^+e^-(\gamma)$
  (if we don't reconstruct the tracks)
- $e^+e^- \to \pi^0/\eta/\eta' \gamma$
  negligible peaking backgrounds

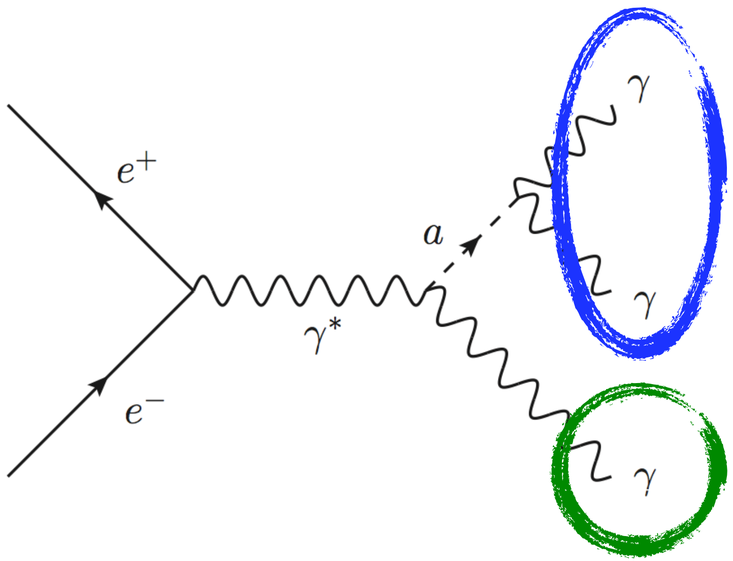
- **Peak hunt** throughout the kinematically-allowed mass spectrum (multiple mass hypotheses)
- **1D fit** of signal peak over smooth background

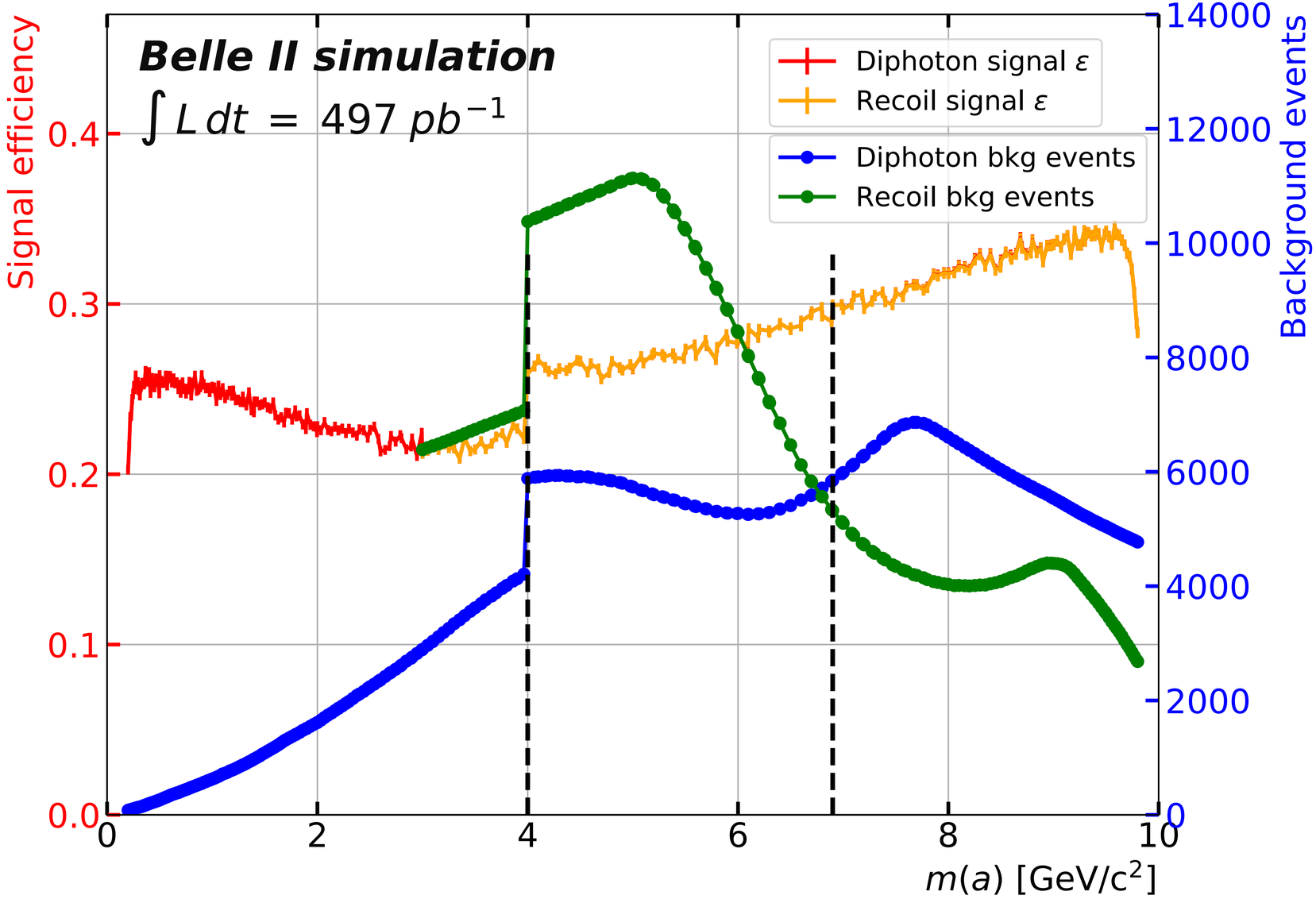
Previous status of searches



Selection performances

- Background rate
- Signal efficiency

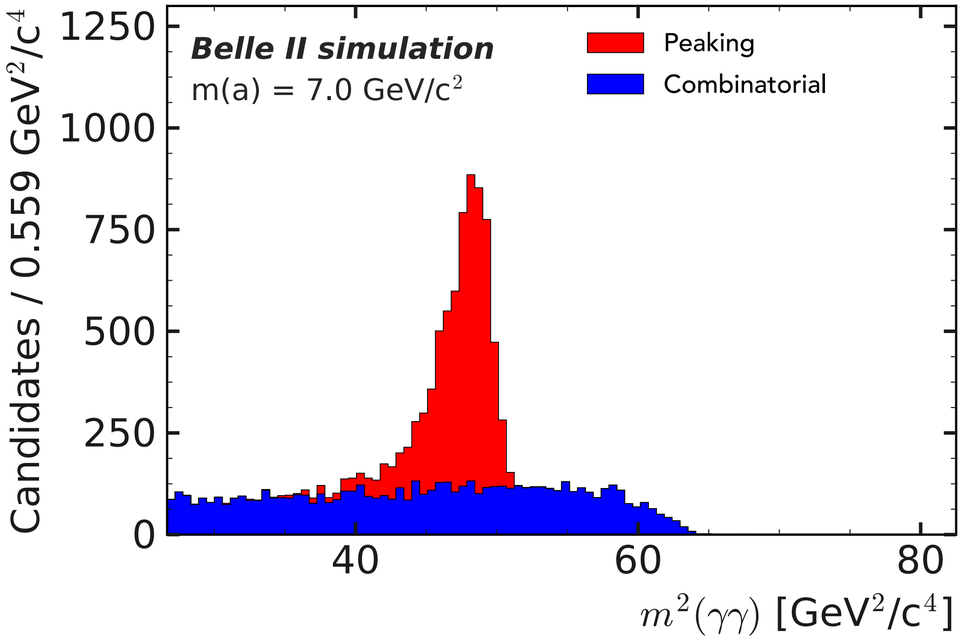




Signal & background modeling

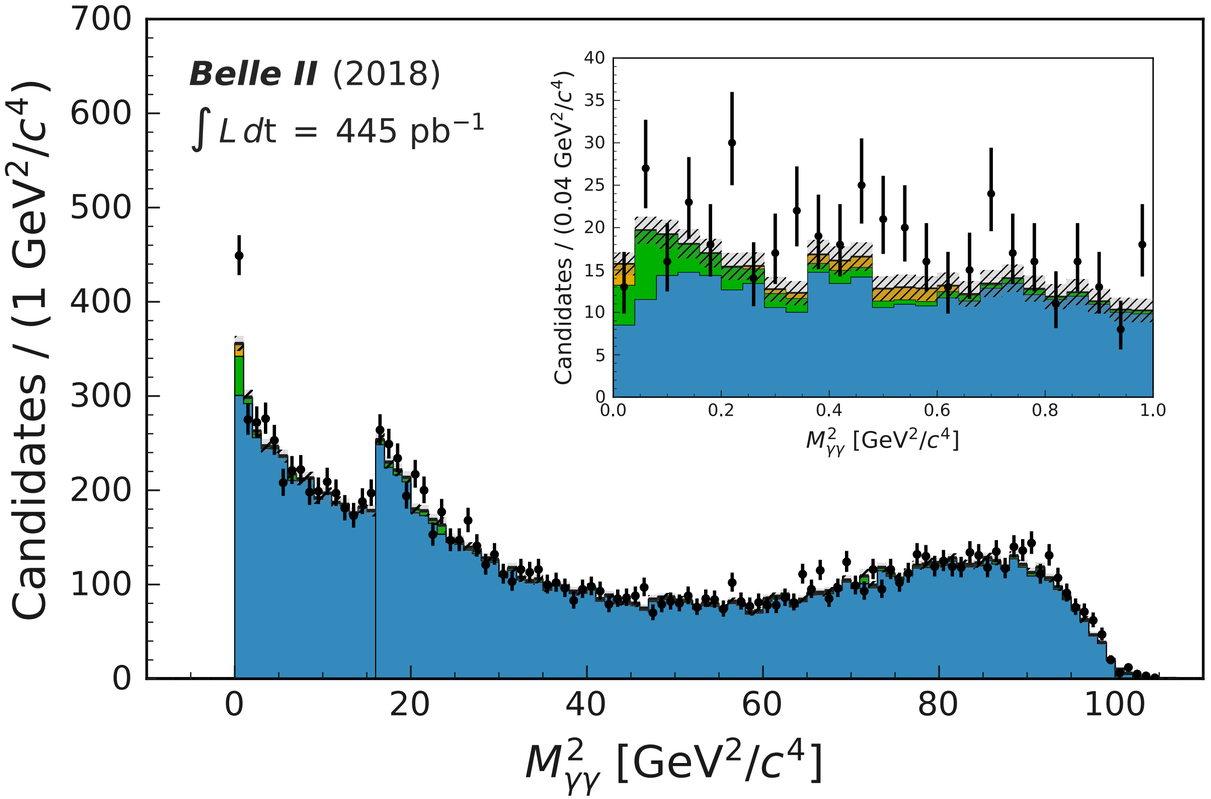
(3 $\gamma \Longrightarrow$ 3 candidates per event)

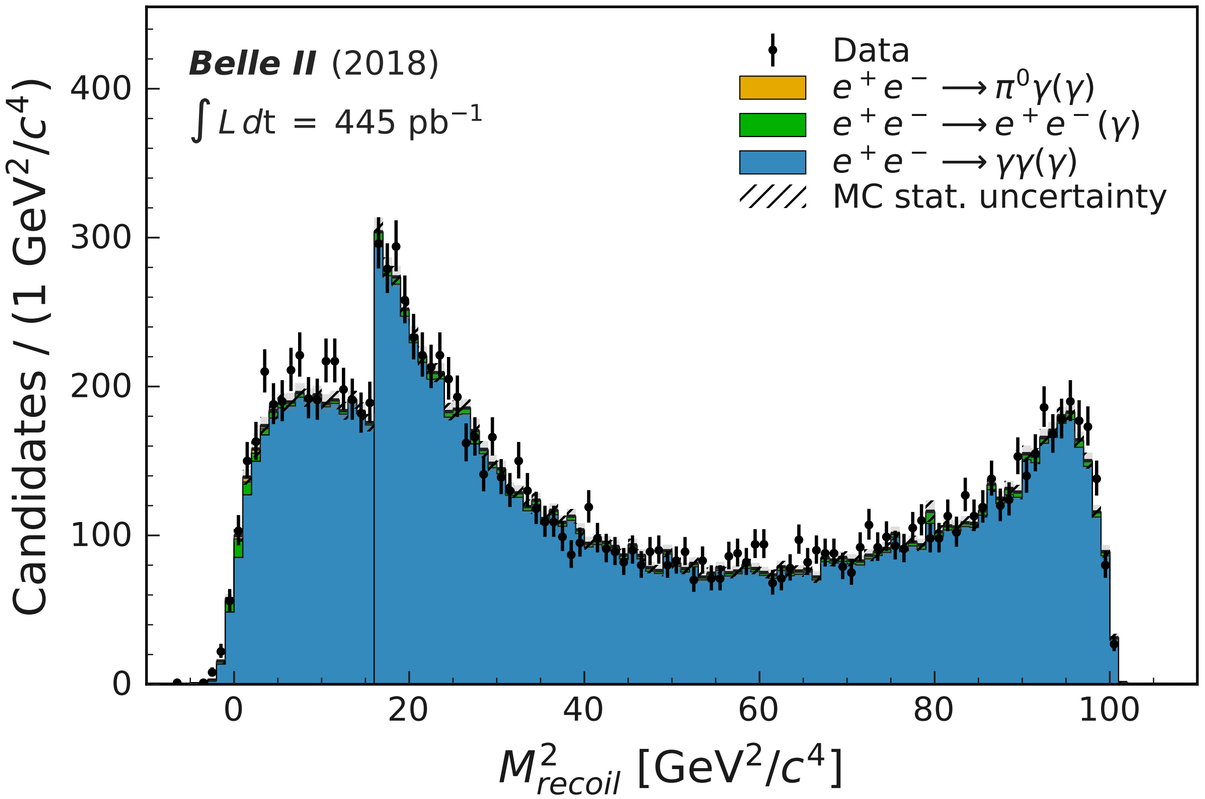
- **Signal**:
  - **Peaking** component: modeled with a **Crystal Ball** (CB):
    fit each MC sample
    interpolate parameters
    fixed for the final fit
  - **Combinatorial** component:
    modeled with a **Kernel Density Estimator** (KDE)
    fixed for the final fit



- **Background**:
  - Modeled with a **polynomial**
  - Choice of **polynomial order** and **fit range**:
    reduced χ^2 and smoothness criteria
  - Polynomial parameters are **floating** for the final fit

Data/MC comparison





Great agreement already in 2018 (data taking for calibration & tuning purposes)

10.1103/PhysRevLett.125.161806

Upper Limit (UL) extraction

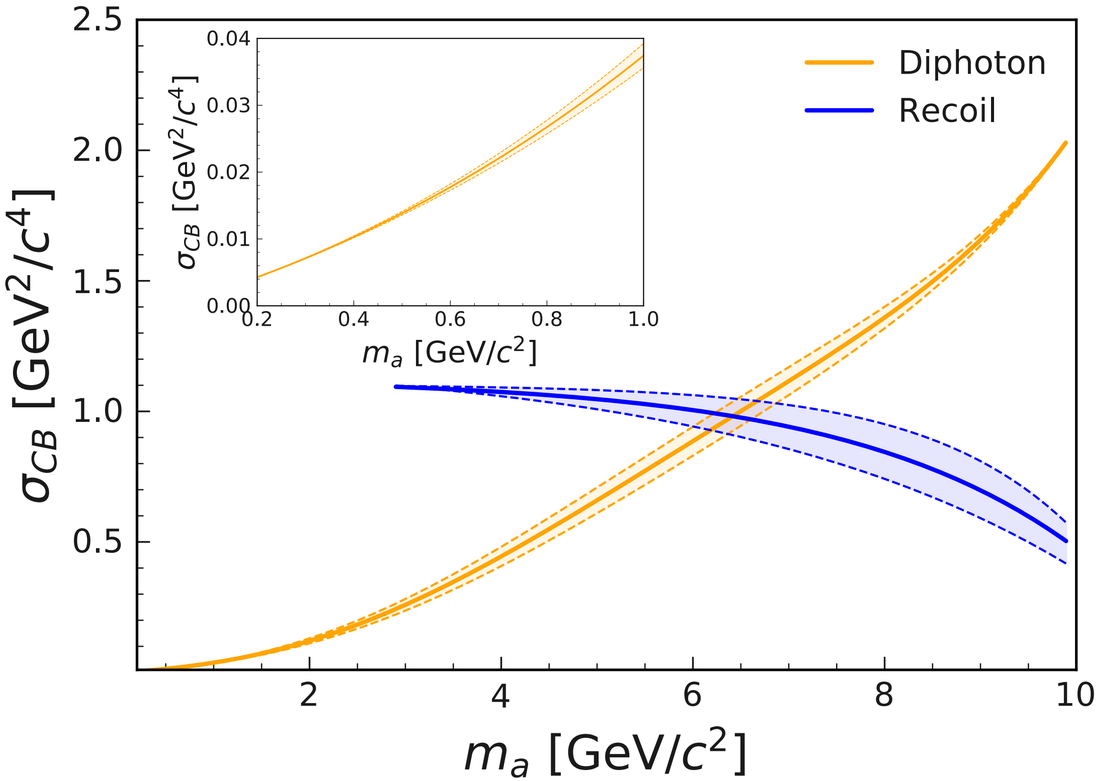
- Binned NLL approach, CLs method
- Allow only positive signal yields, i.e. cross section $\sigma_{a\gamma\gamma} \geq 0$
- ALP mass scan in **steps of 0.5** σ_{CB} to search for signal peaks
- If **no global significance > 3** is found (with systematics): we **set limits**
  - No *local* significance > 3 has been found

Systematic uncertainties

Systematic uncertainties are **small wrt statistical** uncertainties.

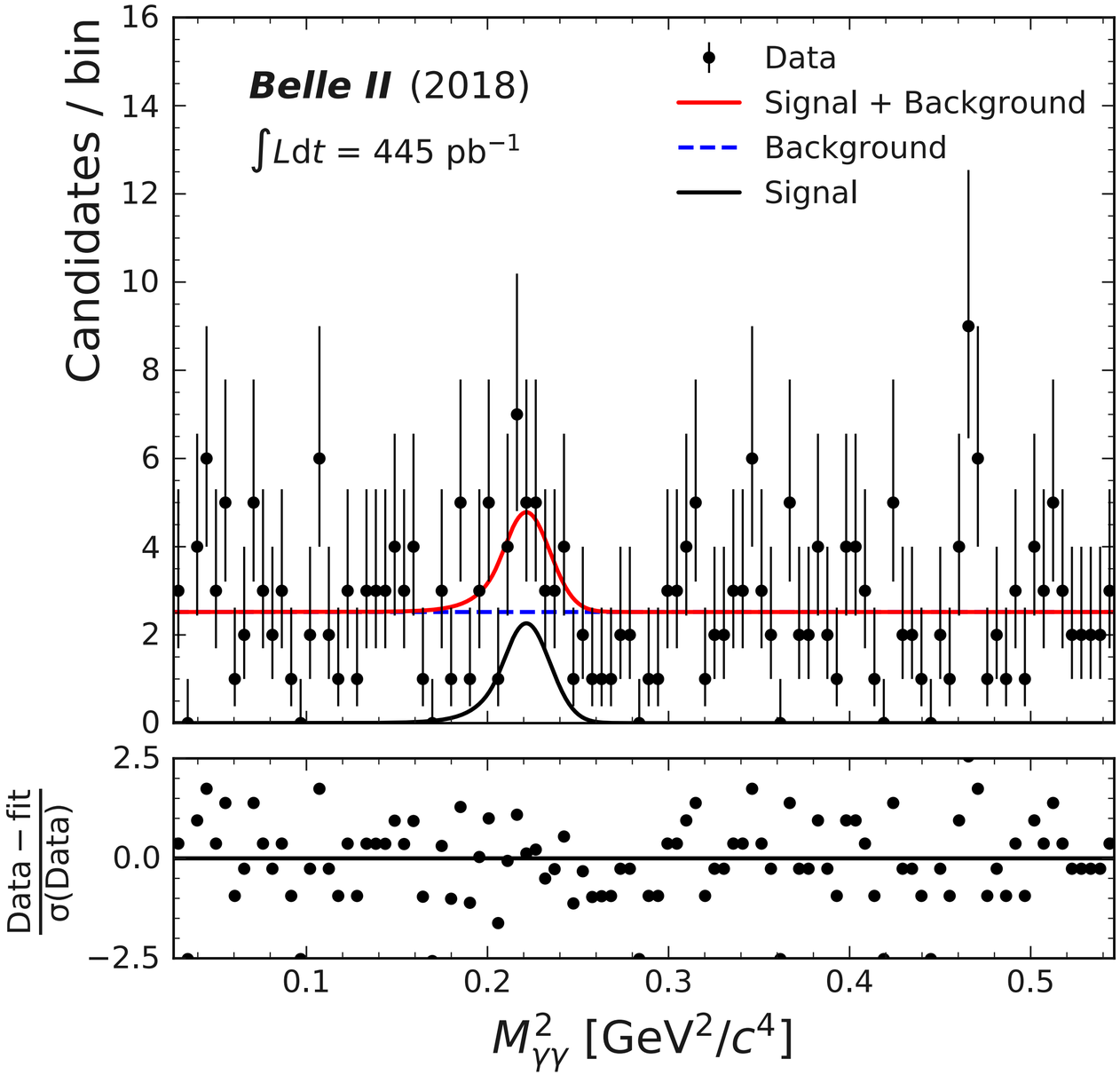
Systematics are from:

- **Choice of background polynomial order & fit range** (least irrelevant):
  - Modify order and range and re-perform UL extraction, take the weakest limit (highest UL ⟺ lowest significance)
- **Signal efficiency**
- **Signal resolution**
  from photon resolution studies



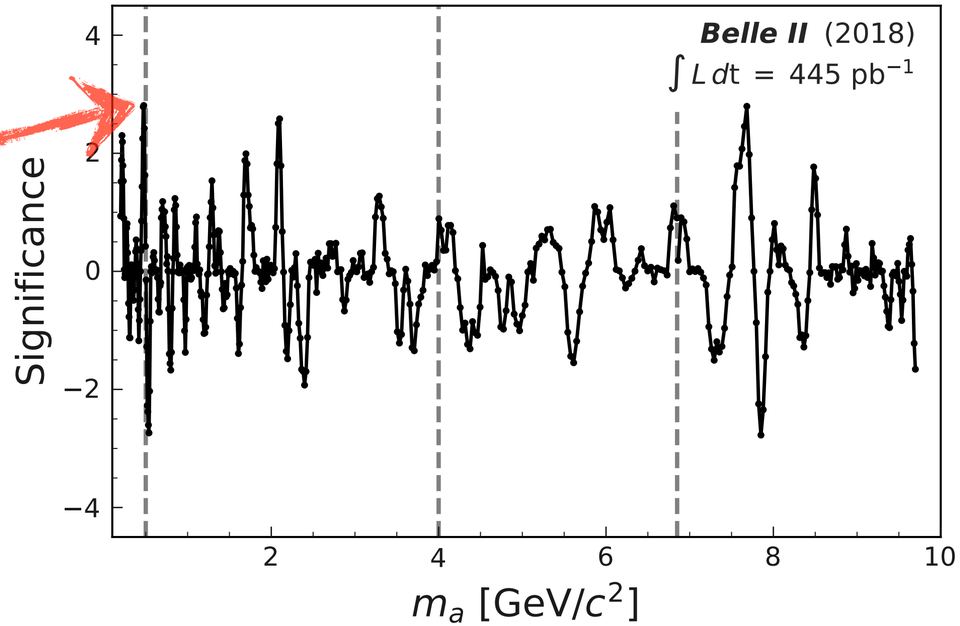
10.1103/PhysRevLett.125.161806

Results



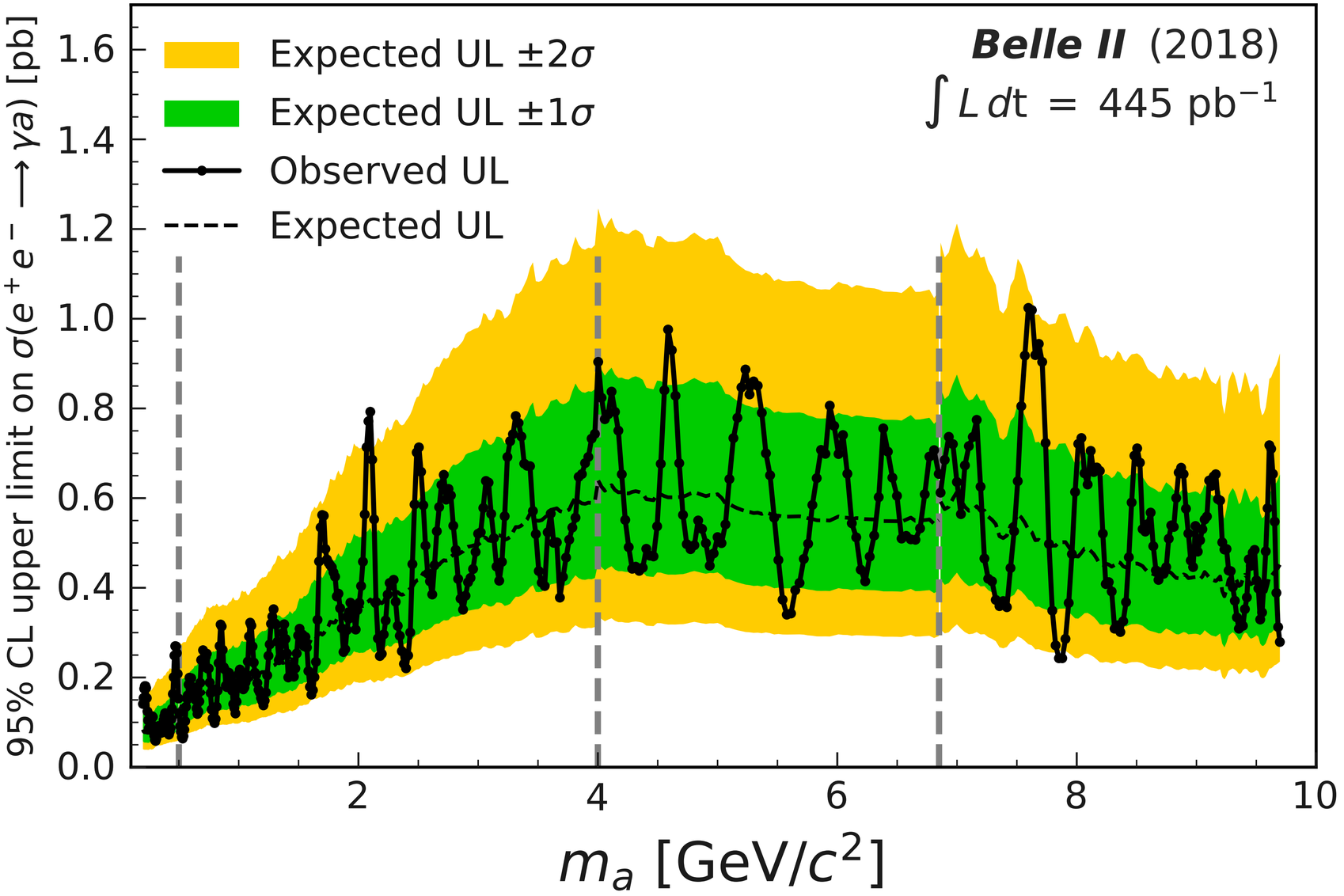
Max local significance:

2.8 σ @ m_a = 0.477 GeV/c²



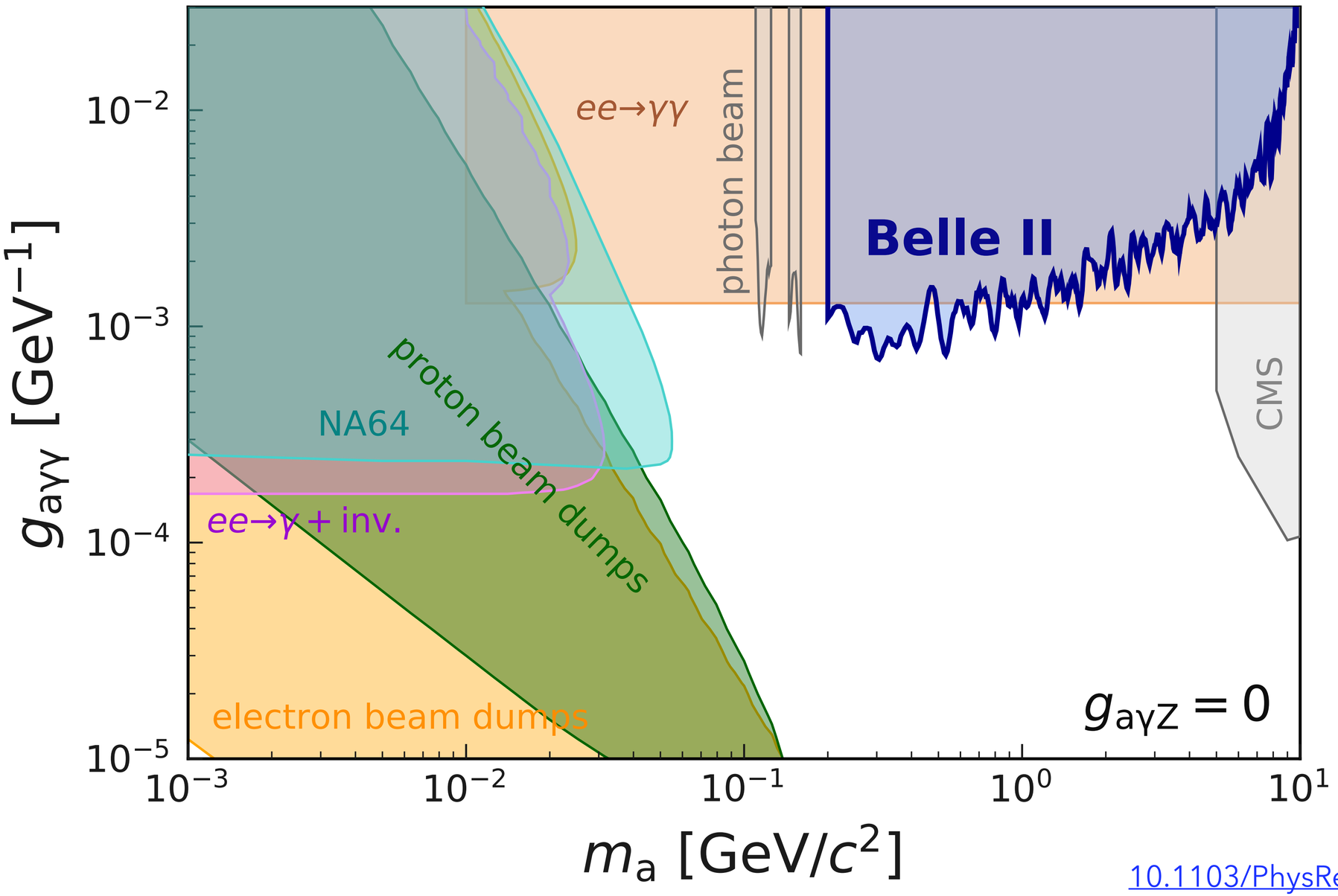
10.1103/PhysRevLett.125.161806

Results



10.1103/PhysRevLett.125.161806

Results



10.1103/PhysRevLett.125.161806

Summary.

Summary

- **Search** for the direct production of a light pseudoscalar **ALP** *a* decaying into two photons
  - $e^+e^- \to \gamma a$, $a \to \gamma\gamma$
  - $m_a \in [0.2, 9.7]$ GeV/c^2
- **No evidence for ALPs**
- Set 95% CL UL on $g_{a\gamma\gamma}$
  - These are the **strongest limits** to date **for $m_a \in$ [0.2, 1] GeV/c^2**
- Results published in **PRL**: 10.1103/PhysRevLett.125.161806

Backup.

Selection variables

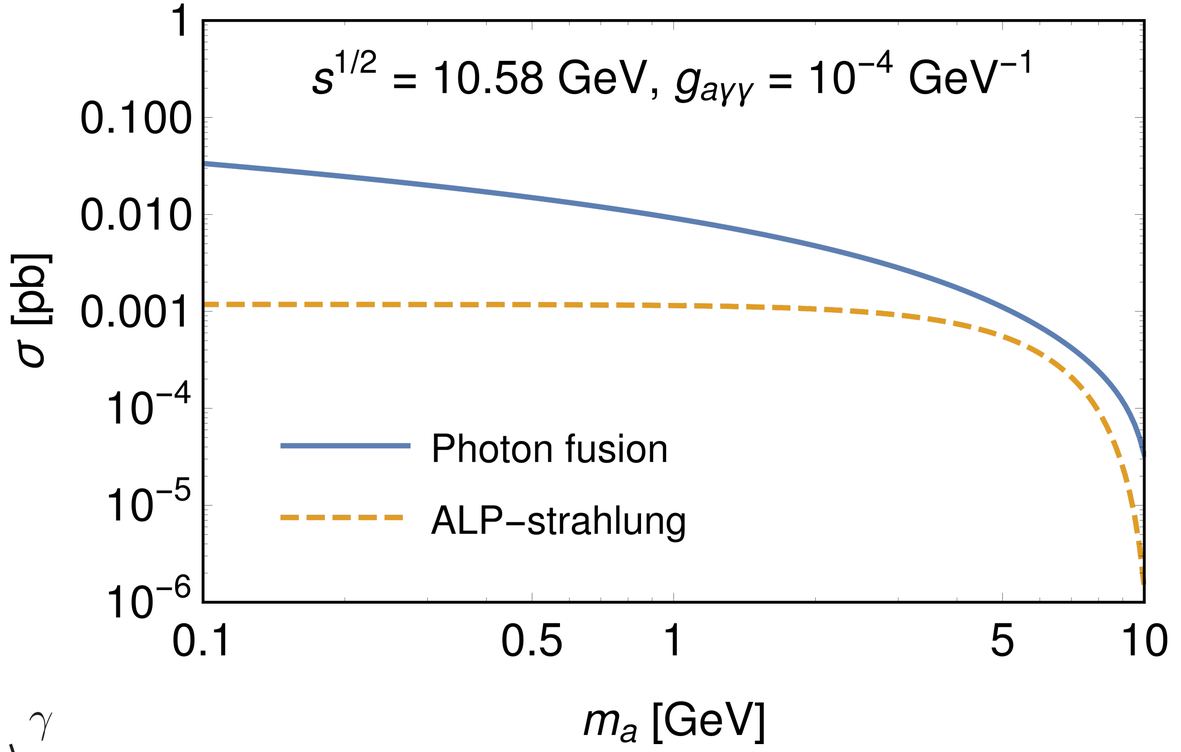
- PFM studies
- Sanity requirements
- Studies on other datasets
- Studies on sidebands

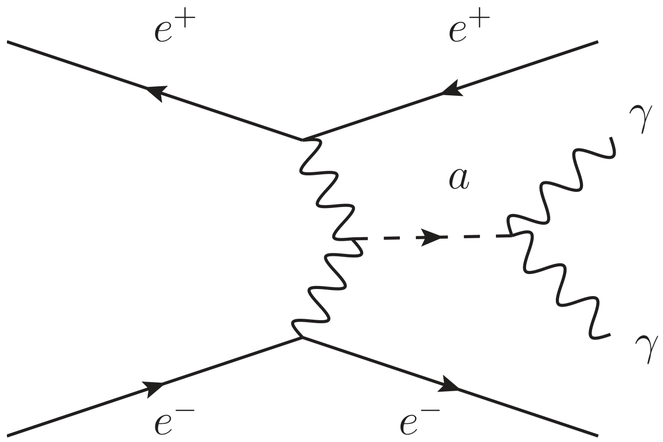
Cuts are listed in the order they are applied

- $E_\gamma \geq 650$ MeV if $m_a \geq 4.0$ GeV/c^2
  $E_\gamma \geq 1000$ MeV if $m_a < 4.0$ GeV/c^2
- $37.3° \leq \theta_\gamma \leq 123.7°$ (barrel acceptance)
- clusterNHits > 1.5
- 3 most energetic γ
- $0.88\ \sqrt{s} \leq m_{\gamma\gamma\gamma} \leq 1.03\ \sqrt{s}$
  (9.31 GeV/c^2 $\leq m_{\gamma\gamma\gamma} \leq$ 10.90 GeV/c^2)
- TimeVar* < 10
- 0 good tracks
- $\Delta\theta \geq 0.014$ rad OR $\Delta\varphi \geq 0.4$ rad
- clusterZernikeMVA of most isolated photon > 0.6

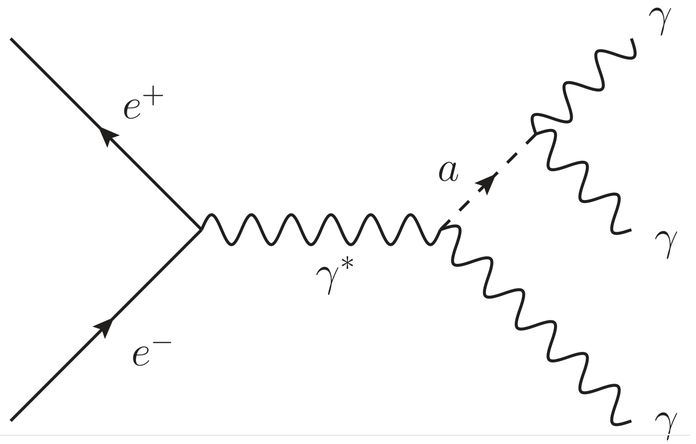
*) TimeVar = $|(t - \sum(t/\Delta t^2)/\sum(1/\Delta t^2))/\Delta t|$

ALPs - theory





Photon-fusion

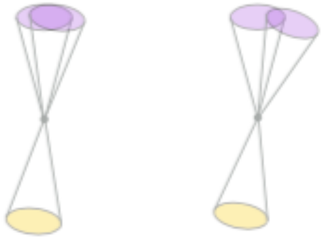


ALP-strahlung

Dolan, Ferber, Hearty, Kahlhoefer, Schmidt-Hoberg, JHEP 1712 (2017) 094

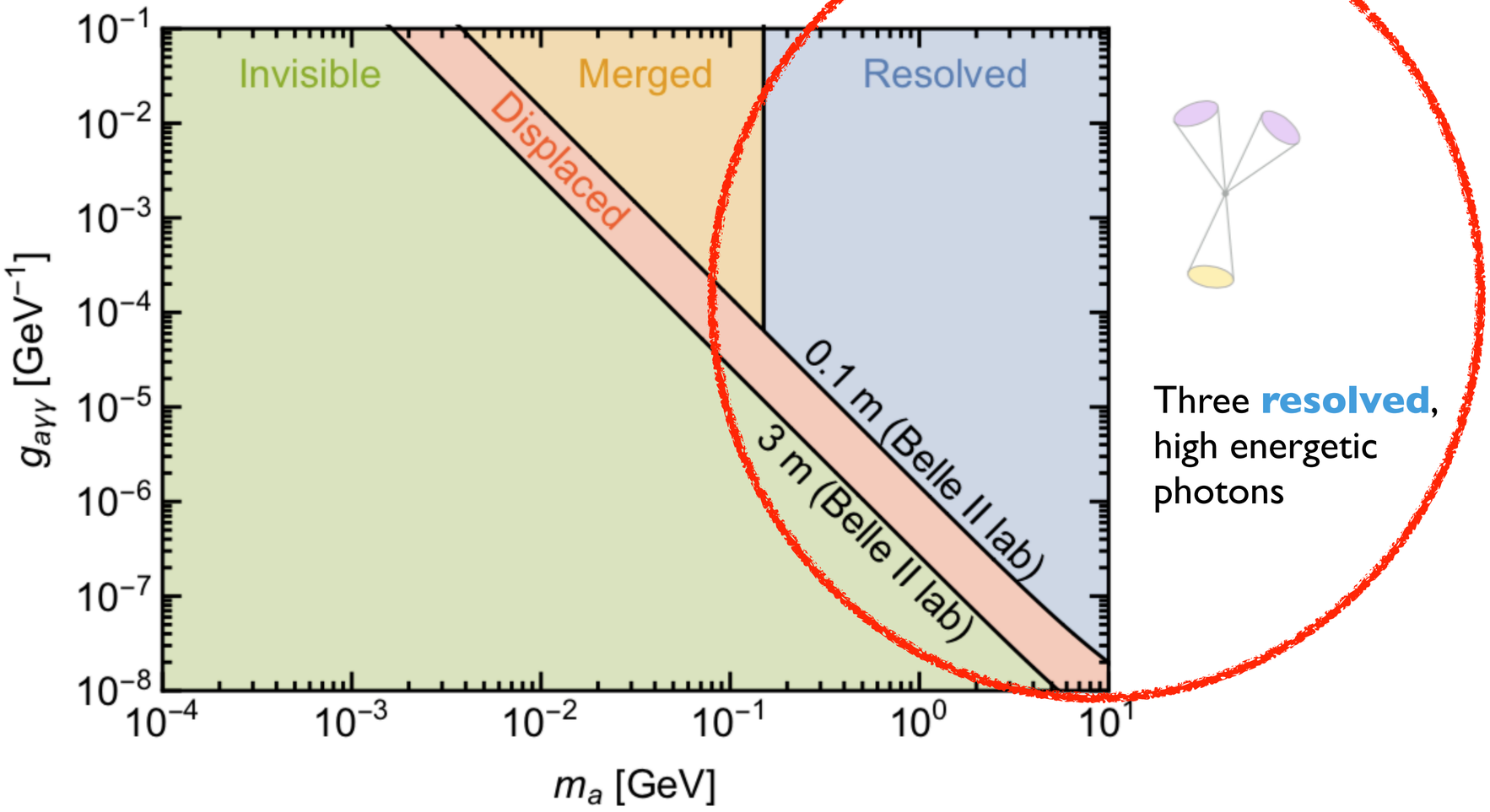
ALPs - analysis

Two of the photons overlap or **merge**



Current focus

ALP decays outside of the detector or decays into **invisible** particles: single photon final state



Three **resolved**, high energetic photons

Dolan, Ferber, Hearty, Kahlhoefer, Schmidt-Hoberg, JHEP 1712 (2017) 094

Selection variables

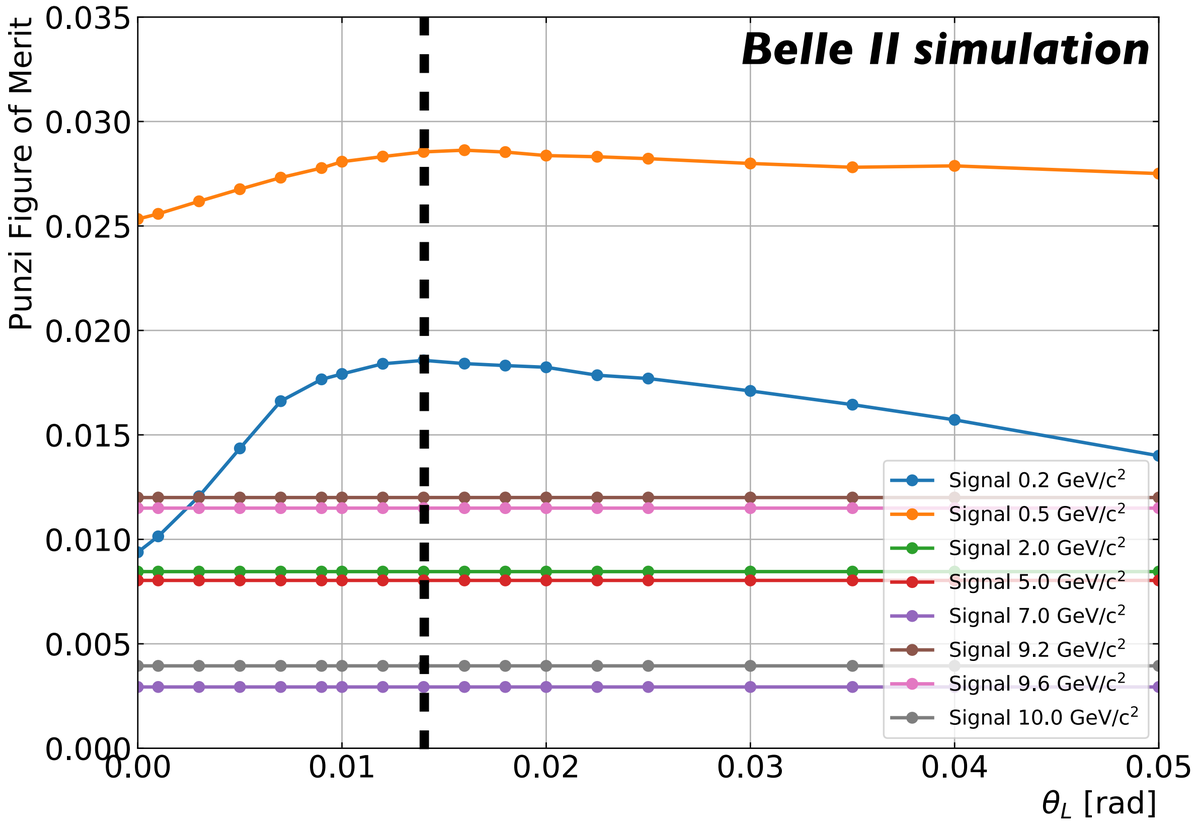
Selection optimization via **maximization** of **Punzi Figure of Merit** (PFM)

- PFM as function of 1 variable
  - Other variables fixed
- Vary cut on that variable
- For multiple ALP masses
- Repeat for all variables

$$PFM = \frac{\epsilon_S}{\frac{a}{2} + \sqrt{B}}$$

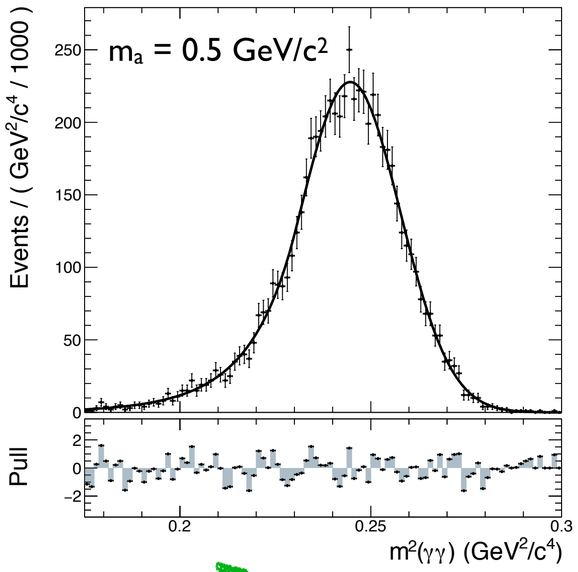
- ϵ_S: Signal efficiency
- B: # bkg candidates passing the cuts
- a: Number of sigmas corresponding to one-sided Gaussian tests at a given significance (a=5)

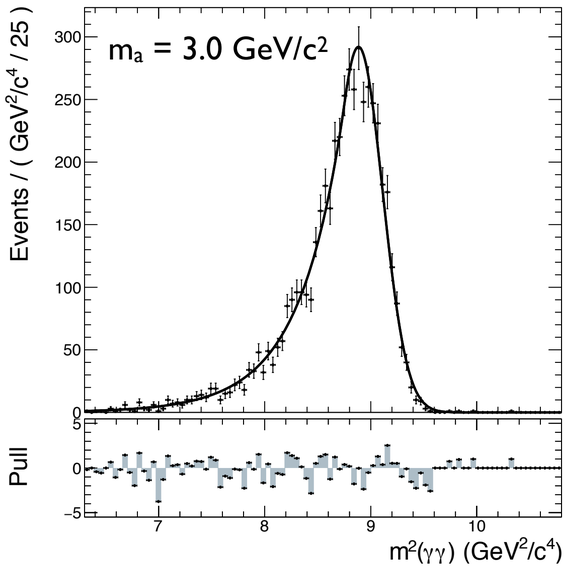
One **example** of PFM scan:

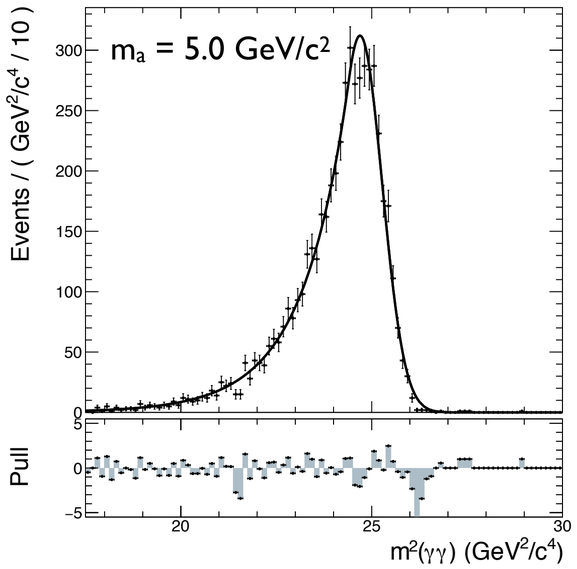


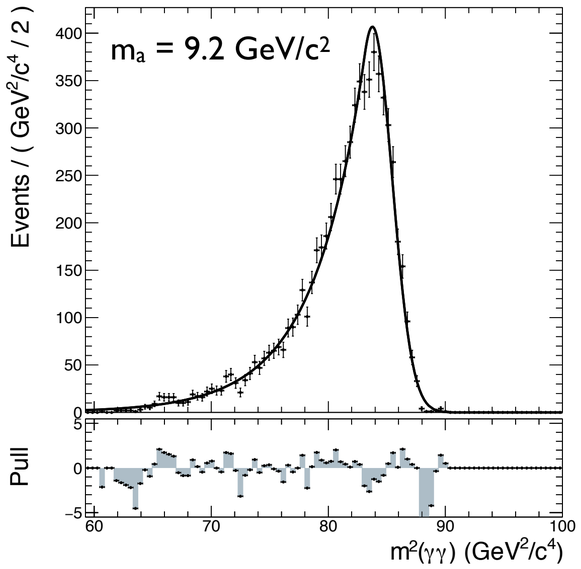
(https://arxiv.org/pdf/physics/0308063.pdf)

Signal peaking



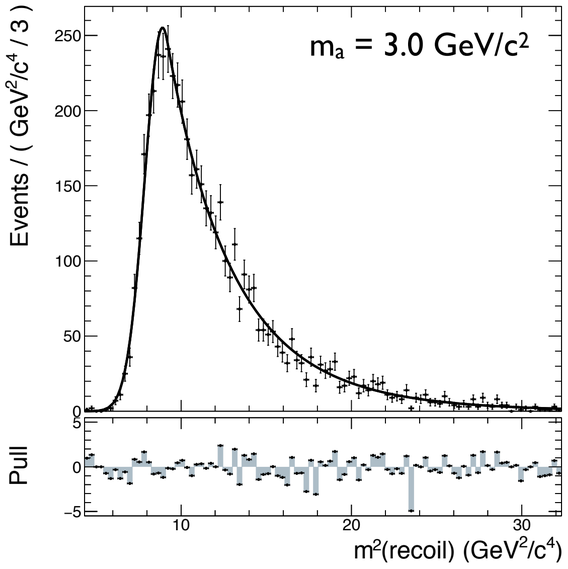


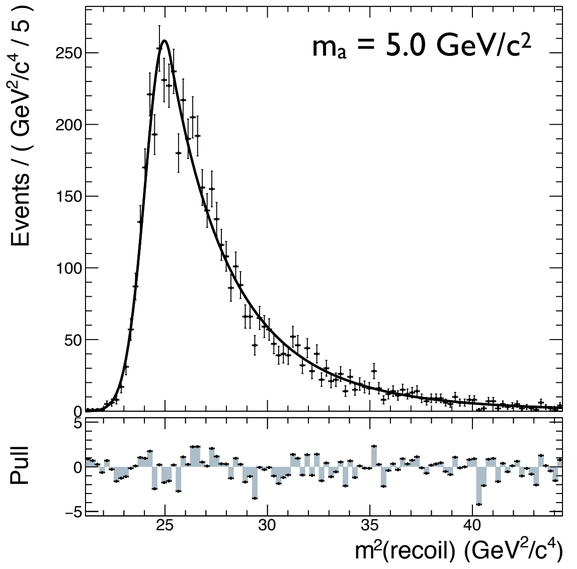


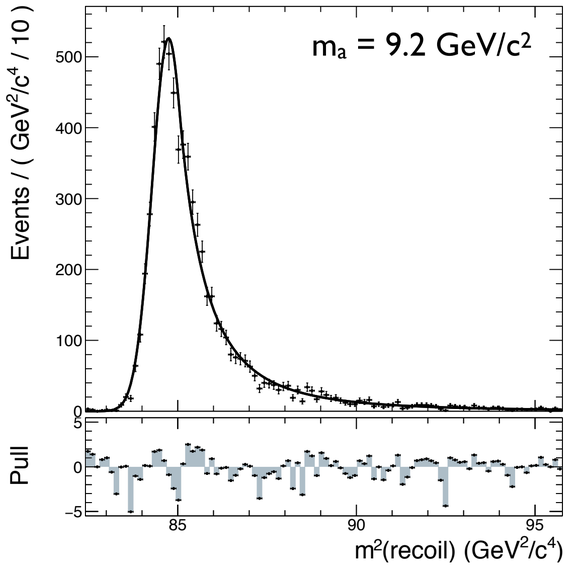


Squared **diphoton** mass

Squared **recoil** mass







Background modeling

Choice of polynomial order & fit range with reduced χ^2 and smoothness criteria

- $m_a \in [0.2, 0.5]$ GeV/c^2 $\Longrightarrow$ 2nd order, fit range $[m_a^2\text{-}20\cdot\sigma_{CB}, m_a^2\text{+}30\cdot\sigma_{CB}]$
- $m_a \in [0.5, 6.85^*]$ GeV/c^2 $\Longrightarrow$ 4th order, fit range $[m_a^2\text{-}20\cdot\sigma_{CB}, m_a^2\text{+}30\cdot\sigma_{CB}]$
- $m_a \in [6.85^*, 9.7]$ GeV/c^2 $\Longrightarrow$ 5th order, fit range $[m_a^2\text{-}25\cdot\sigma_{CB}, m_a^2\text{+}25\cdot\sigma_{CB}]$