#### Dark matter at Belle II

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## Light dark matter at accelerators

- Ordinary matter in our universe is outnumbered by dark matter 5:1 (in mass).
- All evidence so far relies only on gravity, including our primary measurement:  $O_{12} = 0.12$

 $\Omega_{\chi} h^2 = 0.12$  Planck, 1807.06209

 But: best explanations of relic density require additional interactions with SM particles



Basis of experimental DM search program

- Due to lack of WIMP observation, light DM ( $m \lesssim 10~{\rm GeV}$ ) has become increasingly attractive



Belle II with leading sensitivity for many light dark sectors

#### **Dark sector mediators**

• Focus here on **portal interactions** with DM mediator:  $\mathcal{O}_{SM} \times \mathcal{O}_{DS}$ 

Lowest mass  
dimension
$$\left\{ \begin{aligned} -\frac{\kappa_{Y}}{2}F'_{\mu\nu}B^{\mu\nu} & \text{Vector portal} \\ (\lambda_{3}\phi + \lambda_{4}\phi^{2})H^{\dagger}H & \text{Higgs portal} \\ \frac{\lambda_{3}\phi + \lambda_{4}\phi^{2}}{f_{a}}F^{\mu\nu}, \frac{\partial_{\mu}a}{f_{a}}\overline{\psi}\gamma^{\mu}\gamma^{5}\psi & \text{Axion portal} \\ \frac{\lambda_{3}\phi + \lambda_{4}\phi^{2}}{g_{N}LHN} & \text{Sterile neutrino portal} \end{aligned} \right.$$

• Further guidance for light DM in reach of Belle II:

**Residual annihilations**  $\chi \bar{\chi} \rightarrow f \bar{f}$  have to be suppressed at late times (low temperatures) to be compatible with cosmo bounds,

in particular from CMB Slatyer, 1506.03811, Planck, 1807.06209

## **Scalar mediators**

• Scalar mediator + fermionic DM avoids strong CMB constraints through velocity-suppressed (p-wave) annihilation  $\langle \sigma v \rangle \sim v^2$ 

$$\mathcal{L} \supset -\frac{1}{2}m_{\phi}^{2}\phi^{2} - m_{\chi}\bar{\chi}\chi - \lambda_{3}\phi H^{\dagger}H - y_{\chi}\bar{\chi}\chi\phi$$

• Mixing after EW symmetry breaking:  $h_{125}$  and new scalar S

$$\sim \frac{m_f}{v} \sin \theta$$

b

Direct production from  $e^+e^-$  suppressed

Instead: heavy quark loops in meson decays

Filimonova et al., 1911.03490, Kachanovich et al., 2003.01788

## **Scalar mediators**

Scalar decay ff XX  $B \rightarrow K + \text{two charged tracks}$ Filimonova et al., 1911.03490  $10^{-1}$ LHCb 10-2  $\pi\pi + KK$ Belle II  $\mu\mu$ 10-3 θ BaBar  $10^{-4}$ HL-LHCb ττ  $10^{-5}$ 2 3 Δ *m*<sub>5</sub> [GeV] Sascha's talk

Requires  $m_S > 2m_{\chi}$ :

Parameter space with right relic density **ruled out** by direct detection + BaBar

Krnjaic, 1512.04119

## **Pseudoscalar ALPs**

- **Pseudoscalar mediator** with fermionic DM avoids direct detection
- Special type of pseudoscalar: **Axion-like particle** defined by shift symmetry  $a \rightarrow a + c$

$$\begin{aligned} \mathscr{L} \supset \frac{1}{2} \partial_{\mu} a \partial^{\mu} a - \frac{m_a}{2} a^2 + \frac{\partial^{\mu} a}{f_a} \sum_{f} \frac{c_f}{2} \bar{f} \gamma_{\mu} \gamma_5 f + \frac{c_{\chi}}{2} \frac{\partial^{\mu} a}{f_a} \bar{\chi} \gamma_{\mu} \gamma_5 \chi \\ + c_{GG} \frac{g_s^2}{(4\pi)^2} \frac{a}{f_a} G_{\mu\nu}^A \tilde{G}^{A,\mu\nu} + c_{BB} \frac{{g'}^2}{(4\pi)^2} \frac{a}{f_a} B_{\mu\nu} \tilde{B}^{\mu\nu} + c_{WW} \frac{g^2}{(4\pi)^2} \frac{a}{f_a} W_{\mu\nu}^A \tilde{W}^{A,\mu\nu} \end{aligned}$$

• Intriguing possibility: coupling to  $W^{A}_{\mu\nu}\tilde{W}^{A,\mu\nu}$  dominates

## **Pseudoscalar ALPs**



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# **Vector mediators**

• Massive vector mediator from broken U(1)'

$$\mathcal{L} \supset g_{\chi} A^{\prime \mu} \bar{\chi} \gamma_{\mu} \chi + m_{A^{\prime}}^2 A^{\prime}_{\mu} A^{\prime \mu}$$

- Either direct charges for SM particles by gauging  $L_{\mu} L_{\tau}$ ,  $B L \dots$
- Or kinetic mixing  $-\epsilon/\cos\theta_W F'_{\mu\nu}B^{\mu\nu}$

$$\rightarrow A' f \bar{f}$$
 couplings  $\sim e \epsilon q_f$ 

• Single-photon search for  $A' \rightarrow \chi \chi$  in association with ISR  $\gamma$ : High sensitivity to on-shell A' production; gap in coverage for  $\bar{\chi} \chi$  production via off-shell A' ( $m_{A'} \gtrsim 10 \text{ GeV}$ ), but Belle II can also be sensitive

Essig et al., 1309.5084, EB et al., 2203.08824

• Simple setup (A' + light fermionic DM) in conflict with CMB

## Vector mediator + inelastic DM

• Inelastic DM avoids CMB (and direct detection) constraints:  $\chi_1 \chi_2 \to A^{'(*)} \to f \bar{f}$ 

annihilation strongly suppressed at late time due to low  $\chi_2$  abundance

- Heavier state decays via  $\chi_2 \to \chi_1 A^{'*} \to \chi_1 f \bar{f}$ , typically long-lived
- New signature:  $e^+e^- \rightarrow \chi_1\chi_2$  followed by displaced decay of  $\chi_2$



# **Dark showers**

- Entire novel class of signatures featuring light LLPs:
- Dark showers in strongly interacting dark sectors resembling SM QCD sector



(multiple) displaced decays number varying from event to event  Projections for Belle II LLP search improve greatly upon existing bounds from BaBar

## Conclusions

- Light dark sectors are well-motivated
- CMB constraints provide model building guidance
- Production at Belle II directly or through meson decay
- Output Decay of dark sector mediators in many cases visible and often displaced
- Belle II almost uniquely positioned among colliders to have leading sensitivity



Great examples in upcoming experimental talks