



Loops in inclusive determinations

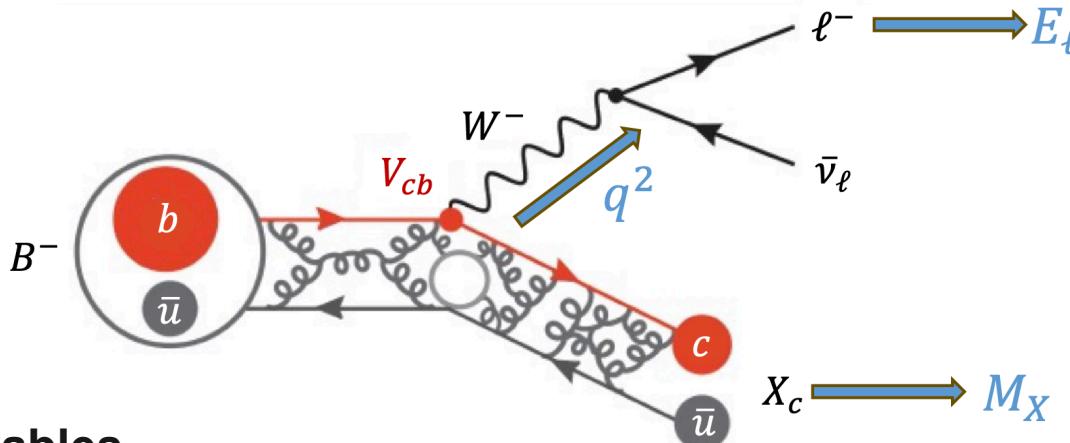
Matteo Fael (CERN)

2023 Belle II Physics Week - KEK - 31 Ott. 2023



Funded by
the European Union

Extraction of V_{cb} from inclusive $B \rightarrow X_c l \bar{\nu}_l$ decays



Observables

- Total rate $\Gamma_{\text{sl}} = \Gamma(B \rightarrow X_c \ell^- \bar{\nu}_\ell)$
- Moments of the differential distribution of an observables O

$$\langle (O)^n \rangle_{\text{cut}} = \int_{\text{cut}} (O)^n \frac{d\Gamma}{dO} dO \Bigg/ \int_{\text{cut}} \frac{d\Gamma}{dO} dO$$

- $O = E_\ell$: energy of the charged lepton in the B rest frame
- $O = M_X^2$: hadronic invariant mass
- $O = q^2$: leptonic invariant mass

see talks by M. Bordone, K. Vos & P. Gambino

Heavy Quark Expansion

Double series expansion in the strong coupling constant α_s and power suppressed terms Λ_{QCD}/m_b

- $\rho = m_c/m_b$
- Total rate

$$\Gamma_{\text{sl}} = \frac{G_F^2 m_b^5 A_{\text{ew}}}{192\pi^3} |V_{cb}|^2 \left[\left(1 - \frac{\mu_\pi^2}{2m_b^2} \right) \left(X_0(\rho) + \frac{\alpha_s}{\pi} X_1(\rho) + \left(\frac{\alpha_s}{\pi} \right)^2 X_2(\rho) + \left(\frac{\alpha_s}{\pi} \right)^3 X_3(\rho) + \dots \right) \right. \\ \left. + \left(\frac{\mu_G^2}{m_b^2} - \frac{\rho_D^3}{m_b^3} \right) \left(g_0(\rho) + \frac{\alpha_s}{\pi} g_1(\rho) + \dots \right) + \frac{\rho_D^3}{m_b^3} \left(d_0(\rho) + \frac{\alpha_s}{\pi} d_1(\rho) + \dots \right) + O\left(\frac{1}{m_b^4}\right) \right]$$

- Moments of differential distribution for some observable O

$$\langle O^n \rangle_{\text{cut}} = (m_b)^{mn} \left[X_0^{(O,n)} + \frac{\alpha_s}{\pi} X_1^{(O,n)} + \left(\frac{\alpha_s}{\pi} \right)^2 X_2^{(O,n)} + \frac{\mu_\pi^2}{m_b^2} \left(p_0^{(O,n)} + \frac{\alpha_s}{\pi} p_1^{(O,n)} + \dots \right) \right. \\ \left. + \frac{\mu_G^2}{m_b^2} \left(g_0^{(O,n)} + \frac{\alpha_s}{\pi} g_1^{(O,n)} + \dots \right) + \frac{\rho_D^3}{m_b^3} \left(d_0^{(O,n)} + \frac{\alpha_s}{\pi} d_1^{(O,n)} + \dots \right) + \frac{\rho_{LS}}{m_b^2} \left(l_0^{(O,n)} + \frac{\alpha_s}{\pi} l_1^{(O,n)} + \dots \right) + O\left(\frac{1}{m_b^4}\right) \right]$$

Why higher order corrections?

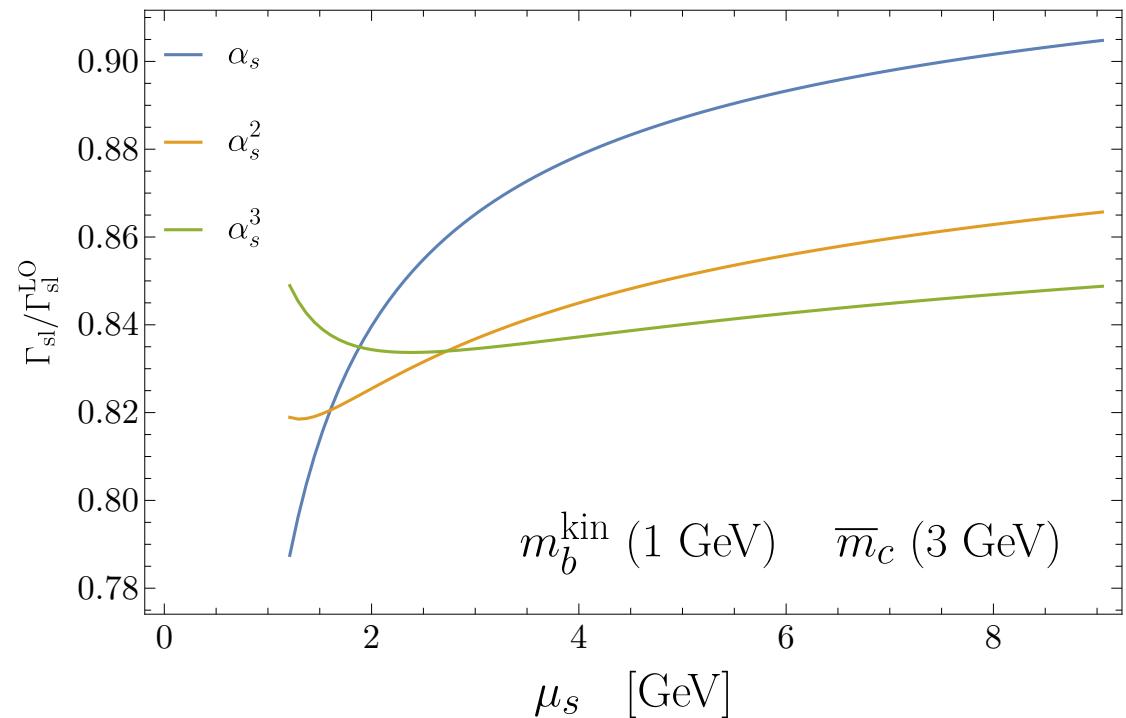
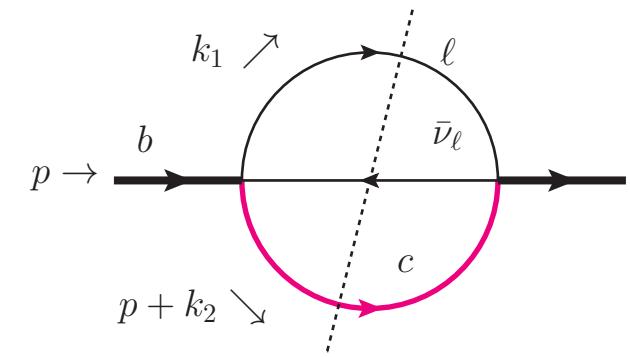
- Missing higher-order terms limit the prediction

$$\Gamma_{\text{sl}} = \frac{G_F^2 m_b^5 A_{\text{ew}}}{192\pi^3} |V_{cb}|^2 \times \left(X_0(\rho) + \frac{\alpha_s}{\pi} X_1(\rho) + \left(\frac{\alpha_s}{\pi}\right)^2 X_2(\rho) + \left(\frac{\alpha_s}{\pi}\right)^3 X_3(\rho) + \dots \right)$$

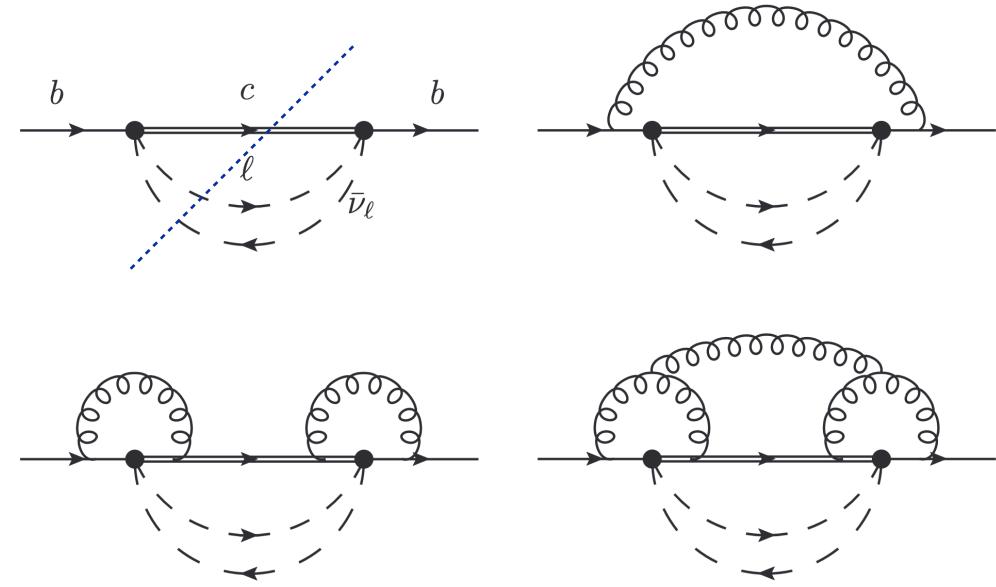
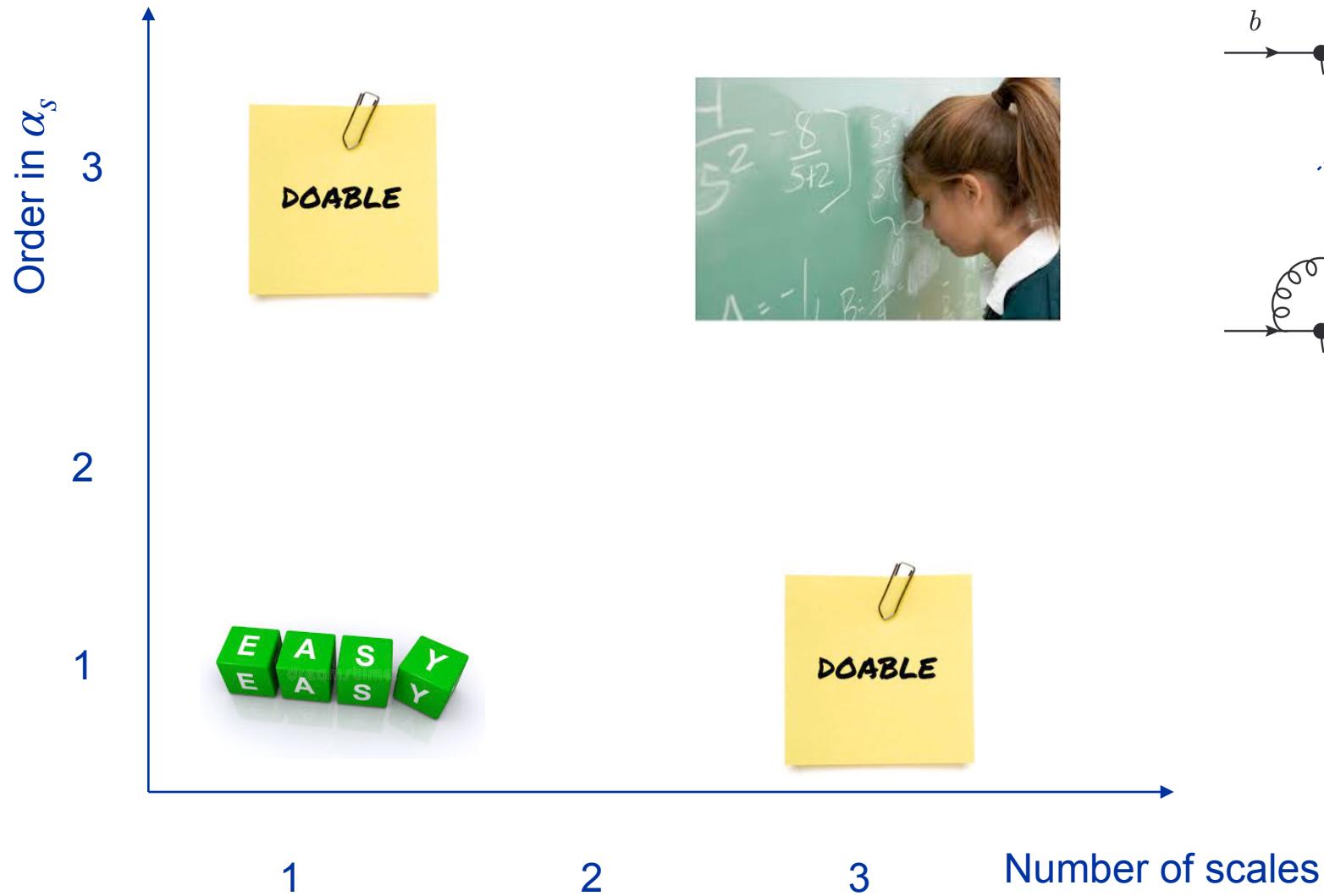
- $\alpha_s(\mu_s), m_b^{\text{kin}}(\mu_{WC}), \bar{m}_c(\mu_c), \mu_G^2(\mu_g)$
- Assess theoretical uncertainties

$$m_b/2 \leq \mu_s \leq 2m_b$$

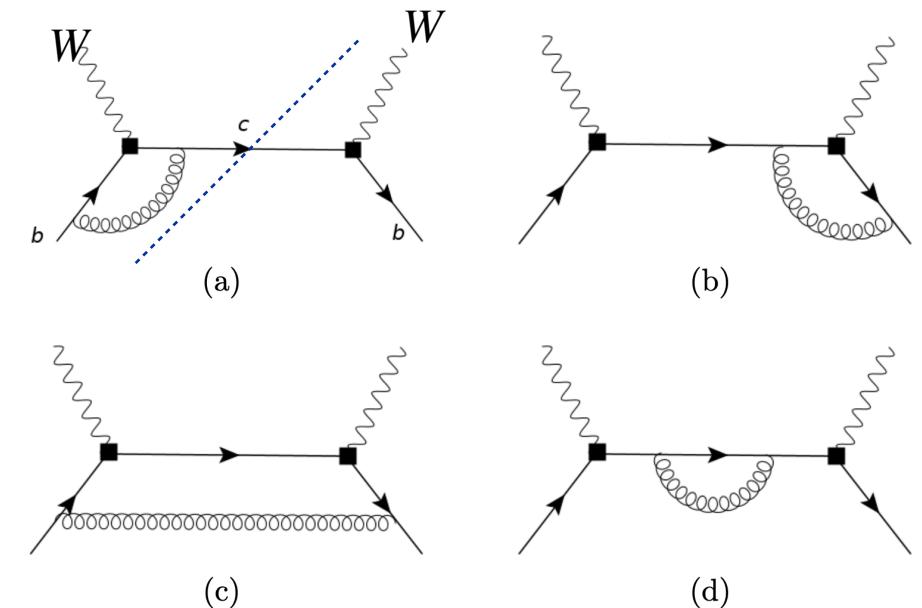
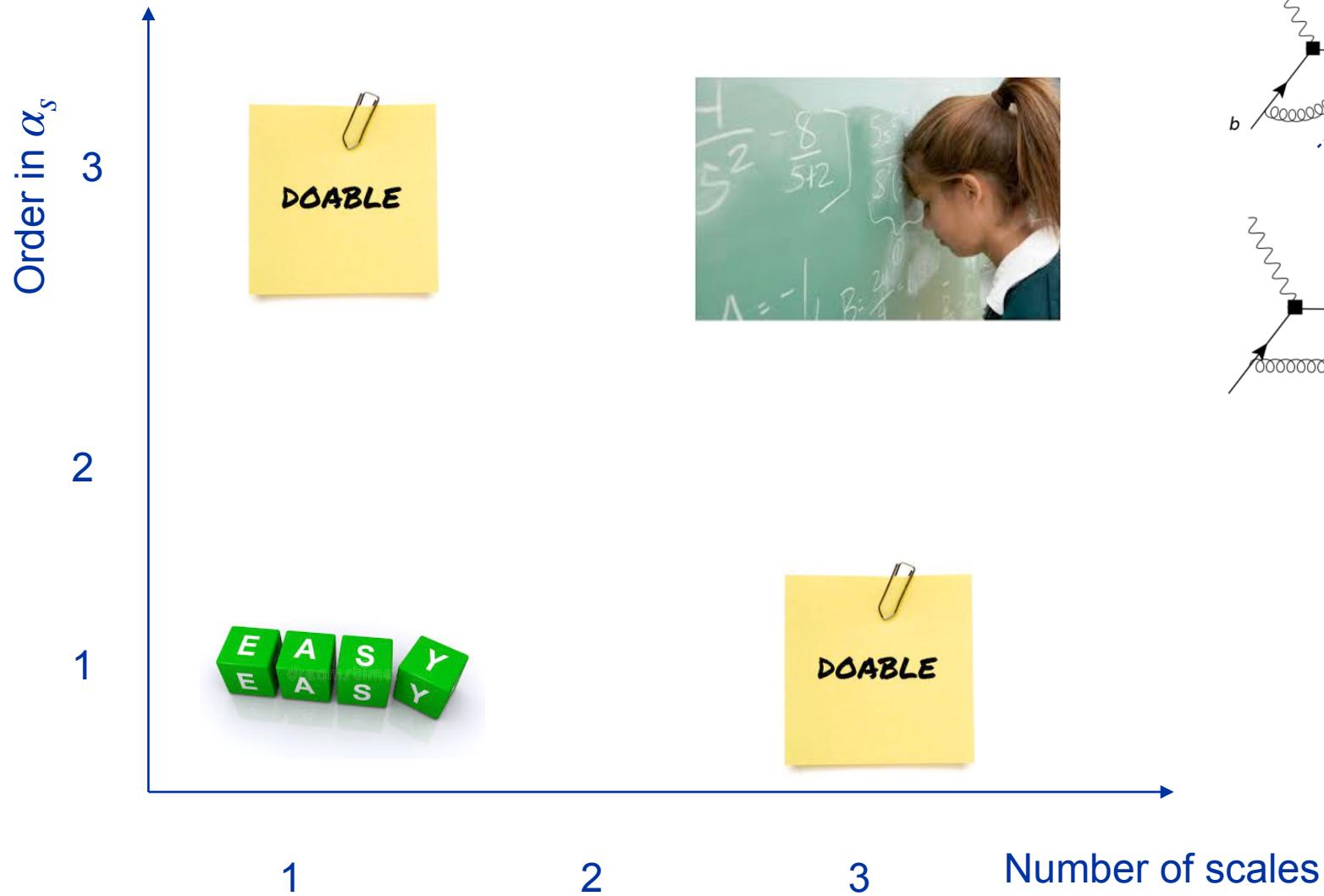
- Reduce impact of theory correlations



MF, Schönwald, Steinhauser, *Phys.Rev.D* 104 (2021) 016003, *JHEP* 08 (2022) 039.



- Scales:**
- $m_b \rightarrow 1$
- m_c



- **Scales:**
- $m_b \rightarrow 1$
- m_c
- q^2
- $q_0 = p_b \cdot q$

State-of-the art

- **Total Rate**

- NNLO Czarnecki, Pak, *Phys.Rev.D* 78 (2008) 114015, *Phys.Rev.Lett.* 100 (2008) 241807
- N3LO ($b \rightarrow c$) MF, Schönwald, Steinhauser, *Phys.Rev.D* 104 (2021) 016003.
- N3LO ($b \rightarrow u$) **NEW** MF, Usovitsch, hep-ph/2310.03685

- **M_x and E_l moments**

- NLO differential rate Aquila, Gambino, Ridolfi, Uraltsev, *Nucl.Phys.B* 719 (2005) 77
- NNLO for moments with $E_{\text{cut}} < E_l$, numerical results for specific E_{cut} and $\rho = m_c/m_b$ Biswas, Melnikov, *JHEP* 02 (2010) 089; Gambino, *JHEP* 09 (2011) 055. Gambino, *JHEP* 09 (2011) 055.

- NLO for μ_π^2 and μ_G^2 Alberti, Gambino, Nandi, *Nucl.Phys.B* 870 (2013) 16, *JHEP* 01 (2014) 147
- N3LO for moments without cuts MF, Schönwald, Steinhauser, *JHEP* 08 (2022) 039.

- **q^2 moments with a lower cut on q^2**

- NLO up to ρ_D^3 Moreno, Mannel, Pivovarov, *Phys.Rev.D* 105 (2022) 5, 054033
- NNLO for moments with $q_{\text{cut}}^2 \leq q^2$ **NEW** MF, Herren, in preparation

- **QED effects** Bordone, Gambino, Haisch, Piccione, hep-ph-2309.02849



NNLO corrections to q^2 spectrum

MF, Herren, in preparation

GOAL:

$$\frac{d\Gamma}{d\hat{q}^2} = \frac{G_F^2 m_b^5}{192\pi^3} |V_{cb}|^2 \left[F_0(\rho, \hat{q}^2) + \frac{\alpha_s}{\pi} F_1(\rho, \hat{q}^2) + \left(\frac{\alpha_s}{\pi} \right)^2 F_2(\rho, \hat{q}^2) \right] + O\left(\frac{1}{m_b^2}\right)$$

$$d\Gamma \simeq \delta(p_L^2 - q^2) W_{\mu\nu} L^{\mu\nu} d\Phi_2(p_b; p_L, p_X) d\Phi_2(p_L; p_\ell, p_\nu) dq^2$$

with $p_L = p_l + p_\nu$



Method employed at NLO in

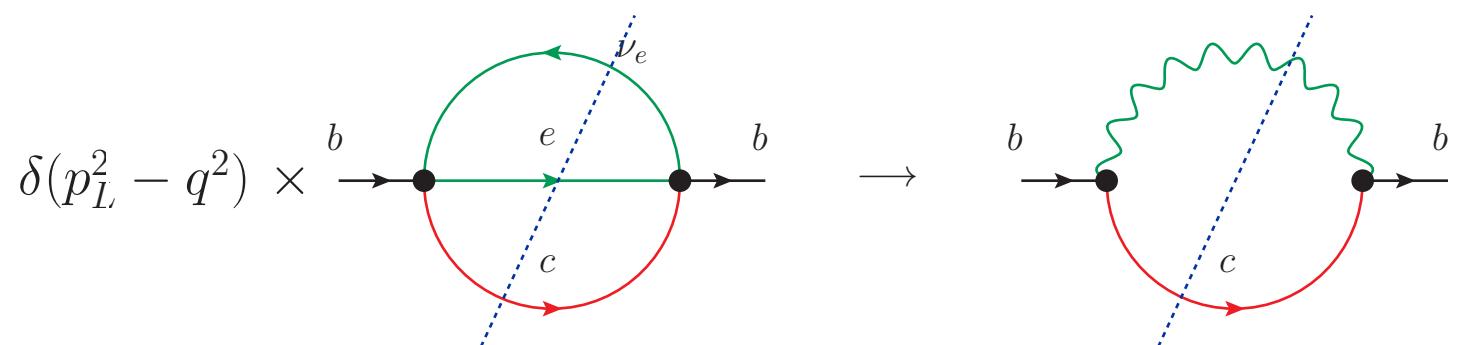
Moreno, Mannel, Pivovarov, *Phys.Rev.D* 105 (2022) 5, 054033

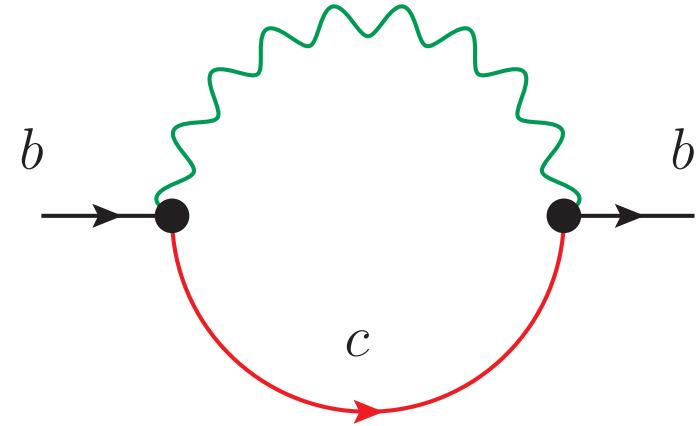
Integration of electron and neutrino phase-space (with $p_L = p_l + p_\nu$)

$$\mathcal{L}^{\mu\nu}(p_L) = \int L^{\mu\nu} d\Phi_2(p_L; p_l, p_\nu) = \frac{1}{384\pi^5} \left(1 - \frac{m_\ell^2}{p_L^2}\right)^2 \left[\left(1 + \frac{2m_\ell^2}{p_L^2}\right) p_L^\mu p_L^\nu - g^{\mu\nu} p_L^2 \left(1 + \frac{m_\ell^2}{2p_L^2}\right) \right]$$

Inverse unitarity $\delta(p_L^2 - q^2) \rightarrow \frac{1}{2\pi i} \left[\frac{1}{p_L^2 - q^2 - i0} - \frac{1}{p_L^2 - q^2 + i0} \right]$

Anastasiou, Melnikov, *Nucl.Phys.B* 646 (2002) 220



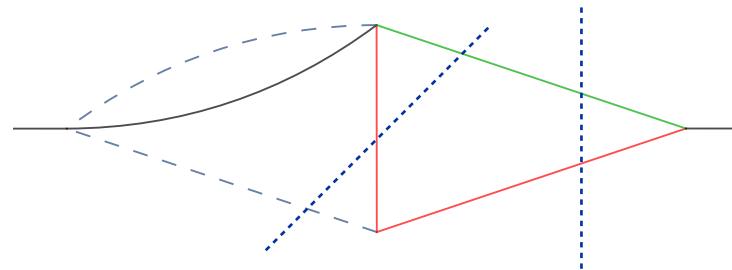


$$\rightarrow \sum_n \underbrace{\frac{p_n(\hat{q}^2, \rho; \epsilon)}{q_n(\hat{q}^2, \rho; \epsilon)}}_{\text{polynomials}} \underbrace{J_n(\hat{q}^2, \rho; \epsilon)}_{\text{master integrals}}$$

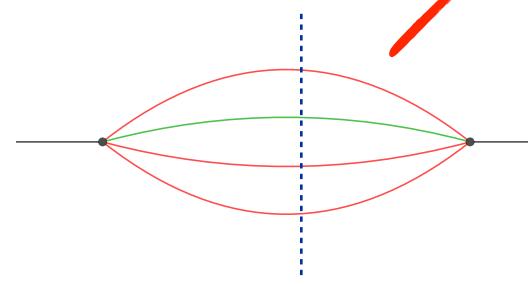
$$\epsilon = \frac{4-d}{2}$$

- **98 master integrals**
- **Neglect cuts through 3 charm**

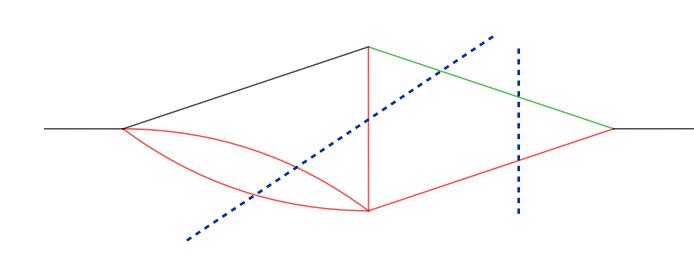
see also: Egner, MF, Schönwald, Steinhauser, HEP 09 (2023) 112



One charm cuts



Three charm cuts



One charm cuts

- **Differential equations**

- Canonical form

Henn, Phys.Rev.Lett. 110 (2013) 251601

Lee, Comput.Phys.Commun. 267 (2021) 108058

$$\frac{\partial \vec{J}}{\partial \rho} = \epsilon \hat{M}_\rho(\hat{q}^2, \rho) \vec{J}(\hat{q}^2, \rho, \epsilon)$$

$$\frac{\partial \vec{J}}{\partial \hat{q}^2} = \epsilon \hat{M}_{q^2}(\hat{q}^2, \rho) \vec{J}(\hat{q}^2, \rho, \epsilon)$$

- Boundary conditions evaluated with AMFlow and subsequent PSLq

Xiao Liu, Yan-Qing Ma, Comput.Phys.Commun. 283 (2023) 108565

- Analytic solution in terms of Generalised Polylogarithms (GPLs)

$$G(a_1, \dots, a_n; z) = \int_0^z \frac{dt}{t - a_1} G(a_2, \dots, a_n; t)$$



$$G(a_1, \dots, a_n; z) = \int_0^z \frac{dt}{t - a_1} G(a_2, \dots, a_n; t)$$

Examples:

$$G(0; z) = \log(z)$$

$$G(x, z) = \log\left(1 - \frac{z}{x}\right)$$

$$\underbrace{G(0, \dots, 0; z)}_n = \frac{\log^n(z)}{n!}$$

$$\underbrace{G(0, \dots, 0, x, z)}_n = -\text{Li}_n\left(\frac{z}{x}\right)$$

Fast numerical evaluation with GiNaC and PolyLogTools

<http://www.ginac.de>

Duhr, Dulat, JHEP 08 (2019) 135

$$G\left(x, \frac{1+x^2}{x}, x, \frac{1}{x}; z\right) \Bigg|_{x=1/2, z=1/3} = 0.00151860208899279\dots$$



$$\frac{d\Gamma}{d\hat{q}^2} = \frac{G_F^2 m_b^5}{192\pi^3} |V_{cb}|^2 \left[F_0(\rho, \hat{q}^2) + \frac{\alpha_s}{\pi} F_1(\rho, \hat{q}^2) + \left(\frac{\alpha_s}{\pi}\right)^2 F_2(\rho, \hat{q}^2) \right] + O\left(\frac{1}{m_b^2}\right)$$

Normalised moments

$$\langle (q^2)^n \rangle_{q_{\text{cut}}^2} = \int_{q^2 > q_{\text{cut}}^2} (q^2)^n \frac{d\Gamma}{dq^2} dq^2 \Bigg/ \int_{q^2 > q_{\text{cut}}^2} \frac{d\Gamma}{dq^2} dq^2$$

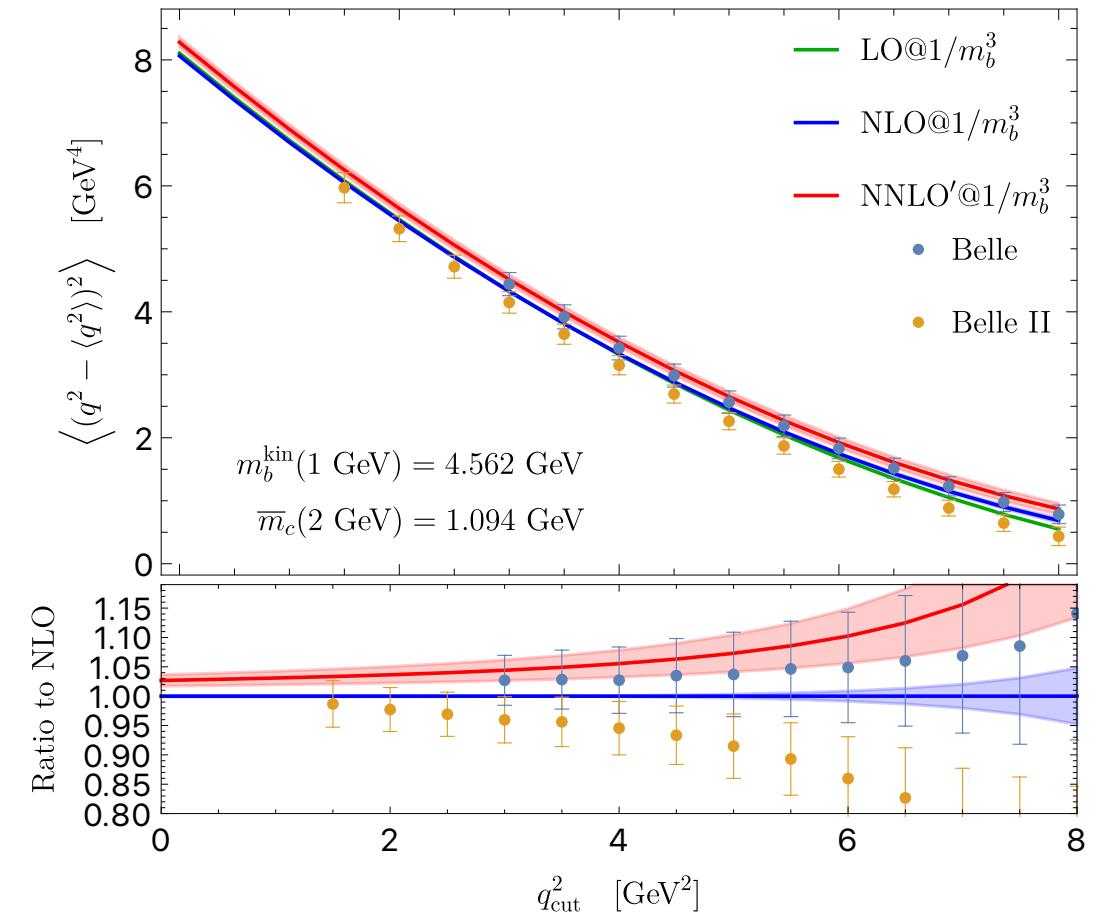
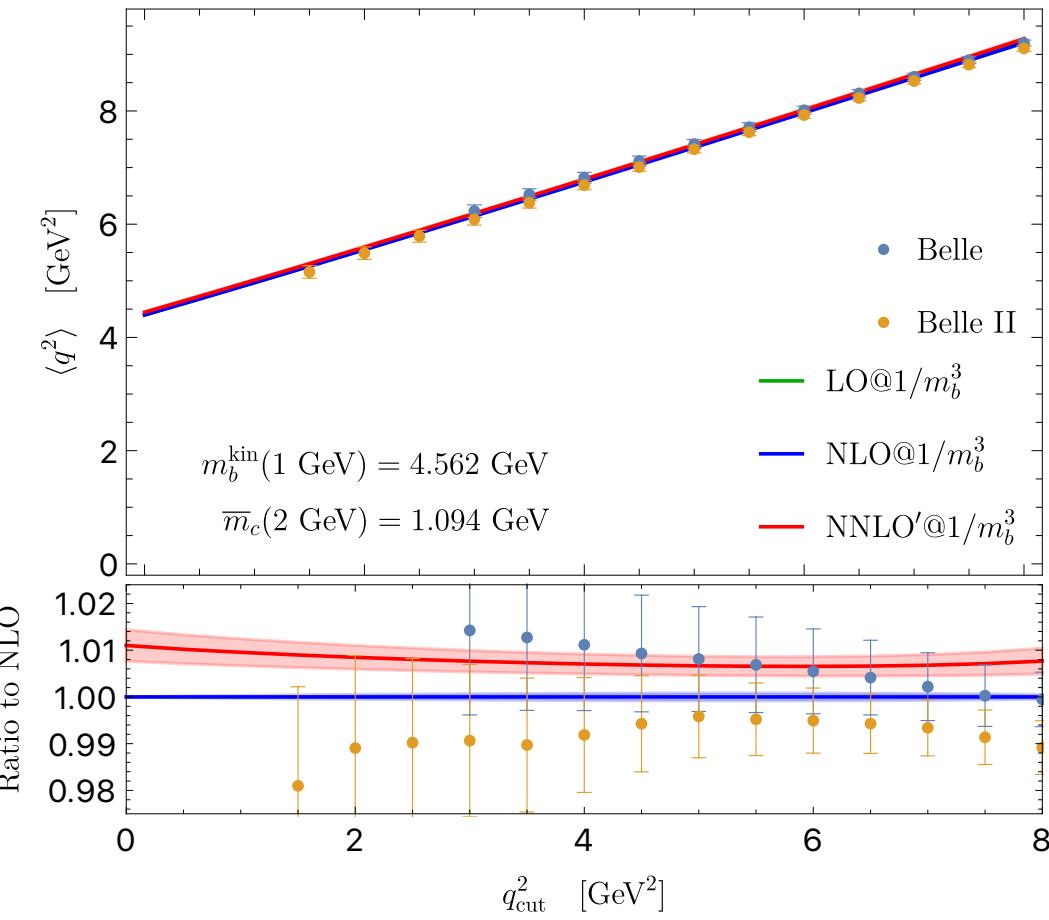
Centralized moments

$$\left\langle (q^2 - \langle q^2 \rangle)^n \right\rangle_{q_{\text{cut}}^2} = \sum_{i=1}^n \binom{n}{i} \langle q^{2i} \rangle \left(\langle q^2 \rangle \right)^{n-i}$$

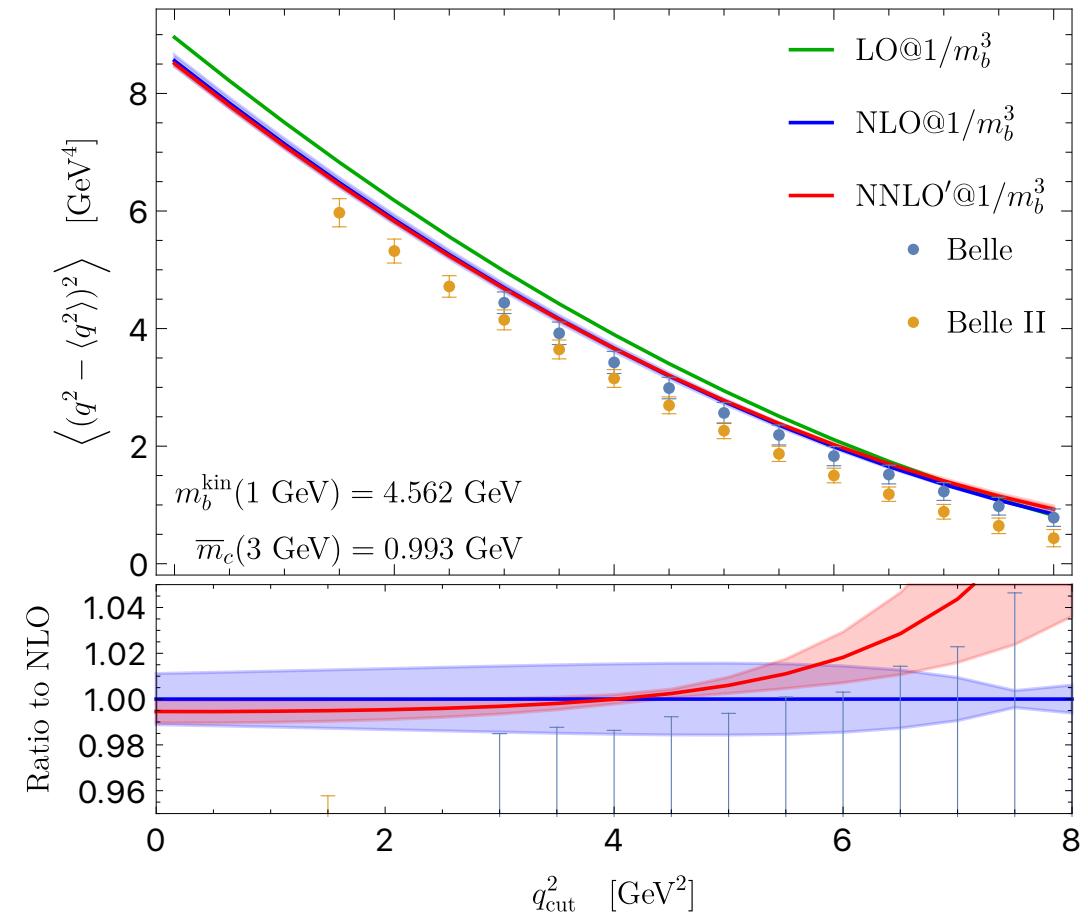
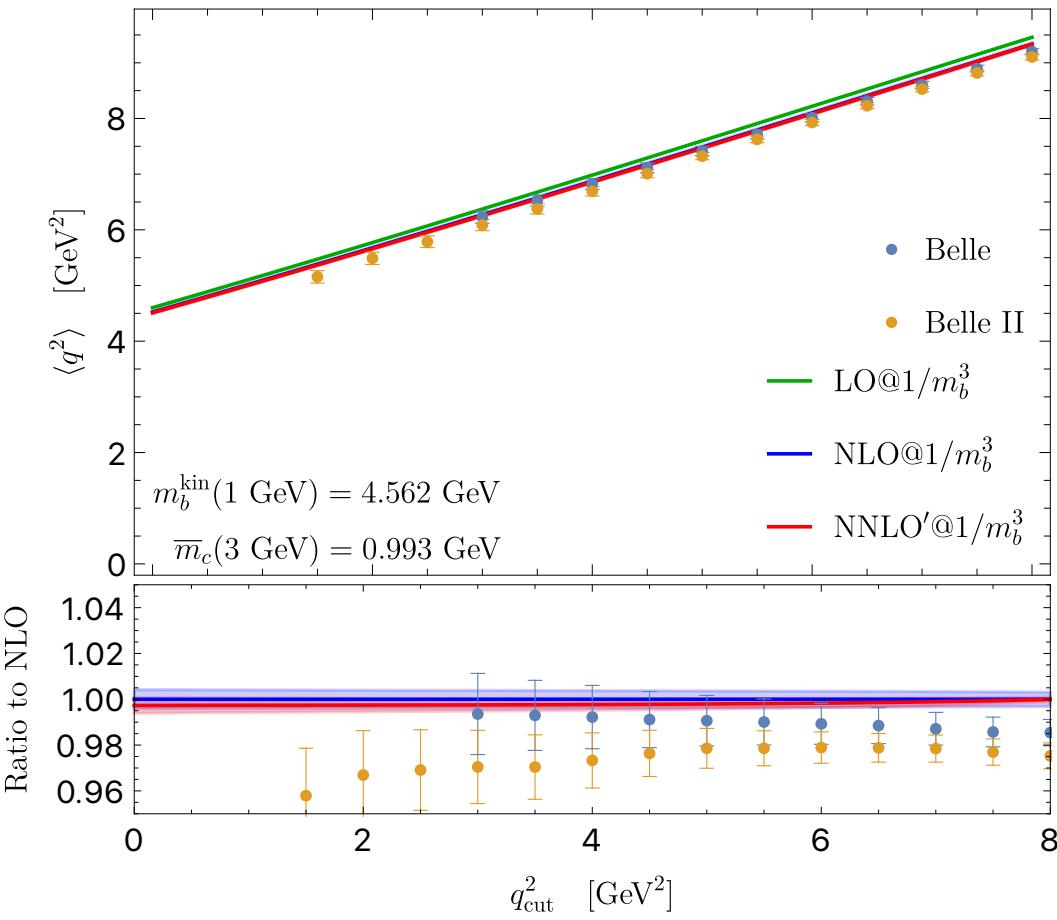
Change of mass scheme

$$m_b^{\text{OS}}, m_c^{\text{OS}} \rightarrow m_b^{\text{kin}}, \bar{m}_c$$





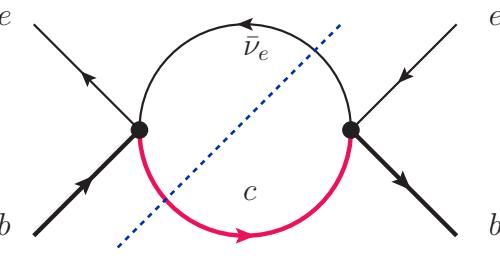
HQE parameters from: Bernlochner et al, JHEP 10 (2022) 068



HQE parameters from: Bernlochner et al, JHEP 10 (2022) 068

Revisiting NNLO corrections to E_l moments

- Similar strategy applies to energy spectrum of the electron:


$$\simeq \frac{1}{E_l} \frac{d\Gamma}{dE_l} \quad \rightarrow \quad \langle E_l^n \rangle_{E_{\text{cut}}} = \int_{E_l > E_{\text{cut}}} (E_l)^n \frac{d\Gamma}{dE_l} dE_l$$

- The master integrals depends on two scales: $\rho = m_c/m_b$ and $E_l = p_b \cdot p_l/m_b$.
- At NLO there are 9 master integrals.
- Perfect numerical agreement with integration of differential rate.

Aquila, Gambino, Ridolfi, Uraltsev, Nucl.Phys.B 719 (2005) 77

- Possibility to extend the calculation at NNLO under study.

MF, Herren, Schönwald, work in progress



Why only E_l moments?

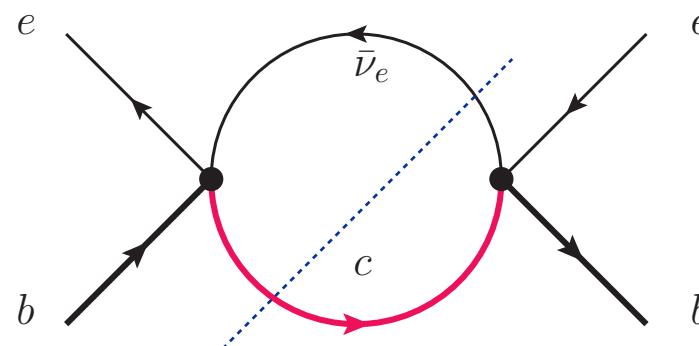
$$\langle (M_X^2)^n \rangle_{E_{\text{cut}}} = \int_{E_l > E_{\text{cut}}} M_X^{2n} \frac{d\Gamma}{dq^2 dq_0 dE_l} dq^2 dq_0 dE_l \quad \left/ \int_{E_l > E_{\text{cut}}} \frac{d\Gamma}{dE_l} dE_l \right.$$

$$\langle (q^2)^n \rangle_{E_{\text{cut}}} = \int_{E_l > E_{\text{cut}}} q^{2n} \frac{d\Gamma}{dq^2 dq_0 dE_l} dq^2 dq_0 dE_l \quad \left/ \int_{E_l > E_{\text{cut}}} \frac{d\Gamma}{dE_l} dE_l \right.$$

Integration order does not matter

$$dM_X^{2n} = (M_X^2)^n \times W_{\mu\nu} L^{\mu\nu} d\Phi_2(p_b; p_l, p_\nu, p_X)$$

$$= (M_B^2 + q^2 - 2M_B q_0)^n \times$$



Why only E_l moments?

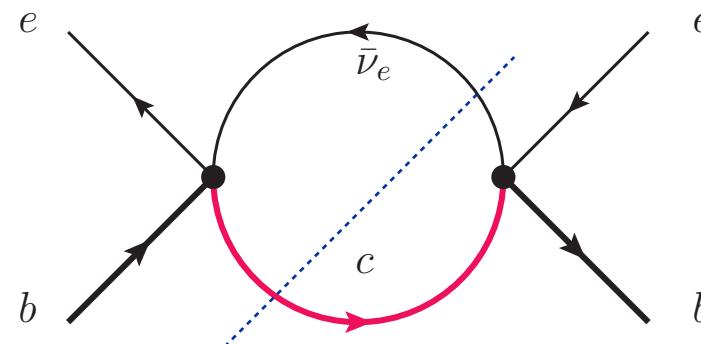
$$\langle (M_X^2)^n \rangle_{E_{\text{cut}}} = \int_{E_l > E_{\text{cut}}} \frac{dM_X^{2n}}{dE_l} dq^2 dq_0 dE_l \left/ \int_{E_l > E_{\text{cut}}} \frac{d\Gamma}{dE_l} dE_l \right.$$

$$\langle (q^2)^n \rangle_{E_{\text{cut}}} = \int_{E_l > E_{\text{cut}}} \frac{dq^{2n}}{dE_l} dq^2 dq_0 dE_l \left/ \int_{E_l > E_{\text{cut}}} \frac{d\Gamma}{dE_l} dE_l \right.$$

Integration order does not matter

$$dM_X^{2n} = (M_X^2)^n \times W_{\mu\nu} L^{\mu\nu} d\Phi_2(p_b; p_l, p_\nu, p_X)$$

$$= (M_B^2 + q^2 - 2M_B q_0)^n \times$$

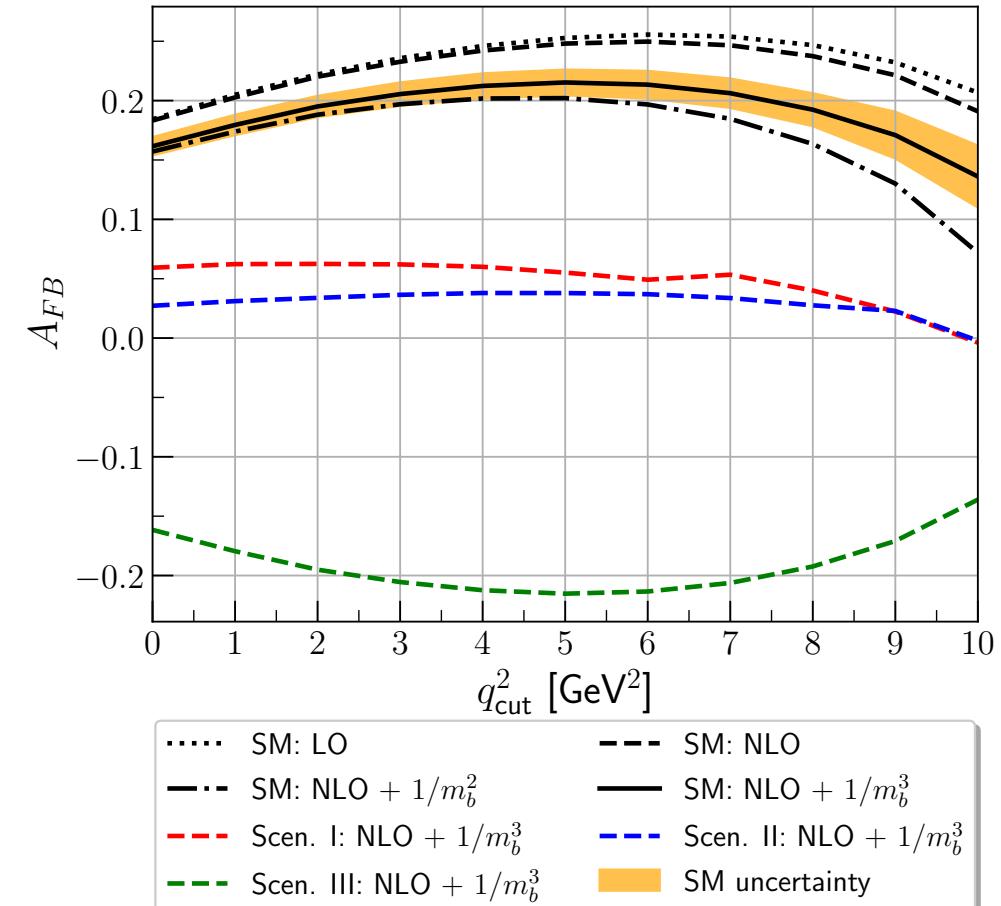
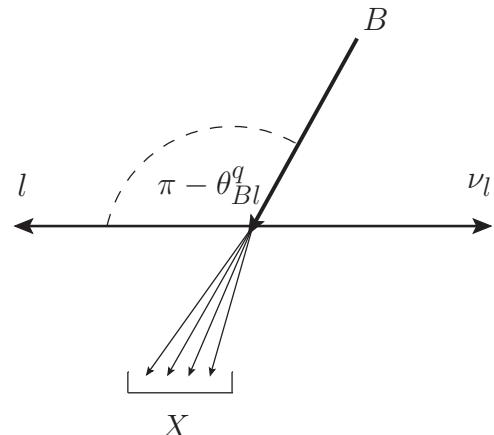


Forward-backward asymmetry

$$A_{FB} = \frac{\int_{-1}^0 \frac{d\Gamma}{dz} - \int_{-1}^0 \frac{d\Gamma}{dz}}{\int_{-1}^0 \frac{d\Gamma}{dz} + \int_{-1}^0 \frac{d\Gamma}{dz}}$$

$$z = \cos \theta = \frac{\nu \cdot p_\nu - \nu \cdot p_\ell}{\sqrt{(\nu \cdot q)^2 - q^2}}$$

$A_{FB}(q_{\text{cut}}^2 = 4 \text{ GeV}) = 24.6(1 - 0.019|_{\alpha_s} - 0.12|_{\text{pow}}) \times 10^{-2}$



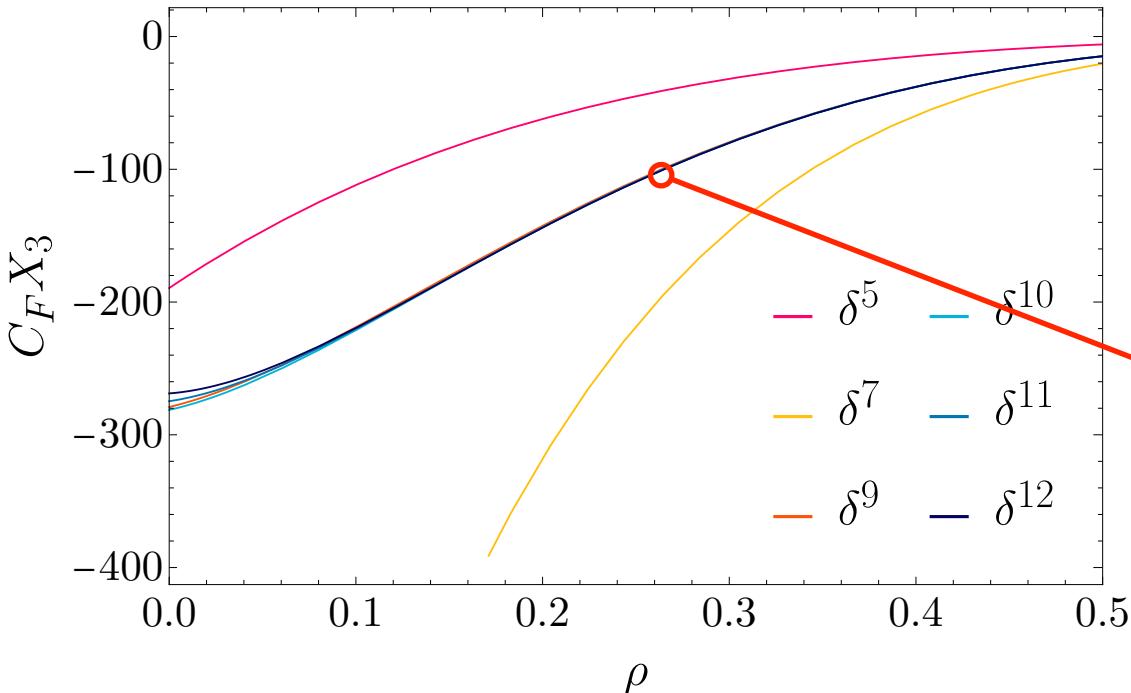
MF, Rahimi, Vos, JHEP 02 (2023) 086

State-of-the art

- **Total Rate**
 - NNLO Czarnecki, Pak, *Phys.Rev.D* 78 (2008) 114015, *Phys.Rev.Lett.* 100 (2008) 241807
 - N3LO ($b \rightarrow c$) MF, Schönwald, Steinhauser, *Phys.Rev.D* 104 (2021) 016003.
 - N3LO ($b \rightarrow u$) **NEW** MF, Usovitsch, hep-ph/2310.03685
- **M_x and E_l moments**
 - NLO differential rate Aquila, Gambino, Ridolfi, Uraltsev, *Nucl.Phys.B* 719 (2005) 77
 - NNLO for moments with $E_{\text{cut}} < E_l$, numerical results for specific E_{cut} and $\rho = m_c/m_b$ Biswas, Melnikov, *JHEP* 02 (2010) 089; Gambino, *JHEP* 09 (2011) 055. Gambino, *JHEP* 09 (2011) 055.
 - NLO for μ_π^2 and μ_G^2 Alberti, Gambino, Nandi, *Nucl.Phys.B* 870 (2013) 16, *JHEP* 01 (2014) 147
 - N3LO for moments without cuts MF, Schönwald, Steinhauser, *JHEP* 08 (2022) 039.
- **q^2 moments with a lower cut on q^2**
 - NLO up to ρ_D^3 Moreno, Mannel, Pivovarov, *Phys.Rev.D* 105 (2022) 5, 054033
 - NNLO for moments with $q_{\text{cut}}^2 \leq q^2$ **NEW** MF, Herren, in preparation



Total rate of $b \rightarrow u$



$$\delta = 1 - \frac{m_c}{m_b}$$

$$\Gamma_{\text{sl}} = \frac{G_F^2 m_b^5 A_{\text{ew}}}{192\pi^3} |V_{cb}|^2 \left(X_0(\rho) + C_F \sum_n \left(\frac{\alpha_s}{\pi} \right)^n X_n(\rho) \right)$$

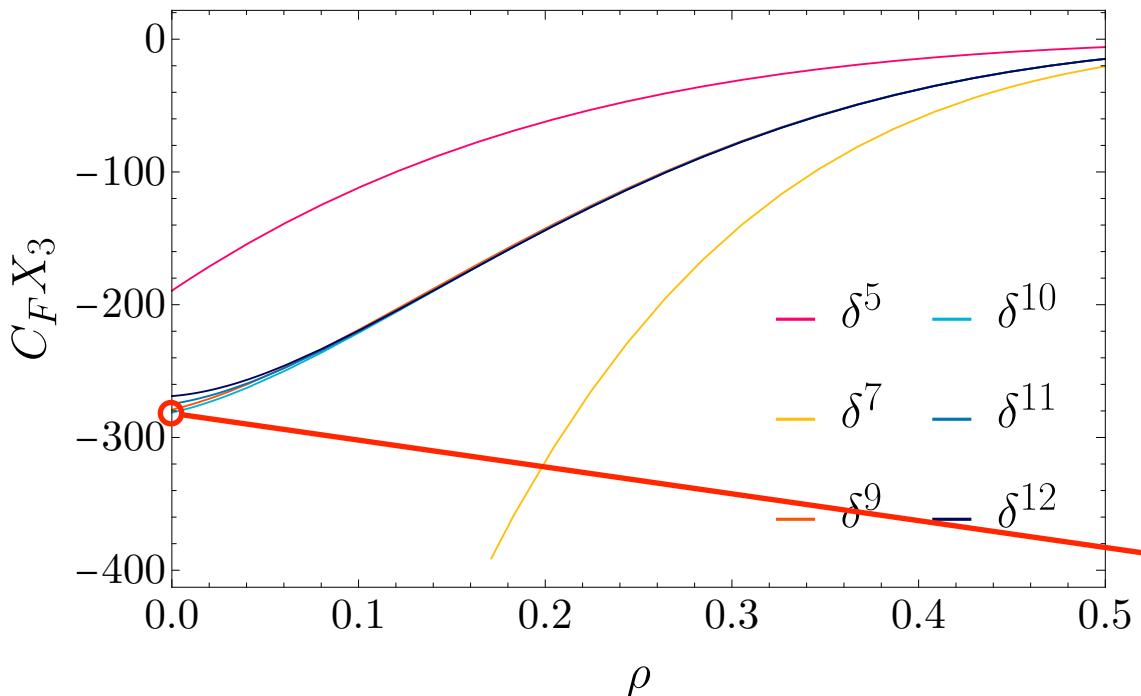
$$C_F X_3(\rho = 0.28) = -91.2 \pm 0.4 (0.4\%)$$

MF, Schönwald, Steinhauser, [Phys.Rev.D 104 \(2021\) 016003](#), [JHEP 08 \(2022\) 039](#).

$$m_b^{\text{OS}} : m_c^{\text{OS}} \quad 1 - 1.78 \left(\frac{\alpha_s}{\pi} \right) - 13.1 \left(\frac{\alpha_s}{\pi} \right)^2 - 163.3 \left(\frac{\alpha_s}{\pi} \right)^3$$

$$m_b^{\text{kin}}(1 \text{ GeV}) : \bar{m}_c(2 \text{ GeV}) \quad 1 - 1.24 \left(\frac{\alpha_s}{\pi} \right) - 3.65 \left(\frac{\alpha_s}{\pi} \right)^2 - 1.0 \left(\frac{\alpha_s}{\pi} \right)^3$$

$$m_b^{1S} : m_c \text{ via HQET} \quad 1 - 1.38 \left(\frac{\alpha_s}{\pi} \right) - 6.32 \left(\frac{\alpha_s}{\pi} \right)^2 - 33.1 \left(\frac{\alpha_s}{\pi} \right)^3$$



MF, Schönwald, Steinhauser, *Phys.Rev.D* 104 (2021) 016003, *JHEP* 08 (2022) 039.

$$\Gamma_{\text{sl}} = \frac{G_F^2 m_b^5 A_{\text{ew}}}{192\pi^3} |V_{cb}|^2 \left(X_0(0) + C_F \sum_n \left(\frac{\alpha_s}{\pi} \right)^n X_n(0) \right)$$

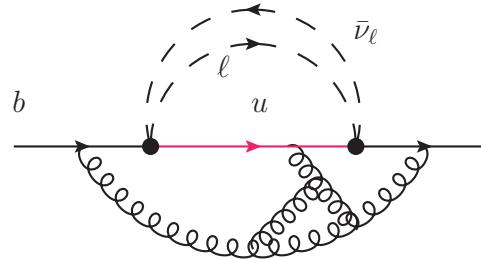
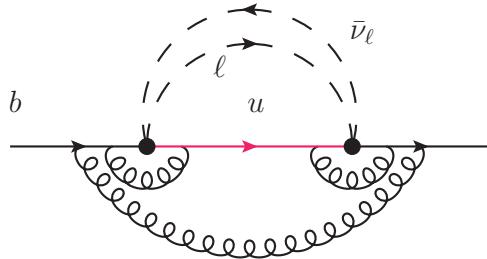
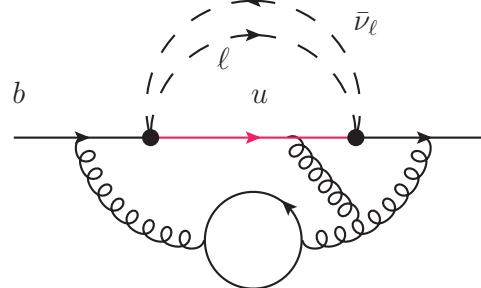
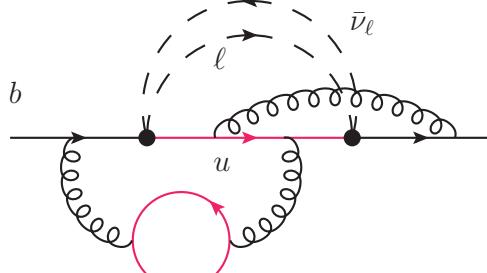
$$C_F X_3(\rho = 0) = -269 \pm 27 (10\%)$$

$$m_b^{\text{kin}}(1 \text{ GeV}) : \bar{m}_c(3 \text{ GeV}) \quad \Gamma_{\text{sl}} \simeq 1 - 0.19|_{\alpha_s} + 0.019|_{\alpha_s^2} + 0.032(9)|_{\alpha_s^3}$$

- Underestimated uncertainty on X_3 ?
- Kinetic mass not good for $b \rightarrow u$?
- Weak-annihilation?
- Large m_c effects at $O(\alpha_s^3)$?

Third order correction to $b \rightarrow u$

MF, Usovitsch, hep-ph-2310.03685



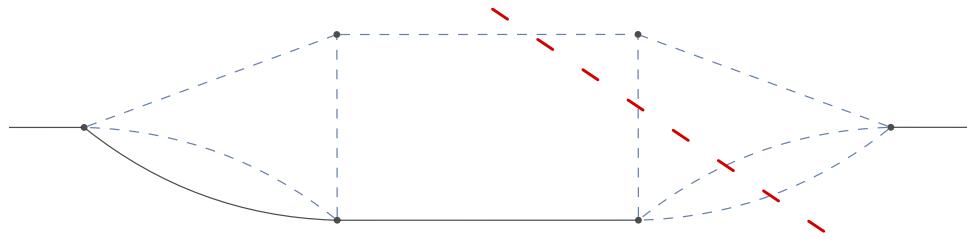
Fermionic corrections

$$X_3 = N_L^2 T_F^2 X_{N_L^2} + N_H^2 T_F^2 X_{N_H^2} + N_H N_L T_F^2 X_{N_H N_L}$$
$$+ N_L T_F (C_F X_{N_L C_F} + C_A X_{N_L C_A})$$
$$+ N_H T_F (C_F X_{N_H C_F} + C_A X_{N_H C_A})$$

$$+ C_F^2 X_{C_F^2} + C_F C_A X_{C_F C_A} + C_A^2 X_{C_A^2}$$

Bosonic corrections

Numerical evaluation of 5 loop integrals



- Challenging integration-by-parts reduction
- 1369 master integrals
- Numerical evaluation with 40 digits with AMFlow
Xiao Liu, Yan-Qing Ma, Comput.Phys.Commun. 283 (2023) 108565
- Parallel calculation large- N_c contributions

Chen, Li, Li, Wang, Wand, Wu, hep-ph/2309.00762

	This work	Ref. [28]	Difference
$T_F^2 N_L^2$	-6.9195	-6.34 (42)	8.3%
$T_F^2 N_H^2$	-1.8768×10^{-2}	$-1.97 (42) \times 10^{-2}$	5.0%
$T_F^2 N_H N_L$	-1.2881×10^{-2}	$-1.1 (1.1) \times 10^{-2}$	12%
$C_F T_F N_L$	-7.1876	-5.65 (55)	22%
$C_A T_F N_L$	42.717	39.7 (2.1)	7%
$C_F T_F N_H$	2.1098	2.056 (64)	2.5%
$C_A T_F N_H$	-0.45059	-0.449 (18)	0.4%

$$\begin{aligned} C_F X_3 &= 280.2 && \text{fermionic} \\ &-536.4 && \text{bosonic, large } N_c \\ &-11.6 (2.7) && \text{bosonic, subleading } N_c \\ &= -267.8 (2.7) && \text{MF, Usovitsch, hep-ph-2310.03685} \end{aligned}$$

$$C_F X_3(\rho = 0) = -269 \pm 27 (10\%)$$

MF, Schönwald, Steinhauser, Phys.Rev.D 104 (2021) 016003, JHEP 08 (2022) 039.



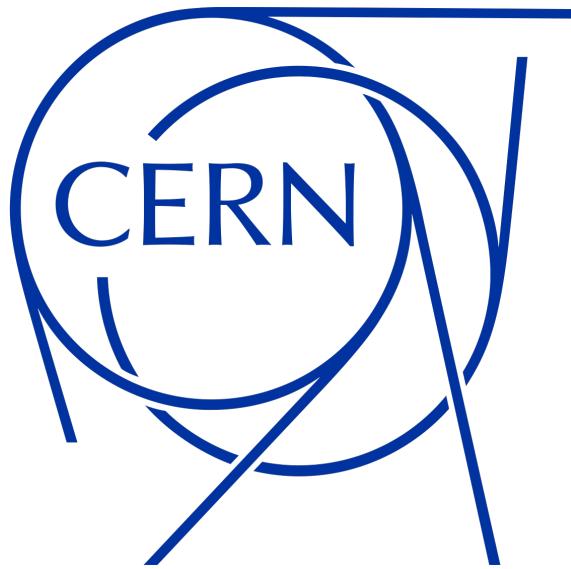
Conclusions

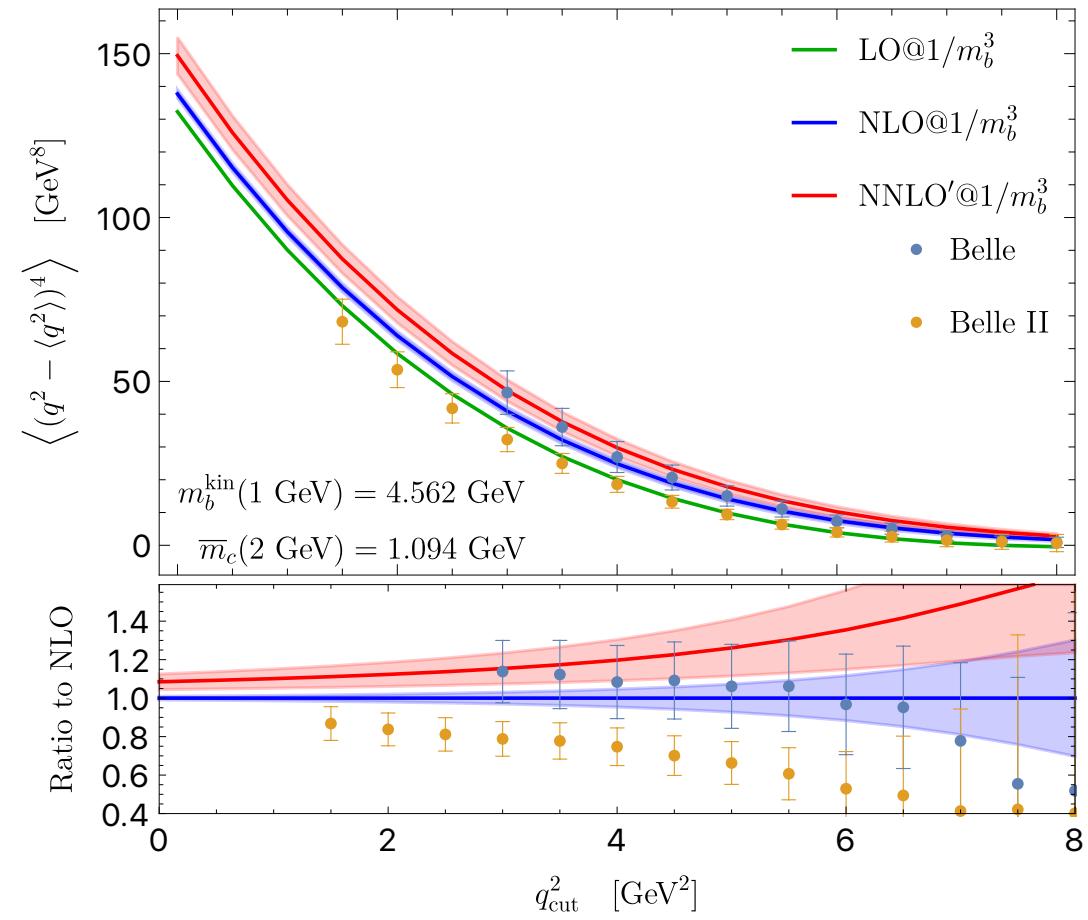
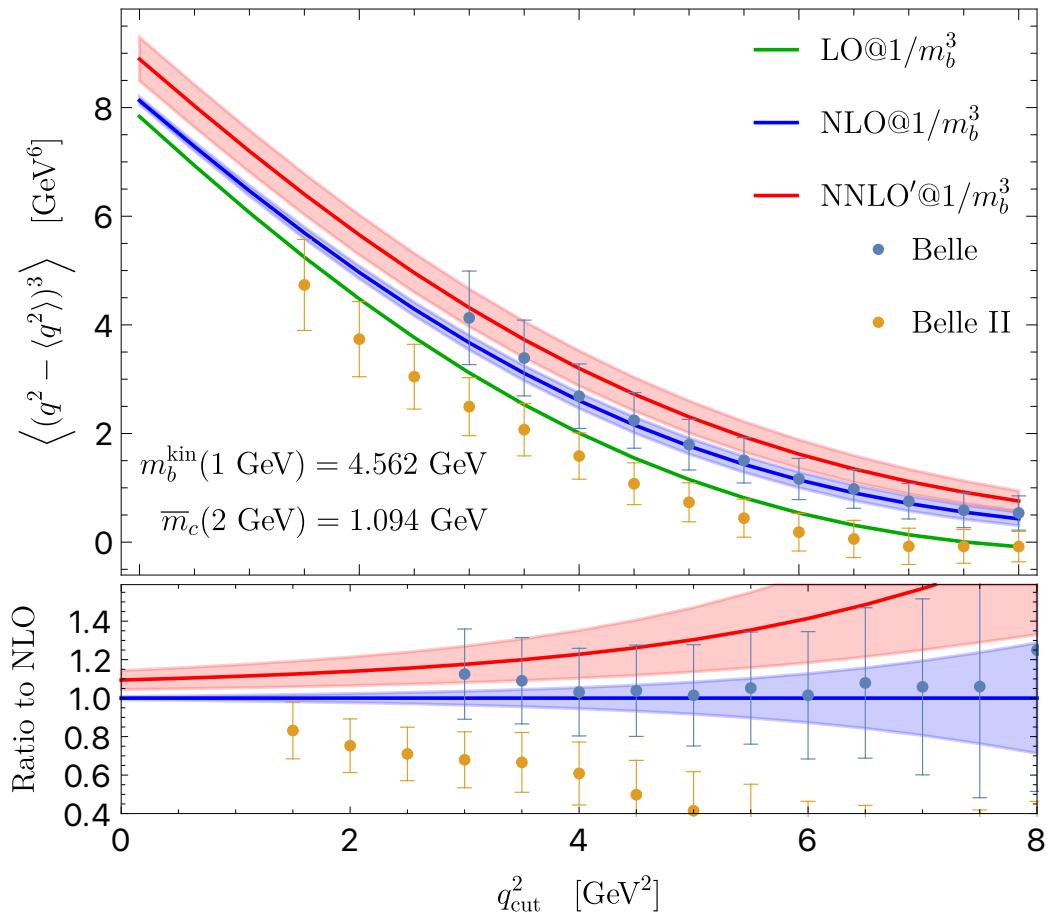
- In the last years the theory of inclusive decays has greatly profited from developments in computational methods for multi-loop integrals.
- Achievements in the last 2-3 years: kinetic mass, total rate, uncut moments at $\mathcal{O}(\alpha_s^3)$
- Differential q^2 spectrum at NNLO analytic:
 - $\langle q^{2n} \rangle_{q_{\text{cut}}^2}$
 - but also $\text{Br}(q_{\text{cut}}^2)$, $\langle M_X^{2n} \rangle_{q_{\text{cut}}^2}$ and $\langle E_l^n \rangle_{q_{\text{cut}}^2}$
- Differential E_l spectrum likely also doable at NNLO analytically
- NLO correction to power suppressed terms (ρ_D and $1/m_b^{4,5}$) may be obtained with the same strategy

Conclusions

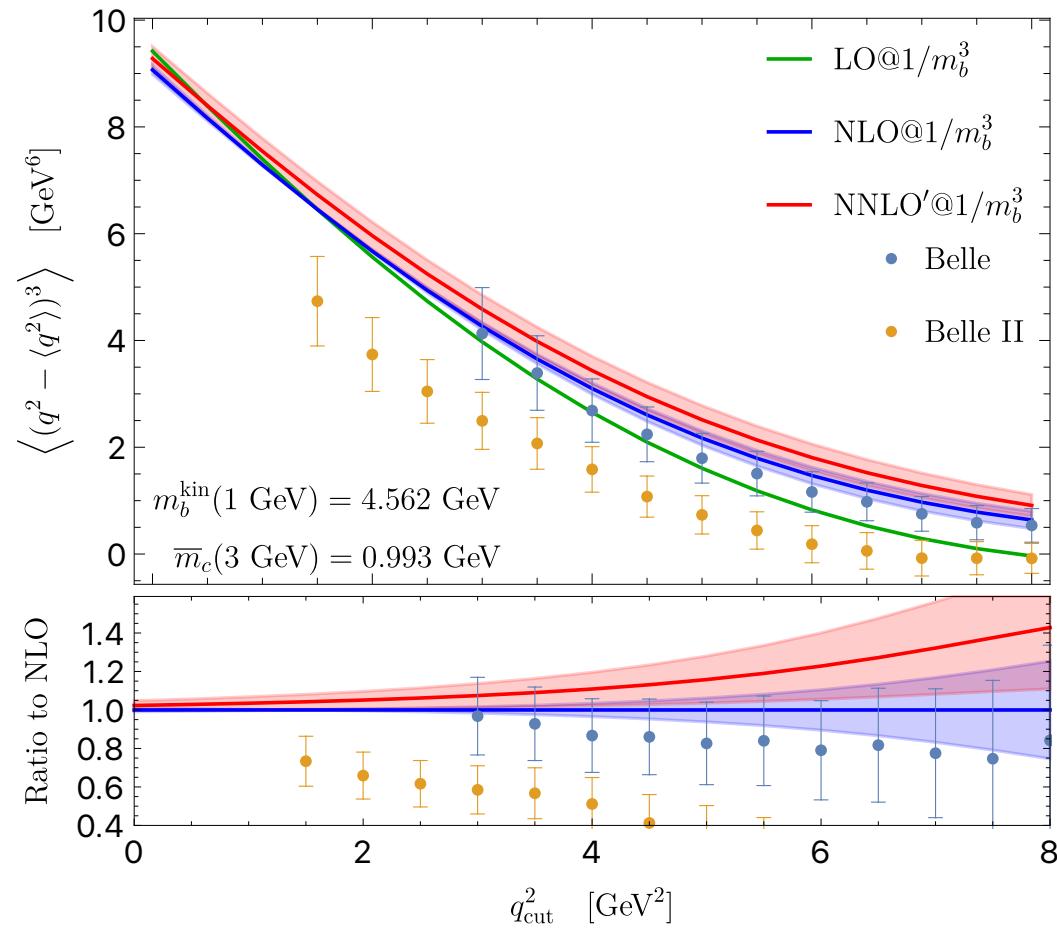
- Kinetic mass at $O(\alpha_s^4)$: reduce $\overline{\text{MS}} \rightarrow \text{kin}$ scheme conversion uncertainty on m_b from 15 MeV to ~ 8 MeV
- Power correction of order μ_{WC}^3/m_b^3 in the kinetic mass definition
- We are working hard to be ready for the new upcoming measurements at Belle III!



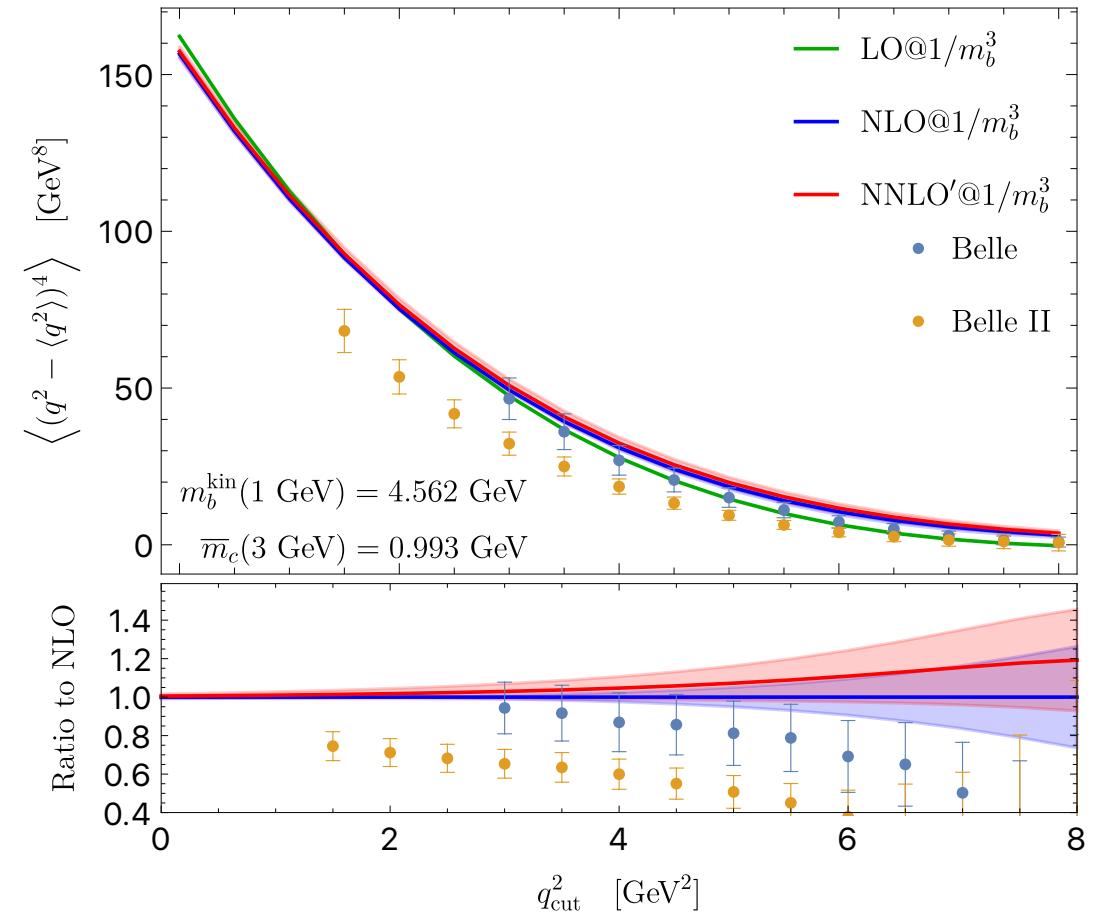




HQE parameters from: Bernlochner et al, JHEP 10 (2022) 068



HQE parameters from: Bernlochner et al, JHEP 10 (2022) 068



Kinetic scheme

- Calculation easily performed in the pole mass scheme for m_b and m_c
- Pole scheme leads to badly behaved perturbation theory
- The fits of $|V_{cb}|_{\text{inc}}$ use the **kinetic scheme**

$$m_b^{\text{pole}} = m_b^{\text{kin}}(\mu) + [\bar{\Lambda}(\mu)]_{\text{pert}} + \frac{[\mu_\pi^2(\mu)]_{\text{pert}}}{m_b^{\text{kin}}(\mu)} + \dots$$

- Also the HQE parameters are re-defined in the kinetic scheme

$$\mu_\pi^2(\mu) = \mu_\pi^2(0) - [\mu_\pi^2(\mu)]_{\text{pert}} \quad \rho_D^3(\mu) = \rho_D^3(0) - [\rho_D^3(\mu)]_{\text{pert}}$$

$$[\mu_\pi^2(\mu)]_{\text{pert}} = \frac{\alpha_s}{\pi} C_F \mu^2 + \dots$$

$$[\rho_D^3(\mu)]_{\text{pert}} = \frac{2}{3} \frac{\alpha_s}{\pi} C_F \mu^3 + \dots$$

$$m_b^{\text{kin}}(1 \text{ GeV}) = 4.573 \text{ GeV} \quad \bar{m}_c(2 \text{ GeV}) = 1.092 \text{ GeV} \quad q_{\text{cut}}^2 = 2 \text{ GeV}^2$$

$$\begin{aligned}\langle q^2 \rangle &= 5.853470329 + 0.2365684899 \text{ api CF} - 5.870252829 \text{ api}^2 \text{ CF} + 12.31040061 \text{ api}^2 \text{ CF flagb0} \\ \langle (q^2 - \langle q^2 \rangle)^2 \rangle &= 6.599610658 + 0.5418960093 \text{ api CF} - 20.80577919 \text{ api}^2 \text{ CF} + 48.85002521 \text{ api}^2 \text{ CF flagb0} \\ \langle (q^2 - \langle q^2 \rangle)^3 \rangle &= 7.003374353 + 3.90484573 \text{ api CF} - 45.2523971 \text{ api}^2 \text{ CF} + 149.8472321 \text{ api}^2 \text{ CF flagb0} - 17\end{aligned}$$

$$m_b^{\text{kin}}(1 \text{ GeV}) = 4.573 \text{ GeV} \quad \bar{m}_c(3 \text{ GeV}) = 0.993 \text{ GeV} \quad q_{\text{cut}}^2 = 2 \text{ GeV}^2$$

$$\begin{aligned}\langle q^2 \rangle &= 6.072411413 - 0.7043840353 \text{ api CF} - 13.53637271 \text{ api}^2 \text{ CF} + 10.36154604 \text{ api}^2 \text{ CF flagb0} \\ \langle (q^2 - \langle q^2 \rangle)^2 \rangle &= 7.448558317 - 3.423938153 \text{ api CF} - 51.49139191 \text{ api}^2 \text{ CF} + 43.98509335 \text{ api}^2 \text{ CF flagb0} - \\ \langle (q^2 - \langle q^2 \rangle)^3 \rangle &= 8.740268785 - 4.734517097 \text{ api CF} - 116.2284133 \text{ api}^2 \text{ CF} + 157.5509862 \text{ api}^2 \text{ CF flagb0} -\end{aligned}$$